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Interactive comment on “Subaqueous speleothems (Hells Bells) formed by the interplay of pelagic redoxcline biogeochemistry and specific hydraulic conditions in the El Zapote sinkhole, Yucatán Peninsula, Mexico” by Simon Michael Ritter et al.

Simon Michael Ritter et al.

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First, the authors are very grateful for the constructive comments and suggestions by anonymous referee #1 and fully appreciate the time and effort taken for reviewing the manuscript. We addressed all issues raised by the referee in a point by point response below.

Referee's major comments:

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Referee: First, I think the authors could expand the introduction beyond speleothems to draw in a wider readership. Hells Bells have been widely speculated about informally by karst researchers for several years, but I think that there is broader biogeochemical story that would be of interest to critical zone and coastal aquifer scientists. The authors don't have to take my suggestion in order for me to recommend publication, but I do think it would help broaden the appeal of the work.

Response: We agree with the referee that there is a wider biogeochemical story to the formation of Hells Bells, but suggest to wait for more data which are not yet available. We have taken additional samples for genome analysis of water and the surface of Hells Bells which are currently under evaluation and we are hopeful that these results will lead to a wider field of interpretation.

Referee: The second comment is a bit more substantive. I was not convinced that Hurricane activity would be likely to be responