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| 4 | Changes in optical characteristics of surface microlayers hint to |
| 5 | photochemically and microbially-mediated DOM turnover in the upwelling |
| 6 | region off the coast off Peru |
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23 The coastal upwelling system off Peru is characterized by high biological activity and a pronounced 24 subsurface oxygen minimum zone, as well as associated emissions of atmospheric trace gases such as N₂O, CH₄ and CO₂. From December 3rd to 23rd 2012, METEOR (M91) cruise took place in the 25 Peruvian upwelling system between 4.59° S and 15.4° S, and 82.0° W to 77.5° W. During M91 we 26 27 investigated the composition of the sea-surface microlayer (SML), the oceanic uppermost boundary 28 directly subject to high solar radiation, often enriched in specific organic compounds of biological 29 origin like Chromophoric Dissolved Organic Matter (CDOM) and marine gels. In the SML, the 30 continuous photochemical and microbial recycling of organic matter may strongly influence gas 31 exchange between marine systems and the atmosphere. We analyzed SML and underlying water (ULW) samples at 38 stations focusing on CDOM spectral characteristics as indicator of 32 33 photochemical and microbial alteration processes. CDOM composition was characterized by 34 spectral slope (S) values and Excitation-Emission Matrix fluorescence (EEMs), which allow to 35 track changes in molecular weight (MW) of DOM, and to determine potential DOM sources and sinks. Spectral slope S varied between 0.012 to 0.043 nm⁻¹ and was quite similar between SML and 36 37 ULW, with no significant differences between the two compartments. Higher S values were 38 observed in the ULW of the southern stations below 15°S. By EEMs, we identified five fluorescent 39 components (F1-5) of the CDOM pool, of which two had excitation/emission characteristics of 40 amino-acid like fluorophores (F1, F4) and were highly enriched in the SML, with a median ratio 41 SML:ULW of 1.5 for both fluorophores. In the study region, values for CDOM absorption ranged from 0.07 to 1.47 m⁻¹. CDOM was generally highly concentrated in the SML, with a median 42 enrichment with respect to the ULW of 1.2. CDOM composition and changes in spectral slope 43 44 properties suggested a local microbial release of DOM directly in the SML as a response to light 45 exposure in this extreme environment. In a conceptual model of the sources and modifications of 46 optically active DOM in the SML and underlying seawater (ULW), we describe processes we think

47 may take place (see graphical abstract): The production of CDOM of higher MW by microbial release through growth, exudation and lysis in the euphotic zone, includes the identified 48 49 fluorophores (F1, F2, F3, F4, F5). Specific amino-acid like fluorophores (F1, F4) accumulate in the 50 SML with respect to the ULW, as photochemistry may enhance microbial CDOM release by a) 51 photoprotection mechanisms and b) cell-lysis processes. Microbial and photochemical degradation 52 are potential sinks of the amino-acid like fluorophores (F1, F4), and potential sources of reworked 53 and more refractory humic-like components (F2, F3, F5). In the highly productive upwelling region 54 along the Peruvian coast, the interplay of microbial and photochemical processes controls the 55 enrichment of amino-acid like CDOM in the SML. We discuss potential implications for air-sea gas 56 exchange in this area.



- 57 58
- 1. Introduction
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61 The Peruvian Eastern Boundary Upwelling System (EBUS), extending along the coast off Peru 62 between 4° and about 40° South, is among the most productive marine ecosystems worldwide 63 (Capone and Hutchins, 2013;Chavez and Messié, 2009;Rosenberg et al., 1983) and it is

64 characterized by high biological activity, involving high export rates of organic carbon both 65 vertically and laterally (Arístegui et al., 2004; Muller-Karger et al., 2005). The high productivity is 66 sustained by winds year-round that promote the upwelling of nutrient-rich deep waters into the 67 euphotic zone, thus favoring phytoplankton photosynthesis and organic matter production (Chavez 68 and Messié, 2009). High rates of organic matter production are counterbalanced by heterotrophic 69 respiration, which provides sinks for the oxygen produced by autotrophs and leads to subsurface 70 Oxygen Minimum Zones (OMZs) (Lachkar and Gruber, 2011). OMZs are expanding worldwide 71 due to reduced solubility at increasing temperatures, as well as a consequence of reduced oceanic 72 ventilation and enhanced stratification (Keeling et al., 2010;Stramma et al., 2008). OMZ become 73 increasingly important as key marine regions for the emission of climate-relevant gases like carbon 74 dioxide (CO₂), methane (CH₄), nitrous oxide (N₂O) and hydrogen sulfide (H₂S) (Paulmier et al., 75 2008;Paulmier et al., 2011). N₂O is a strong greenhouse gas and ozone-reactive: 30% of its 76 atmospheric concentration has an oceanic source (Solomon et al., 2007), of which, up to 75% is 77 supported by OMZs (Bange et al., 2001). Therefore, OMZs are key environments to assess the 78 oceanic contribution to the concentration of atmospheric gases. Defining the processes that regulate 79 gas fluxes across the water-air interface is a central objective to better understand the reciprocal 80 relationship between changes in our climate and marine environments.

81 The uppermost oceanic layer in contact with the atmosphere is the sea-surface microlayer (SML), 82 which mediates major climate-relevant processes including air-sea gas exchange and sea-spray 83 aerosol emission (Liss and Duce, 2005). This interface between a liquid (hydrosphere) and a gas 84 phase (atmosphere) accumulates organic matter of biological origin, creating a sort of "skin" of 85 surface-active compounds able to damp capillary waves and "capping the flux" of gases across the 86 water-air interface (GESAMP, 1995). Natural organic compounds in the SML include a vast array 87 of photosynthesis products including carbohydrates, amino acids and lipids, as well as other carbon-88 rich compounds like dissolved organic matter (DOM) and marine gels (e.g. Cunliffe et al., 2013). 89 The DOM pool represents a continuum of molecular weights and biological lability ranging from

90 refractory to labile DOM being utilized rapidly by microorganisms (Benner, 2002;Carlson, 2002), 91 or photochemically degraded (Kieber, 2000). These compounds, produced in the oceanic photic 92 zone and brought to the SML through rising bubbles (Hardy, 1982), contribute to the enrichment of 93 a natural surface biofilm and favor specific SML heterotrophic communities that are very active in 94 recycling this organic material (Hardy, 1982;Cunliffe et al., 2011). While bulk dissolved organic 95 carbon is not generally enriched in the SML, specific DOM fractions are present occasionally at 96 much higher concentrations than in the underlying water (Cunliffe et al., 2013). These enriched 97 pools of organic matter include marine gel particles (Wurl and Holmes, 2008), chromophoric 98 dissolved organic matter (CDOM) (Zhang and Yang, 2013; Tilstone et al., 2010) and phenolic 99 material (Carlson, 1982;Carlson and Mayer, 1980).

100 CDOM is the principal light-absorbing constituent of DOM, strongly absorbing UV (100 - 400 nm) 101 and visible radiation (400 - 700 nm), and it can comprise 20%-70% of the DOM in oceanic waters 102 (Coble, 2007). CDOM plays a major role in the attenuation of UV wavelengths and can reduce the 103 availability of underwater photosynthetically active radiation for primary production (Bracchini et 104 al., 2011). Photolysis of CDOM promotes the formation of low molecular weight (LMW) 105 compounds from the breakdown of high molecular weight DOM (HMW-DOM), facilitating the 106 bioavailability of carbon uptake for microbial growth from biologically refractory material, and 107 representing an important loss pathway for CDOM in the oceans (Kieber et al., 1989). Other major 108 by-products of CDOM photolysis are carbon monoxide (CO), which often exists at supersaturated 109 concentrations in the oceans' surface (Blough, 2005, and references therein), CO₂ (Miller and Zepp, 110 1995) and reactive chemical species (Loiselle et al., 2012). To initiate a photochemical reaction, 111 light must first be absorbed and in this respect the SML is very well exposed to elevated solar 112 radiation (Liss and Duce, 2005). CDOM photolysis may affect biological processes within the SML 113 as well as the structure of accumulated organic matter. Optical properties and photochemical 114 cycling of DOM have been widely investigated in the ocean: CDOM alters light spectra in the 115 surface ocean and its spatial and temporal distribution have been used in characterizing water 116 masses exchange (Nelson and Siegel, 2013). However, processes within the SML remain poorly 117 understood. Possible effects of photochemistry on SML chemical composition have been discussed 118 in the past (Blough, 2005), but still little is known on CDOM fluorophores, sources and sinks 119 (Tilstone et al., 2010;Zhang and Yang, 2013). To discern sources, sinks and modification of DOM 120 in surface waters, whether microbially or photochemically-induced, we investigated optical 121 properties of organic sea-surface microlayers and underlying water samples in the highly productive 122 EBUS off Peru. We applied optical spectroscopy measurements combined with chemical and 123 biological analysis to identify different compounds within the CDOM pool and their partitioning 124 between the SML and the underlying water. The use of excitation-emission matrix fluorescence 125 spectroscopy (EEMs) allowed us to discriminate different compound classes in the SML and 126 underlying water based on their excitation and emission maxima (Coble, 1996).

127 At present, the oceans are subject to many changes in physical and chemical properties like pH, 128 temperature, and dissolved oxygen concentration, which potentially will affect the biological 129 cycling of carbon (Riebesell et al., 2009;Keeling et al., 2010;Bopp et al., 2002). Whether the oceans 130 are sources or sinks of carbon depends on the production rate of organic matter with respect to its 131 biological degradation (Del Giorgio and Duarte, 2002), and high DOM degradation in the SML 132 might represent a net source of CO_2 to the atmosphere (Garabétian, 1990). It is well known that the 133 composition of the SML reflects biological processes of the euphotic zone (Galgani et al., 2014;Gao 134 et al., 2012; Matrai et al., 2008; Bigg et al., 2004), and that elevated concentrations of organic matter 135 may accumulate in the SML in highly productive regions like the Peruvian EBUS (Engel and 136 Galgani, 2016). The enrichment of light-absorbing DOM in the SML may increase the 137 photochemical formation and fluxes of reactive chemical species at the surface, with potentially 138 important consequences for the composition of the SML itself and for the fate of compounds 139 passing through this interface (Blough, 2005). Last but not least, the photochemical DOM 140 breakdown may increase the biological availability of carbon, thus increasing heterotrophic 141 respiration and CO_2 flux to the atmosphere.

142 CDOM contributes to the dissolved organic carbon (DOC) pool, but while DOC is a bulk measure, 143 CDOM is a characteristic of DOM rather than a discrete class of compounds (Nelson and Siegel, 144 2013). Positive correlations have been observed between CDOM and DOC in coastal systems and 145 plankton enclosures (Loginova et al., 2015), but the strenght of these correlations varies much 146 across regional and seasonal differences (Blough and Del Vecchio, 2002). CDOM is a precursor for 147 photochemical reactions that may drive the emission of trace gases from photochemically-altered 148 DOM (e.g. Ciuraru et al., 2015). Therefore, in upwelling areas associated with OMZs, CDOM 149 characteristics in the SML are worth to be investigated as they may impact the exchange of gases 150 between the ocean and the atmosphere.

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2. Material and methods

153 **2.1. Study area**

154 The R/V METEOR cruise M91 was an integrated biogeochemical study in the upwelling region off 155 Peru, with the aim to assess the importance of oxygen minimum zones (OMZs) for the air-sea 156 exchange of gases relevant for climate and tropospheric chemistry (Bange 2013). A total of 39 157 samples for SML and underlying water were collected in December 2012 between 5°S and 16°S off 158 the Peruvian coast. Data that we report here additionally from what previously described by Engel 159 and Galgani (2016) refer to 38 stations. For easiness of comparison, table 1 recalls salinity, water 160 temperature, radiation and wind speed, as already described in the companion manuscript (Engel 161 and Galgani, 2016).

Some stations were revisited for multiple sampling (Table 2): S7 and S7_2; S12_1, S12_2, and S12_3; S16_1, S16_2, S16_3; S20 and S20_2. These stations were sampled within a time frame of evolution of CDOM optical properties in the SML and ULW, as we were interested in monitoring the evolution of CDOM optical Whenever possible, we sampled at sunrise, midday and sunset. For security reasons, it was not possible to sample later than sunset, as the zodiac operations were not allowed out at dark. Exact 168 latitude and longitude were not always possible to retrieve after a certain time, but were similar for169 the stations sampled in a few hours time lag.

170 The sampling approach for the SML was chosen as a silicate glass plate of 500 mm (length) x 250 mm (width) x 5 mm (thickness) with an effective sampling area of 2000 cm^2 as indicated in Engel 171 172 and Galgani (2016). Briefly, the glass plate was inserted into the water perpendicular to the surface and withdrawn at a controlled rate of ~ 20 cm s⁻¹ as first suggested by Harvey and Burzell (1972). 173 Different devices can be applied to sample the SML. The glass plate approach we choose collects a 174 175 thinner SML (~60 - 150 µm) when compared to i.e. the Garrett Screen (150 - 300 µm), one of the 176 mainly recognized practices introduced by Garrett in 1965 (Cunliffe et al., 2011; Garrett, 1965). 177 The glass plate was chosen because it allows the sampling of enough volume required for analysis 178 while keeping a minimal dilution with underlying water. Sampling was performed on a rubber 179 boat; in order to obtain a well-standardized procedure and to minimize biases by sampling, the same 180 person always took the samples with a repeatable withdrawal speed of the SML. The rubber boat 181 was positioned as far upwind of the ship as possible and away from the path taken by the ship in 182 order to avoid any potential surface contamination. The outboard motor of the rubber boat was 183 switched off and samples were collected in upwind clean waters.

184 Before collecting the sample into the bottle, we let the plate drain for 20 s approximately. Then, the 185 sample retained on both sides of the plate was removed with a Teflon wiper, and the procedure 186 repeated about 20 times to obtain the necessary volume for analysis. The exact amount of dips per 187 sample has been tracked. The first sample was discarded and used to rinse the collecting bottle (HCl 188 10% cleaned and Milli-Q rinsed). Glass plate and wiper were acid cleaned (HCl 10%) and Milli-Q 189 rinsed prior use, and at sampling site they were copiously rinsed with in situ seawater to minimize 190 any contamination with alien material during transport and handling. Underlying seawater (ULW) was collected right after SML at about ~ 20 cm depth by opening an acid cleaned (HCl 10%) and 191 192 Milli-Q rinsed glass bottle and closing it underwater. The thickness (d, m) of our reference SML 193 that we were able to collect was estimated as follows:

194 (1) d = V/(A x n)

195 Where V is the SML volume collected, i.e. 60-140 mL, A is the sampling area of the glass plate (A =196 2000 cm²) and n is the number of dips. During this cruise, the apparent sampling thickness of the 197 SML ranged between 45 and 60 μ m, with an overall mean of 49 \pm 8.9 μ m (Engel and Galgani, 198 2016). Many factors may influence the thickness of the SML such as withdrawal rate, dipping time, and plate dimensions. With a withdrawal speed of ~ 20 cm s⁻¹, the apparent SML thickness was in 199 200 accordance with previous findings at similar withdrawal rate reporting $60 - 100 \mu m$ (Harvey and 201 Burzell, 1972) and 50 - 60 µm (Zhang et al., 1998). The sampling thickness was very well 202 comparable among all stations, indicating that no major biases due to sampling procedure may have 203 occurred.

After sampling, bottles were stored in the dark and the samples immediately processes in the laboratory onboard, within maximum 30 minutes from sampling.

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207 **2.2.** Chemical and biological analyses

208 Dissolved organic matter (DOM): Sampling, calibration and analysis procedure for dissolved 209 organic carbon (DOC) and for dissolved hydrolysable amino acids (DHAA), have been described in 210 details in Engel and Galgani (2016). Additionally, to track DOM diagenetic state and 211 bioavailability, we used the carbon-normalized yields of dissolved amino acids to DOC, expressed 212 as DHAA%-DOC (Amon and Fitznar, 2001;Benner, 2002;Kaiser and Benner, 2009;Davis and 213 Benner, 2007). Amino acids generally comprise a large fraction of bioavailable organic matter and 214 are preferentially consumed by microbial activity quite rapidly. In surface waters they may be easily 215 photodegraded too. Therefore, the amount of carbon included in amino acids is considered as a 216 good indicator of DOM diagenesis and a value of ~ 2% of DHAA%-DOC may indicate the 217 threshold between labile and semi-labile and refractory DOM (Davis and Benner, 2007).

218 Samples for chromophoric and fluorescent DOM (CDOM and FDOM) were filtered through 0.45

 μm PES syringe filters and collected into 40 mL pre-combusted (8 h, 500° C) amber glass vials.

220 Samples were stored in the dark at 4° C with no other treatment than pre-filtering. Since storage 221 procedures may affect the absorbance and fluorescence properties of DOM, absorbance and 222 fluorescence readings were performed directly on-board within a few hours from sampling or the 223 next day according to Schneider-Zapp and colleagues (2013). Prior to measurements, samples were 224 stored in the dark and acclimatized at room temperature. For CDOM, triplicate absorbance 225 measurements were made on a Shimadzu 1800 UV-Visible Spectrophotometer in the range 220 to 226 700 nm with 0.5 nm increments, in a 10 cm path-length quartz cuvette against Milli-Q water as a 227 reference. For FDOM, 3-D fluorescence spectroscopy was performed with a Varian Cary Eclipse 228 Fluorescence Spectrophotometer equipped with a xenon flash lamp and data assembled into 229 Excitation/Emission matrices (EEMs) which enable to individuate single DOM fluorophores 230 (Coble, 1996) and to perform parallel factor analysis PARAFAC (Stedmon and Bro, 2008). 231 Samples have been acclimatized and scanned at a fixed 20°C temperature (Cary Single Cell Peltier 232 Accessory, VARIAN) in 1 cm path length quartz cuvette. Scans were performed at 600 nm/min 233 using an excitation range (Ex) of 240-450 nm with 5 nm increments and recorded emission (Em) in 234 the range 242-600 nm with 2 nm increments. Samples were run in a mode of 5 nm slit for both 235 excitation and emission and 0.1 s integration time.

Particulate Organic Carbon (POC) and gel particles: Total numbers of gel particles were
determined by microscopy after Engel (2009). A detailed description of the method used during
M91 cruise can be found in Engel and Galgani (2016). POC data were retrieved after Engel and
Galgani (2016). We refer to this companion publication for further analytical details.

240 Phytoplankton and heterotrophic bacteria: Samples, calibration and analysis for phytoplankton
241 and heterotrophic bacteria counts for M91 are described in details in Engel and Galgani (2016).

242

243 2.3. Data analysis

244 **CDOM:** The measured absorbance at every wavelength λ was converted to absorption coefficient 245 $a(\lambda), (m^{-1})$, according to the following equation (Bricaud et al., 1981):

246 (2)
$$a(\lambda) = 2.303 A_{\lambda}/L$$

247 where A_{λ} is the absorbance and L is the path-length of the cuvette (here 0.10 m). Absorbance is an 248 optical characteristic of CDOM, which allows quantifying the amount of CDOM in the samples. 249 Therefore, the absorption coefficient $a(\lambda)$ is considered as a proxy for CDOM concentration. To 250 estimate CDOM concentration, we calculated the absorption coefficient at 325 nm as often used for 251 the open ocean (Swan et al., 2009;Nelson and Siegel, 2013). The dependence of a on the wavelength was determined by analyzing the spectral slope parameter S (nm⁻¹) in the discrete 252 253 wavelength ranges of 275-295 nm and 350-400 nm, determined by linear regression of log-254 transformed absorption spectra against the wavelength (Bricaud et al., 1981;Helms et al., 2008):

255 (3)
$$a(\lambda) = a(\lambda_0)e^{-S(\lambda - \lambda_0)}$$

where $a(\lambda_0)$ is the absorption coefficient at a reference wavelength λ_0 . S measured in the wavelength 256 range 275-295 nm (S(275-295), nm⁻¹) and 350-400 nm (S(350-400), nm⁻¹) as well as slope ratio 257 258 (SR) defined as S(275-295): S(350-400) are good indicators to characterize CDOM (Helms et al., 259 2008). SR is characterized by lower values for terrestrial CDOM compared to CDOM produced by 260 autochthonous marine sources and instead of S alone, could be a more sensitive indicator of 261 photochemically induced changes in the molecular weight of the CDOM pool as an increase in SR 262 suggests photodegradation processes, while a decrease is often related to microbially altered CDOM 263 (Helms et al., 2008). Both S(275-295) and SR increase with a) irradiation (photobleaching), b) with 264 decreasing DOM molecular weight, and c) at higher salinity reflecting mixing of water masses 265 along a salinity gradient. As such they are useful as tracers to determine mixing and coastal inputs. 266 We also determined the SUVA₂₅₄ index, that is, the specific ultraviolet absorbance (A) at 254 nm normalized to DOC concentration. This index was shown to correlate significantly with increasing 267 268 aromaticity of DOM (Weishaar et al., 2003):

269 (4) SUVA₂₅₄ (mg C L⁻¹ m⁻¹) =
$$A(254)$$
 (m⁻¹)/ DOC (mg L⁻¹)

FDOM: The 3-D recorded spectra were corrected for the instrumental biases both for excitationand emission using correction curves provided by the manufacturer (Stedmon and Bro, 2008).

272 Additionally, spectra were corrected against a Milli-Q water blank run every day before the samples 273 to remove water Raman peaks. No correction for inner filter effects was applied to our data as for each sample the relative $a(\lambda)$ value was below 10 m⁻¹ (Lawaetz and Stedmon, 2009;Stedmon and 274 Bro, 2008). As an example, a(254) was on average $2\pm 2 \text{ m}^{-1}$ for SML and $1.6\pm 1.3 \text{ m}^{-1}$ for underlying 275 276 water (ULW) samples. Fluorescence spectra were normalized to Raman Units (R.U.) by integrating 277 the Raman peak of 350 nm Ex and 382 to 407 nm Ex extracted by the daily Milli-Q water blank. 278 Calibration to R.U. was done with the FDOMcorrect toolbox for Matlab (The MathWorks Inc.) 279 incorporated in DrEEM toolbox (Murphy et al., 2013). We choose to normalize to R.U. as these 280 units are widely used in open ocean measurements and we could compare our results.

281 PARAFAC analysis was applied to EEMs in order to identify and quantify independent underlying 282 components of the CDOM pool, and was performed by the N-way toolbox for Matlab in DrEEM 283 (Murphy et al., 2013). After normalization to R.U. units, data were smoothed to remove remaining 284 scatter peaks, Raman and Rayleigh signals by creating a sub-dataset. We then performed a 285 preliminary outlier analysis generating models with 3 to 7 factors with non-negativity constraints, 286 comparing the spectra to unconstrained models. When dilution dominates the dataset, components 287 are strongly correlated. To investigate biases due to dilution, we performed a test for correlations 288 between the components, as suggested by the DrEEM tutorial by Murphy and colleagues (2013). 289 We then normalised the dataset by the DrEEm function *normeem* to reduce the co-linearity related 290 to the concentration, thus giving low-concentrated samples a possibility to enter the model, 291 followed by the outlier test again on the normalised data. After visually comparing the spectra and 292 looking at the error residuals for models with 4 to 7 components, we then compared the models by 293 the sum of squared errors (SSE) expressed as a function of wavelength, choosing the models with 294 lower SSE. At this stage, we choose models with 5, 6 and 7 components and reversed the 295 normalization to obtain the unscaled scores before validation. Models with 5, 6 and 7 components 296 were validated by split half analysis " $S_4C_6T_3$ " (see Murphy et al. 2013) where it was ensured that in 297 each test the dataset halves being compared had no samples in common. The validation was

successful for 5-components model, for all comparison. The maximum fluorescence intensities of the five fluorophores at specific Ex-Em wavelengths ranges are described in table 3. Figures with the model comparison for both excitation and emission for the 5-components model are included in the supplementary material (Figures S1 and S2).

302 In fluorescence spectroscopy, the humification index (HIX), first introduced by Zsolnay et al. (1999), is a powerful tool to study CDOM dynamics in soils, as humification is associated with a 303 304 shift to longer emission wavelengths (Senesi, 1990). It has been first applied to aquatic CDOM in 305 estuarine waters by Huguet and colleagues (2009), and is calculated as the ratio H/L of two spectral 306 region areas of the emission spectrum scanned at 254 nm excitation. Area L is calculated between 307 the emission wavelengths 300 nm and 345 nm, and area H between 435 nm and 480 nm. When the 308 degree of aromaticity of CDOM increases, the emission spectrum at excitation 254 nm is shifted 309 towards the red (longer wavelengths), implying an increase in H/L ratio and in HIX. High HIX 310 implies maximum fluorescence intensity at long wavelengths and therefore the presence of complex 311 molecules like HMW aromatic CDOM (Senesi et al., 1991). We applied a slight modification to the 312 HIX index for our samples, introducing the "SMHIX" index, where SM stands for Surface 313 Microlayer. As we did neither have the scanned excitation wavelength of 254 nm, nor the scanned 314 spectrum at excitation 345 nm and 435 nm, we calculated SMHIX index as follows:

315 (5) SMHIX =
$$(\sum I_{434 \to 480}) / \sum I_{300 \to 346})$$

Where $\sum I_{434\to480}$ is the sum of all fluorescence intensities at every emission wavelength between 434 nm and 480 nm, and $\sum I_{300\to346}$ is the sum of all fluorescence intensities at every emission wavelength between 300 nm and 346 nm, both scanned with excitation = 255 nm.

319 Enrichment Factors: Enrichment Factors (EF), allow tracking of accumulation patterns of any 320 compound in the SML with respect to the underlying water (ULW) and comparison among 321 different compounds. EF are calculated according to the following:

322 (6) $EF = [X]_{SML} / [X]_{ULW}$

Where [X] is the concentration of a given parameter in the SML or ULW, respectively (GESAMP, 1995). EF > 1 indicates an enrichment, EF < 1 indicates a depletion in the SML. EFs are normally used for quantitative parameters, i.e., measured in abundance and concentration such as DOC, DHAA, CDOM, marine gels and cell abundances. Here, we applied the EF calculation for qualitative ratios and indexes too, like *S*(275-295), SR, SMHIX, SUVA₂₅₄, DHAA%-DOC. We kept the same wording, which is useful to describe differences between SML and ULW for both quantitative and qualitative parameters.

330 Statistical tests in data analysis have been accepted as significant for p < 0.05. Calculations, 331 statistical tests and illustration were performed with Microsoft Office Excel 2010, Sigma Plot 12.0 332 (Systat), Prism (GraphPad), Ocean Data View and Matlab R2009b (The MathWorks Inc.).

333

334 3. Results

Results on dissolved organic carbon and amino acids, gel particles (TEP and CSP), phytoplankton and bacterial abundance and the relative enrichment of these components in the SML of our sampling sites have been described elsewhere (Engel and Galgani, 2016). Here, we focus on the optical properties of DOM to identify possible sources, sinks and dynamics in the SML and underlying water of the Peruvian upwelling region.

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3.1. CDOM optical absorption properties

In the upwelling region off Peru, values for CDOM absorption coefficient a(325) ranged from 0.09 to 1.47 m⁻¹ in the SML and from 0.07 to 1.47 m⁻¹ in ULW. Highest values were observed at stations S10_1 to S10_4 along the coast for both SML and ULW. CDOM was enriched in the SML at most stations (Figure 3), with median EF for a(325) = 1.2 in a range varying between 0.4 and 2.8. A median EF = 1.2 means that at least 50% of our observations accounted for a CDOM-enriched SML. Besides the southern transect, higher EF values were observed at the northern stations S2 and S2_2, and in the southern coastal upwelling stations S15_1 to S15_3. Lower EFs and EFs < 1, indicating a depletion of CDOM in the SML, were observed at higher distance from the coast(Figure 4).

The spectral slope parameter between 275 and 295 nm (S(275-295), nm⁻¹) is a good indicator for 351 352 CDOM molecular weight as an increase of this parameter indicates decreasing molecular weight, 353 thus revealing accumulation or degradation processes of bioavailable CDOM (Helms et al., 2008). In our samples, S(275-295) ranged from 0.012 to 0.038 nm⁻¹ in the SML and from 0.017 to 0.043 354 nm^{-1} in ULW. In general, S(275-295) was quite similar between SML and ULW, and no statistically 355 356 significant differences were found between SML and ULW for S(275-295). Higher spectral slopes 357 were observed in the ULW of the southern stations below 15°S (S19, S19_2, S20, S20_2, S1778). 358 In the coastal stations \$10 1 to \$10 4 and \$14 1 to \$15 3 lower \$(275-295) values were 359 determined for both SML and ULW. Median enrichment factor (EF) for S(275-295) was 1 (Figure 360 3), thus indicating similar molecular weight of CDOM compounds in the SML and ULW. Lower 361 EFs were observed in the northernmost and southernmost stations and along the coast.

362 The SUVA₂₅₄ and SMHIX indexes are related to the degree of CDOM aromaticity and to its humic content, respectively. In our study, SUVA254 ranged from 0.49 to 1.74 mg C L⁻¹ m⁻¹ in the SML, 363 364 with highest values at coastal southern stations S10 1 to S10 4 and S14 1 to S17 2. Similar values were recorded for ULW, ranging from 0.49 to 1.21 mg C L⁻¹ m⁻¹. Generally, SUVA₂₅₄ values in our 365 samples were comparable to the Pacific Ocean with a typical SUVA₂₅₄ of 0.6 mg C L^{-1} m⁻¹ (Helms 366 et al., 2008;Weishaar et al., 2003). Median EF for SUVA254 was 1.1, with higher values in 367 368 correspondence of northern stations and coastal southern stations (S2, S2_2, S15_1 to S15_3 and 369 S19 to S1778) where the higher EF for a(325) were also observed (Figures 3 and 4). SMHIX 370 ranged from -1.33 to 2.05 for SML and from -0.1 to 3.03 for ULW, with highest values in ULW. 371 Enrichment factors showed an overall depletion of high-humic acid containing CDOM in the SML 372 (Figure 3), with median EF = 0.8. Higher humic acid enrichment in the SML was observed on the 373 southern transect S19 to S1778 (Figure 4), where we recorded the highest enrichment of CDOM (as 374 *a*(325)) as well.

15

The carbon-normalized yields of dissolved amino acids (DHAA%-DOC) as indicator of DOM diagenetic state, ranged from 1.4% to 8.1% in SML samples and from 0.9% to 3.6% in ULW samples, indicating relatively more labile DOM in the SML. This observation was supported by the enrichment factors (EF), which showed a general enrichment of more labile DOM in the SML (Figure 3), with median EF values for DHAA%-DOC of 1.5. Highest EFs were recorded in the northernmost stations S1 to S3, and on the southernmost transect S19 to S1778.

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3.2. PARAFAC analysis for CDOM fluorophores

383 Five optically active components were identified by PARAFAC analysis with the DrEEM toolbox 384 in Matlab (Murphy et al., 2013), hereafter named F1, F2, F3, F4 and F5 (Figure 5). The spectral 385 characteristics of the five identified components were compared to previous studies as described in 386 table 3. F1 had an excitation range of 250-290 nm with emission peaks between 320 and 350 nm, 387 which corresponds to peak T of the amino-acid like fluorescence of tryptophan, derived by *in-situ* 388 primary production (Coble, 1996). This component (F1) was generally enriched in the SML 389 (Figures 6, 7) with a median EF = 1.5, between a minimum EF of 0.5 and a maximum EF of 3.3. 390 Potential loss processes of this compound are its destruction by UV light and microbial degradation 391 (Stedmon and Markager, 2005b). F1 has also been related to protein-like fluorescence of 392 extracellular polymeric substances (Liu and Fang, 2002). Fluorescence intensities of F1 were the 393 lowest compared to the other fluorophores, but significantly higher in the SML compared to the 394 ULW (Mann-Whitney Rank Sum Test, p < 0.001, n = 38). Both in SML and ULW, fluorescence 395 intensities of F1 were positively correlated to components F3, F4 and F5 (Spearman Rank Order 396 Correlation coefficient C = 0.37, p < 0.001, n = 76 with F3; C = 0.41, p = 0.001, n = 57 with F4; C 397 = 0.38, p < 0.001, n = 76 with F5).

Component F2 had a short wavelength excitation range (250-260 nm) with emission at longer
wavelengths (500-520 nm), corresponding to peak A of fulvic acids and humic acids (Stedmon and
Markager, 2005a;Singh et al., 2010;Yamashita and Jaffé, 2008;Coble, 2007;Santín et al., 2009). F2

401 showed a regional enrichment in the SML, with highest values at the northernmost stations S2 to S3 402 and at stations S10_1 to S10_4 (Figure 6). F2 enrichment was not ubiquitous (Figure 7), with 403 median EF = 1, ranging from a minimum EF = 0.5 and a maximum EF = 3.6. F2 positively 404 correlated with bacterial abundance and temperature (Table 4) and to F3 and F5 components 405 (Spearman Rank Order Correlation coefficient C = 0.74, p < 0.001, n = 76 with F3, and C = 0.71, p 406 < 0.001, n = 76 with F5).

407 Component F3 was characterized by a clear excitation peak at 265 nm, with emission maxima in the 408 longer wavelength range 520-540 nm. Component F3 showed a median EF = 1.1 (minimum EF =409 0.3, maximum EF = 4.7), indicating a slight enrichment in the SML (Figure 7), with higher 410 accumulations close to the coast at stations S19 2 to S1778 and at the edge of the continental shelf 411 at stations S4 and S8 (Figure 6), in correspondence with the highest enrichment of gel particles in 412 the SML (Engel and Galgani, 2016). In our study F3 was positively correlated with the abundance 413 of bacteria, proteinaceous particles and increasing SUVA₂₅₄ (Table 4). It showed an inverse 414 correlation to salinity (Table 4). Besides F1 and F2, F3 was significantly correlated to F5 415 (Spearman Rank Order Correlation coefficient C = 0.87, p < 0.001, n = 76).

416 Component F4 was not detectable at all stations, but showed high enrichment in the SML close to 417 the coast and along the continental shelf at stations S10 1 to S10 4, S13 1 to S13 3, S14 1 to 418 S15 2 (Figure 6). F4 was generally enriched in the SML (Figure 7) with median EF = 1.5 (in a 419 minimum-maximum EF range of 0.4 - 14.9) and with significant differences in fluorescence 420 intensity compared to the ULW (Mann-Whitney Rank Sum Test, p < 0.001, n = 38). F4 featured 421 characteristics of an amino-acid like fluorophore with excitation/emission maxima in the range 250-422 265/284-320 in the fluorescence peak T region of tyrosine (Coble, 1996; Murphy et al., 2008; Aoki 423 et al., 2008; Yamashita and Tanoue, 2003) and phenylalanine (Yamashita and Tanoue, 424 2003; Jørgensen et al., 2011) (Table 3). F4 was negatively correlated to bacterial abundance (Table 425 4), and to slope ratio SR with SR = (S(275-295):S(350:400)). F4 was also negatively correlated to 426 SMHIX, indicating a low humic-acid content of this fluorophore. As for F1, it positively correlated

with SUVA₂₅₄ and DHAA%-DOC (Table 4). Interestingly, F4 showed the highest fluorescence
intensities among all samples.

429 Component F5 was quite difficult to identify, as we found no comparable spectra in the literature. It 430 showed typical characteristics of allochtonous humic-like material with excitation/emission ranges 431 in the peak A and C regions, which have been observed in bay and offshore waters (Mostofa et al., 432 2013). F5 had the highest fluorescence intensities both in the SML and ULW but was not clearly 433 enriched in one or the other compartment (Figure 7). EF ranged from a minimum of 0.5 and a 434 maximum of 3, with median value = 1.1. Highest enrichments in the SML were observed at 435 northern stations S4 and S4_2, at stations S10_1 to S10_4, and in the southern stations S20 to 436 S1778 (Figure 6). F5 was similar in characteristics to component F3, and positively correlated to 437 bacterial abundance and proteinaceous CSP particles (Table 4). Component F5 was also positively 438 correlated to all other fluorophores F1, F2, F3 as described, and to F4 (Spearman Rank Order 439 Correlation coefficient C = 0.34, p = 0.009, n = 57).

440 On the revisited stations, only component F1 showed a direct dependency on light exposure, 441 significantly decreasing in fluorescence – thus concentration – with increasing global radiation 442 intensity ($r^2 = 0.56$, p = 0.013, n = 10). Components F2 to F5 showed no significant change with 443 increased irradiation (Spearmank Rank Order Correlation analysis).

444

445 **3.3.** Changes in CDOM properties related to the biological and physical environment

Both in the SML and ULW, CDOM optical properties as absorption coefficient a(325), S(275-295), and SUVA₂₅₄ were compared to salinity, temperature, wind speed and particulate organic carbon (POC) (Table 5). Data on POC have been described in detail in Engel and Galgani (2016). CDOM absorption coefficient a(325) decreased at higher salinity, temperature and wind speed in the SML and ULW, with stronger dependency on these physical parameters in the SML (Table 5). In both compartments, there was a positive correlation of a(325) to POC. The spectral slope parameter S(275-295), indicator for DOM molecular weight, source, and degradation processes (Helms et al., 453 2008), increased at higher salinity and temperature (Figure 8d) in the SML and ULW. It did not 454 show any correlation to wind speed, but a significant negative correlation to POC in both 455 compartments (Table 5). Moreover, an increase of bacterial and phytoplankton cells led to a lower 456 S(275-295) both in the SML and ULW (Figures 8a, b). The dependency of S(275-295) on bacteria 457 in the SML (Spearman Rank Order Correlation Coefficient C = -0.59, p < 0.001, n = 35) was 458 stronger than in the ULW (C = -0.38, p = 0.02, n = 36), potentially indicating a higher bacterial 459 CDOM contribution. S(275-295) was also negatively correlated to phytoplankton abundance with a 460 stronger relationship in the ULW (C = -0.64, p = 0.001, n = 22) than in the SML (C = -0.47, p =461 0.004, n = 35). In the SML, we observed a significant decrease in S(275-295) with increasing 462 abundance of gelatinous proteinaceous particles (CSP) (Figure 8c), while in the underlying water a 463 lower S(275-295) was highly related to increasing concentration of polysaccharidic gels (TEP). In 464 both SML and ULW, higher salinity, temperature and wind speed were related to lower SUVA₂₅₄ 465 indexes, as indicators of DOM aromaticity. A positive correlation was observed instead between 466 SUVA₂₅₄ and POC (Table 5). An increment in temperature was inversely correlated to DOM 467 lability, and therefore bioavailability, expressed as DHAA%-DOC, indicating a higher degree of 468 DOM degradation (Spearman Rank Order Correlation coefficient C = -0.68, p < 0.001, n = 29 in the 469 SML and C = -0.66, p < 0.001, n = 29 in the ULW). DHAA%-DOC was also lower at higher 470 salinity (Spearman Rank Order Correlation coefficient C = -0.42, p = 0.02, n = 29 in the SML and C = -0.63, p < 0.001, n = 29 in the ULW). As for S(275-295), we observed similar trends in SR 471 472 (data not shown): SR was negatively correlated to DHAA%-DOC (Spearman Rank Order 473 correlation coefficient C = -0.50, p < 0.001, n = 75) and to both gel particles abundance (Spearman 474 Rank Order correlation coefficient C = -0.37, p < 0.001, n = 75 for TEP and C = -0.33, p = 0.004, n475 = 75 for CSP). SR did not show any significant correlation to total bacteria or phytoplankton 476 abundance, but was significantly lower in the SML, with a median EF = 0.9 (Mann-Whitney Rank 477 Sum Test, p = 0.013, n = 38). Furthermore, DHAA%-DOC was significantly higher in the SML 478 (Mann-Whitney Rank Sum Test, p = 0.036, n = 38).

479

480 **4. Discussion**

481 **4.1. CDOM enrichment and production in the top surface layer of the ocean**

482 The enrichment of organic material in the SML has been mainly related to biological processes in 483 the euphotic zone below the surface (Hardy, 1982;Liss and Duce, 2005). EBUS are among the most 484 productive regions in the ocean and therefore interesting systems to investigate the relationship 485 between organic matter accumulation and SML biogeochemical properties. The Peruvian EBUS is 486 associated with an extensive OMZ and a key region for the study of gas fluxes from the ocean 487 (Paulmier et al., 2008; Paulmier and Ruiz-Pino, 2009; Keeling et al., 2010). The presence of an 488 organics-enriched surface layer may strongly affect gas exchange between the marine and the 489 atmospheric systems (Engel and Galgani, 2016). The characterization of CDOM via its optical 490 properties adds relevant information to the organic matter composition in the SML, as it allows 491 discriminating between terrestrial and marine sources of DOM that may be equally enriched at the 492 surface. Moreover, it helps tracking changes in DOM "quality" deriving from higher DOM 493 exposure to solar radiation at the sea-surface than deeper in the water column. As such, microbial 494 and photochemical DOM turnover in the SML may contribute to the atmospheric emission of gases 495 and chemical reactive species, and interfere with the microbial carbon loop in the ocean.

496 In the Peruvian EBUS, we observed a general enrichment of CDOM in the SML with respect to the ULW, based on values of the specific absorption coefficient $a(\lambda)$ measured at 325 nm. Higher 497 498 values for CDOM absorbance were observed in the coastal upwelling stations characterized by 499 lowest salinity, temperature and highest enrichment of organic components, both in the particulate 500 and dissolved fraction (Engel and Galgani, 2016). It is commonly observed that spectral loadings of 501 allochtonous/terrestrial-like CDOM decrease with increasing salinity (Murphy et al., 2008). 502 However, we did not observe such trend in our samples. Instead, we found a negative correlation of 503 amino-acid like fluorophore F1 to salinity and temperature, and no clear enrichment of humic-acid 504 like fluorophores F2, F3 and F5 in the SML. Therefore, we think that in the SML of the study

505 region the contribution of terrestrially derived CDOM, if any, is overwhelmed by the high 506 productivity of the upwelling system. Organics enriched in the SML such as the amino-acidic 507 compounds F1 and F4 found at the upwelling stations may therefore reflect other processes rather 508 than input of allochtonous CDOM from land. DOC concentrations in the SML were related to DOC 509 concentrations in the ULW (Engel and Galgani, 2016), and the same was true for CDOM 510 absorption coefficient a(325) (Spearman Rank Order Correlation coefficient C = 0.82, p < 0.001, n511 = 38), implying a direct dependency of SML CDOM on the organic matter in the ULW (Zhang and 512 Yang, 2013). CDOM absorption coefficient a(325) as well as its spectral slope S(275-295) did not 513 show any correlation to changes in DOC concentrations neither in the SML, nor in the ULW, but 514 were significantly related to DOM diagenesis (DHAA%-DOC) POC, and abundance of autotrophic 515 and heterotrophic microorganisms suggesting a recent production of labile or semi-labile substrates 516 driven by *in-situ* microbial or photochemical processes in the underlying euphotic zone or at the 517 immediate sea surface. A closer look on CDOM spectral properties revealed significant differences 518 between SML and ULW. According to Helms et al. (2008), an increase in S(275-295) and SR 519 suggests CDOM photodegradation and decreasing molecular weight. DHAA%-DOC is used here 520 as an indicator for DOM diagenesis, thus, the extent of microbially altered DOM. The higher 521 DHAA%-DOC, the more labile, bioavailable, recent and less altered DOM in the sample. We 522 observed a negative correlation when comparing DHAA%-DOC and POC to S(275-295) and to SR. 523 The higher DHAA%-DOC, the lower S(275-295) and SR. Microorganisms adopt several strategies 524 against tough environments; the correlation between DHAA%-DOC to S(275-295) and SR was 525 stronger in the SML than in the ULW, suggesting an accumulation of HMW-DOM related to the 526 contribution of microorganisms directly in the SML or in the proximity due to cell lysis or 527 exudation, which has been previously proposed (Tilstone et al., 2010). Thus, the close correlations 528 of optical parameters to POC and marine gels lead to hypothesize that autochtonous CDOM 529 produced in the very surface ocean can actually be incorporated in the gelatinous organic carbon 530 pool.

531

532 **4.2. CDOM composition**

533 The analysis of EEMs allowed the identification of five fluorescent components both in the SML 534 and ULW, of which two (F1 and F4) showed an amino-acid like fluorescence of autochtonous 535 material, and three (F2, F3 and F5) had the characteristics of fulvic-acid like or humic-acid like 536 CDOM (Table 3). These classes of fluorophores are commonly found in marine environments 537 (Coble, 2007; Mostofa et al., 2013), but EEMs analyses of SML samples are scarce and up to now 538 revealed the enrichment in humic-acid like fluorophores only (Zhang and Yang, 2013). Phenolic 539 materials deriving from humic and fulvic acids transported by river drainage, and from macroalgae 540 polyphenols, are often enriched in the SML, and indicate the presence of surface slicks (Carlson, 541 1982;Carlson and Mayer, 1980). Here, we observed a significant enrichment of amino-acid like 542 fluorophores F1 and F4 with respect to ULW, in good accordance with previous reports on amino 543 acids enrichment in the SML (Kuznetsova et al., 2004;Cunliffe et al., 2013;Tilstone et al., 2010), 544 and with our own observations for the Peruvian EBUS (Engel and Galgani, 2016). F1 has shown 545 the greatest production rates during algal blooms, whereas its major sinks are UV light and 546 microbial degradation (Stedmon and Markager, 2005b). Moreover, it is assumed that F1 relates to 547 the fluorescence of amino acids still bound in the proteinaceous matrix (Stedmon and Markager, 548 2005b). Based on these previous findings and on our results (Table 4), we suggest that F1 is a 549 tryptophan-like fluorophore, originating by *in-situ* primary production, relatively labile as it features 550 an increase in fluorescence intensity correlated to increasing DHAA%-DOC, and possibly included 551 in the gel particles surface matrix. F4 showed very high fluorescence intensities compared to F1, 552 F2 and F3. In the literature, F4 has been associated to the fluorescence of amino acids in peptides 553 (Stedmon and Markager, 2005b). Similarly to F1, F4 showed a positive correlation to DHAA%-554 DOC, as to indicate its labile nature. The aromatic content of DOM is highly responsible for its photoreactivity (e.g. Mopper et al., 2014); F4 correlation to DOM lability (DHAA%-DOC) and 555 556 aromatic content (SUVA₂₅₄) was weaker than for F1. In our study, this may indicate F4 as an

557 intermediate product of photochemically-driven aggregation or microbial degradation of labile 558 CDOM. F4 has been linked to the fluorescence of tyrosine and phenylalanine (e.g. Coble, 559 1996; Murphy et al., 2008; Jørgensen et al., 2011) and both amino acids were enriched in the SML of 560 the Peruvian EBUS (Engel and Galgani, 2016). Recently, Cao and colleagues (2014) performed a 561 laboratory study aimed at understanding intermolecular interactions between N₂O and phenol 562 (C₆H₅OH), cresol (CH₃C₆H₄OH), and toluene (CH₃C₆H₅), which are representative aromatic 563 compounds and useful models of various biomolecules such as the aromatic amino acids tyrosine 564 and phenylalanine containing a benzene ring. Their experiment was carried out in a Ne matrix at 565 about 8 K (-265.15 °C) with millimolar concentrations of the aromatic compounds (Cao et al., 2014), therefore in a setting not at all comparable to our experimental setup. Cao and colleagues 566 found interesting π non-covalent interactions between N₂O and the aromatic compounds, 567 568 suggesting an interaction of N₂O with tyrosine and phenylalanine of great interest for biological 569 processes. We mention it here, as the presence of these specific amino acids in the SML of the 570 Peruvian EBUS may interfere with the exchange of N₂O between the ocean and the atmosphere, as 571 suggested previously (Engel and Galgani, 2016). The enrichment of fluorophores F1 and F4 in the 572 SML could be partly due to the upwelling of colder nutrient-rich waters that boost primary 573 production in the euphotic zone. Salinity and temperature gradients may thus explain the variation 574 of F1 in the SML (Table 4), reflecting local upwelling and DOM production. The observed 575 accumulation of amino-acid like CDOM may additionally derive from a local microbial release 576 within the SML itself due to cell disintegration, or as protection strategy for the exposure to UVB 577 light in a demanding environment (Ortega-Retuerta et al., 2009). Mycosporine-like amino acids 578 (MAAs), for example, serve as a natural microbial UV sunscreen against photodamage (Garcia-579 Pichel et al., 1993) and have been observed in enriched concentrations in the SML (Tilstone et al., 580 2010). Major losses of autochtonous protein-like fluorophores in the SML may be related to 581 photochemical and microbial degradation: negative correlations of F1 and F4 to SR may hint to 582 photochemical degradation, recalling that an increase in SR is usually related to photobleached

material (Helms et al., 2008). The negative correlation of F4 to bacterial abundance may be insteadan indication of a microbial sink of this fluorophore.

585 The fulvic acid or humic acid-like components F2, F3 and F5 were ubiquitous in SML and ULW, 586 with no significant differences in fluorescence intensities between the two compartments. F2 and F3 587 have been previously observed in coastal marine environments (e.g. Jørgensen et al., 2011;Ishii and 588 Boyer, 2012). In the literature, component F2 has been characterized as of terrestrial origin, 589 allochtonous in marine environments, found in bays, rivers and coastal waters. It is assumed to 590 reflect small-sized molecules, being resistant to photodegradation, biologically not available, and 591 mainly derived from photobleached terrestrial-like humic acids in marine waters with highest 592 concentrations near the surface (Ishii and Boyer, 2012). In this study we did not find a correlation 593 of F2 to global radiation but a positive correlation to temperature and to bacterial abundance (Table 594 4). We also observed an increase of bacterial abundance with increasing sea-surface temperature, 595 which is well supported by existing literature (e.g. Morán et al., 2015). Higher temperature also 596 stimulates the activity of marine bacteria (e.g. Piontek et al., 2009). Thus, as F2 probably reflects 597 the fluorescence of highly degraded small molecules, we may characterize F2 as the ultimate 598 product of microbial CDOM degradation in the surface ocean, not bioavailable anymore. F3-like 599 fluorophores have been identified as an intermediate product of terrestrially derived DOM, still 600 subject to further photochemical degradation (Stedmon et al., 2007). Earlier studies attributed this 601 optical behavior to fulvic acid C-like components showing a peak in region A. According to Ishii 602 and Boyer (2012), F3 like components may comprise larger hydrophobic molecules that are 603 photodegradable by UV light, of terrestrial or microbial origin, biologically degraded and produced. 604 Moreover, F3 appearance has been related to apparent oxygen utilization (Yamashita et al., 2010), 605 further suggesting a microbial source of this material (Jørgensen et al., 2011). In this study, F3 606 showed a slight enrichment in the SML and was related to heterotrophic bacteria as well as to CSP 607 particles, possibly indicating its origin in microbial reworking of larger organic compounds. F5 608 showed characteristics of humic acid fluorophores, with fluorescence maximum ranges to the lower

609 end of F3 emission indicating a more pronounced CDOM alteration with respect to F3. Showing 610 similar correlations to heterotrophic bacteria and CSP, F5 may as well derive from a microbial in-611 situ reworking of larger organic molecules both in the SML and ULW contributing to the size 612 continuum and reactivity of the gel particles pool in surface waters. In fact, a net production and 613 accumulation of humic-like CDOM in surface waters may occur in upwelling regions (Nieto-Cid et 614 al., 2005; Jørgensen et al., 2011), whereas photochemical loss is thought to be the major removal 615 mechanism of this material (e.g. Mopper and Schultz, 1993). In this study, fulvic acid/humic acid-616 like fluorophores well correlated among each other, suggesting a common underlying origin.

617 Based on CDOM absorption and fluorescence characteristics, we propose a conceptual model for 618 the control of CDOM production and loss in the SML and ULW by microbial and photochemical 619 processes (see graphical abstract). In this model, the accumulation of CDOM in the SML is the 620 result of a) the biological production of CDOM in the ULW and deeper water column, stimulated 621 by the upwelling of nutrient-rich waters to the sunlit surface and b) the local microbial release of 622 CDOM as a response to elevated solar radiation. Previous and our own observations on amino-acid 623 fluorophores (F1, F4), as well as on the enrichment of CSP and amino acids in the SML described 624 elsewhere (Engel and Galgani, 2016), suggest a rapid turnover of fresh DOM in the sea-surface 625 itself. On one hand, microbes release fresh DOM directly within the SML or in the upper first 626 centimetres, as a consequence of high light exposure. On the other hand, and both in the SML and 627 ULW, microbial and photochemical degradation would lead to the loss of amino-acid like 628 fluorophores (F1, F4) and to the accumulation of less labile and humic-like components completely 629 degraded (F2) or still subject to further photochemical degradation (F3, F5).

630

631 **4.3.** Implications for surface ocean dynamics and future perspectives

Optical properties of DOM in the Peruvian EBUS revealed a SML characterized by amino-acid like
CDOM fluorophores. CDOM enrichment in the SML has been observed in different marine regions
associated with enrichment in phenolic compounds, MAAs and humic acids (Carlson, 1982;Carlson

635 and Mayer, 1980; Tilstone et al., 2010; Zhang and Yang, 2013). MAAs for example (LMW-DOM) 636 are well known as microbial sunscreen in aquatic environments (Bhatia et al., 2011;Shick and 637 Dunlap, 2002), and were observed in higher concentrations in the SML during surface slicks 638 development (Tilstone et al., 2010). Here, the accumulation of amino-acid like CDOM may have a 639 major microbial source directly in the SML or the immediate subsurface water, whereas fulvic 640 acid/humic acid-like CDOM likely originated in the sunlit zone below by microbial and 641 photochemical processing of upwelled organic material. Accumulation of amino acids in the SML 642 has been related to a reduced bacterial activity, being the SML an extreme environment where the 643 consumption of amino acids may be lower (Santos et al., 2012). A reduced bacterial activity may 644 thus also explain the amino acids enrichment in the SML of the Peruvian EBUS (Engel and 645 Galgani, 2016). We may assume that in the top layer of the ocean, and at higher extent in the SML, 646 exposure to light may have determined three main processes: 1) microbial release amino-acid like 647 CDOM as a sunscreen function, 2) increased availability of biological substrate by CDOM 648 photolysis and 3) further photochemical degradation of microbially-altered CDOM. Photochemistry 649 is able to alter the HMW fraction making it more available for microbial attack (Kieber et al., 650 1989), but at the same time it may lead to a net loss of bioavailable substrates (Kieber, 2000). 651 Therefore, the interplay of photochemistry and microbial activity controls the accumulation and loss 652 of organic compounds at the sea-surface, implying consequences on gas fluxes worth deeper 653 investigations in climate-relevant marine regions such as the OMZ off Peru. As an example, high 654 microbial DOM respiration can lead to higher production of CO₂ in the SML (Garabétian, 1990), 655 whereas high concentrations of isoprene may be released from photosensitized DOM reactions in 656 the SML, proving an abiotic source of this gas uncoupled from biological production (Ciuraru et al., 657 2015).

658 It remains unclear whether in the Peruvian EBUS an increase in bioavailable carbon may have 659 implied a higher heterotrophic respiration and CO_2 production in the SML, and this is an attractive 660 hypothesis for future studies in this direction. It may be suggested however, that a net DOM 661 production in the SML may take place independently of the biological productivity of the 662 underlying waters as a sole microbial response to light exposure. We assessed the enrichment of 663 light-absorbing proteinaceous organic material in the SML of a highly productive oceanic system, 664 which may interfere with correct estimates of primary production from remote measurements. To 665 conclude, we suggest that further primary production estimates may take into account the CDOM 666 enrichment in the first centimeters of the ocean.

667

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Tables

| | Salinity [PSU] | Temperature [°C] | Global Radiation [W m ⁻²] | Wind Speed [m s ⁻¹] |
|---------|----------------|------------------|---------------------------------------|---------------------------------|
| Average | 34.9 | 19.2 | 539 | 5.5 |
| SD | 0.2 | 1.7 | 352 | 2.1 |
| Min | 34.4 | 15.9 | 10 | 0.6 |
| Max | 35.3 | 21.9 | 1088 | 9.0 |

Table 1. Data on average, maximum and minimum salinity, water temperature, global radiation and wind speed during M91. Data were retrieved from Dship data server of R/V Meteor.

Table 2. Stations with multiple measurements. Metadata with date, local and UTC time of sampling, coordinates, and average global radiation retrieved from Dship data server of R/V Meteor.

| Station Ship ID | Nr. | Station nr. | Samples | Date | Time [UTC] | Time [Local] | Lat, S [°] | Long, W [°] | Average Global Radiation [W m ⁻²] |
|--------------------|-----|----------------|---------|----------|---------------|-----------------|------------|-------------|--|
| 1733-5 | 1 | S7 | sml/ulw | 08-12-12 | 11:30 | 6:30 | 9°31.258' | 79°17.886' | 10 |
| 1733-9 | 1 | S7_2 | sml/ulw | 08-12-12 | 19:45 | 14:45 | 9°32.75' | 79°18.43 | 837 |
| 1752-2 | | S12_1 | sml/ulw | 13-12-12 | 12:00 | 7:00 | 12°55.20' | 78°.42.00' | 380.5 |
| 1752-7 | 2 | S12_2 | sml/ulw | 13-12-12 | 20:30 | 15:30 | 12°59.79' | 78°.41.00' | 704.5 |
| 1752-9 | | S12_3 | sml/ulw | 13-12-12 | 23:10 | 18:10 | 12°55.20' | 78°.42.03' | 47 |
| 1764-4 | | S16_1 | sml/ulw | 17-12-12 | 12:40 | 7:40 | 14°7.708' | 76°52.759' | 381 |
| 1764-6 | 3 | S16_2 | sml/ulw | 17-12-12 | 17:40 | 12:40 | 14°11.11' | 76°55.95' | 1043 |
| 1764-9 | | S16_3 | sml/ulw | 17-12-12 | 22:00 | 17:00 | 14°11.10' | 76°55.99' | 161.5 |
| 1777-2 | Λ | S20 | sml/ulw | 22-12-12 | 18:00 | 13:00 | 15°31.174' | 75°36.015' | 1088 |
| 1777-10 | 4 | S20_2 | sml/ulw | 23-12-12 | 15:00 | 10:00 | 15°36.42' | 75°38.60' | 1046 |

Table 3. Fluorescent components identified in this study in both SML and ULW samples, according to their Ex/Em maxima ranges (nm), maximum fluorescence intensity range Fmax (R.U.), corresponding peaks individuated in previous studies (peak name, region, Ex/Em ranges) and properties.

| Components of this study | Ex/Em maxima [nm] | Fmax range [R.U.] | Literature peak name (region, Ex/Em) | Reference | Properties |
|--------------------------|----------------------|-------------------------|--|-----------|--|
| F1 | 250-290/ 320-350 | 0.001- 0.228 | (T) (275/340) | А | Protein-like fluorescence of tryptophan Autochtonous material. Source: in situ primary production |
| | | | 6(B) (280/338) | В | Protein-like fluorescence of tryptophan, autochtonous material. Source: algal growth. Sink: microbial reworking, UVB. |
| | | | T (280-285/340- 350) | С | Protein-like, extracted from EPS. |
| F2 | 250-260/ 500-520 | 0.048- 1.709 | 2(A) (250/504) | D | Fulvic acid C-like allochtononus material present in all environments. Terrestrial/autochthonous fulvic acid fluorophore group. |
| | | | 1(A) (250/520) | Е | Fulvic acid C-like. Bay waters, allochtonous. |
| | | | 2(A) (<260/>500) | F | Humic Acid C-like, river and coastal waters, allochtonous. Terrestrial humic. |
| | | | 1(A) (<230-260/400- 500) | G | Small sized molecules, photoresistant and biologically not available. Source: photochemistry, terrestrially derived humic acids in marine waters, highest concentrations near the water surface. |
| | | | 2(A?) (250/504) | Н | UVA humic-like, fulvic acid, terrestrial, autochtonous |
| | | | C2(-) (256/>500) | Ι | Humic acid C-like, estuaries of the Iberian peninsula, allochtonous. |
| F3 | 265/520-540 | 0.019- 1.640 | 2(A+C) (<240-275/434- 520) | G | Larger molecules, hydrophobic compounds, photodegradable by UVA light. Source: terrestrial or microbial, intermediate inputs of minimal exposure to sunlight, biologically degraded and produced. |
| | | | C1 (~275/400-550) | L | Humic-like CDOM microbially produced. |
| | | | 1(A/C) (<260/466) | 0 | Humic-like CDOM oxidized <i>in situ</i> by microbial processes. |
| F4 | 250-265/ 284-320 | 0.002- 6.507 | (T) (275/300) | J | Protein-like fluorescence of tyrosine. Autochtonous material. Source: in situ primary production, North Pacific and Atlantic Ocean. |
| | | | 4(T) (275/306(338)) | В | Fluorescence of tryptophan and tyrosine in peptides. Greatest production rates during establishment of algal bloom. Source: algae in exponential growth phase. Sinks: not identified (microbial uptake or aggregation?) |
| | | | (B) (275/310) | А | Tyrosine-like, marine waters, autochtonous. |
| | | | C(T) (270-290/250- 365) | К | Autochtonous protein-like hydrophobic acid fraction from phytoplankton cultures. |
| | | | C3(T) | L | Protein-like fluorescence of phenylanine. |
| | | | Standard (255-265/284- 285) | М | Protein-like fluorescence of phenylanine. Source: standard. |
| | | | (B) (265-280/293- 313) | М | Protein-like fluorescence of tyrosine. Source: autochtonous. |
| F5 | 270-275/ | 0.023- | (A,C) | Ν | Humic acid C-like or A-like, allochtonous |

| 540-550 | 1.714 | (<260-270/>508) | material |
|---------|-------|-----------------|---------------------------|
| | | | in bay and marine waters. |

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Table 4. Spearman Rank Order Correlation coefficients (*C*) between fluorescent components (F1-5) and total bacterial and phytoplankton cells, TEP and CSP particles, SUVA₂₅₄, *S*(275-295), *SR*, *a*(325), DHAA%-DOC, SMHIX, salinity and temperature measured in our study, both in the SML and ULW. Statistical significance was accepted for p < 0.05. n = number of samples. Only statistically significant correlations are shown. Bold characters indicate negative correlations.

| Component [R.U.] | Statistics | Bacteria [cells mL ⁻¹] | Phytoplankton [cells mL ⁻¹] | TEP [L·1] | CSP [L ⁻¹] | SUVA ₂₅₄ [mg C L ⁻¹ m ⁻¹] | <i>S</i> (275-295) [nm ⁻¹] | SR | a(325) [m ⁻¹] | DHAA%- DOC [%] | SMHIX | Salinity [psu] | Temperature [° C] |
|---------------------|-------------|---------------------------------------|--|--------------|---------------------------|--|---|------------------------|------------------------------|----------------------|--------------------------|------------------------|----------------------|
| E1 | С | | 0.285 | 0.281 | | 0.620 | - 0.257 | - 0.387 | 0.406 | 0.696 | - 0.342 | - 0.261 | - 0.323 |
| F1 | р n | - | 57 | 0.014 76 | | < 0.001 76 | 0.025 76 | < 0.001 75 | < 0.001 76 | < 0.001 76 | 0.003 76 | 0.023 76 | 0.004 76 |
| F2 | C p n | 0.393 < 0.001 71 | - | - | - | - | - | - | - | - | 0.225 0.050 76 | - | 0.238 0.038 76 |
| F3 | C p n | 0.355 0.002 71 | - | - | 0.411 < 0.001 76 | 0.305 0.007 76 | - 0.221 0.055 76 | - 0.226 0.051 76 | - | - | - | - 0.273 0.017 76 | - |
| F4 | C p n | - 0.409 0.003 52 | - | - | - | 0.346 0.008 56 | - | - 0.410 0.002 56 | - | 0.392 0.008 56 | - 0.536 < 0.001 57 | - | - |
| F5 | C p n | 0.270 0.023 71 | - | - | 0.402 < 0.001 76 | - | - | - | - | - | - | - | - |

Table 5. Spearman Rank Order Correlation (*C*) between CDOM optical properties both in the SML and ULW with salinity (C_{PSU}), water temperature (C_T), wind speed (C_U) and particulate organic carbon (C_{POC}). Significant correlations (p < 0.01) are marked in bold (except ^a = p < 0.05). *n* is the number of samples, except * = 36 samples.

| SML | $C_{\rm PSU}$ | CT | Cu | C _{POC} | n |
|--|--|-----------------------------------|--|---------------------------------------|---------------|
| CDOM <i>a</i> (325) | -0.420 | -0.728 | -0.535 | 0.579 | 38 |
| S(275-295) | 0.640 | 0.616 | 0.318 | -0.597 | 38 |
| SUVA ₂₅₄ | -0.380 ^a | -0.634 | -0.460 | 0.537 | 38 |
| | | | | | |
| ULW | $C_{\rm PSU}$ | CT | Cu | C _{POC} | n |
| ULW CDOM <i>a</i> (325) | С _{РSU} -0.329 ^а | С _т -0.637 | C _U -0.386 ^a | С _{РОС} 0.656* | n 38 |
| ULW CDOM <i>a</i> (325) <i>S</i> (275-295) | C _{PSU} -0.329 ^a 0.493 | С _т -0.637 0.613 | <i>C</i> _U -0.386 ^a 0.24 | С _{РОС} 0.656* -0.622* | n 38 38 |

Figures' legend

Figure 1. Maps showing all sampled stations. Stations with multiple measurements are: (1) S7/7_2, (2) S12_1/3 and S12_2, (3) S16_1, S16_2/3, (4) S20 and S20_2. The locations of S7 and S7_2; S12_1 and S12_3; S16_2 and S16_3 coincide, as sampling was performed at different times.

Figure 2. CDOM absorption coefficient a(325), $[m^{-1}]$, in SML and underlying water (ULW) and spectral slope parameter between 275 and 295 nm, S(275-295), $[nm^{-1}]$.

Figure 3. Box and Whiskers plot of enrichment factors for CDOM absorption coefficient a(325), aromaticity (SUVA254), DOM diagenetic state (DHAA%-DOC), spectral slope S(275-295), and modified surface microlayer humification index (SMHIX). The horizontal lines of the boxes represent 25%, 50% (median) and 75% percentiles (from bottom to top). In the boxes, crosses represent the mean. Whiskers represent minimum and maximum values, and circles are outliers. Outliers are staggered to better visualize them. To identify the station, see outliers' labels and color legend. For a(325), SUVA254 and S(275-295) n = 38. For SMHIX, n = 37 and for DHAA%-DOC n = 29.

Figure 4. Enrichment factors (EF) in the Peruvian upwelling region. From the top left, EF for absorption coefficient measured at 325 nm both in SML and ULW, spectral slope parameter *S*(275-295) as indicator for changes in DOM molecular weight, SUVA254 as indicator for DOM aromatic content, DHAA%-DOC as indicator of DOM lability, and SMHIX as indicator of humic content of DOM.

Figure 5. (Above) Contour plots of five fluorescent components as identified by PARAFAC analysis and (below) relative spectral loadings of overlaid spectra for the 5-components model validated with 3 split comparisons. The axes of contour plots have been scaled to better visualize the fluorescence intensities (R.U.). A figure with the complete spectrum is included in the supplementary material (Figure S3). The dashed black line in the spectral loadings indicates excitation maxima for each component, the solid black line indicates emission peaks.

Figure 6. Distribution of enrichment factors (EF) for fluorescent components F1, F2, F3, F4, F5 identified in this study. Maximum EF for F4 has been recorded at station S10_2, with a value of 14.9. For visualization purposes, this data point is not included in the figure and fluorescence intensities have been scaled down to a maximum EF = 6.

Figure 7. Box and Whiskers plot of enrichment factors for fluorescent components F1, F2, F3, F4 and F5. The horizontal lines of the boxes represent 25%, 50% (median) and 75% percentiles (from bottom to top). In the boxes, crosses represent the mean. Whiskers represent minimum and maximum values, and circles are outliers. Outliers are staggered to better visualize them. To identify the station, see outliers' labels and color legend. For F4, n = 24. For all other components, n = 38.

Figure 8a-d. (a) Linear regression between bacterial abundance $[10^{6} \text{ cells mL}^{-1}]$ and spectral slope $S(275-295) \text{ [nm}^{-1}]$ in SML and ULW. (b) Linear regression (ULW) and Spearman Rank Order Correlation (SML) between phytoplankton abundance $[10^{4} \text{ cells mL}^{-1}]$ and spectral slope $S(275-295) \text{ [nm}^{-1}]$. (c) Linear regression between CSP abundance $[10^{8} \text{ particles L}^{-1}]$ and spectral slope $S(275-295) \text{ [nm}^{-1}]$ in the SML and between TEP abundance $[10^{8} \text{ particles L}^{-1}]$ and spectral slope $S(275-295) \text{ [nm}^{-1}]$ in the ULW. (d) Linear regression between temperature [°C] and $S(275-295) \text{ [nm}^{-1}]$ in SML and ULW. Black triangles: SML, open dots: ULW.



Figure 8



Figure 9



Figure 10



Figure 11



Figure 12



Figure 13



Figure 14



Figure 15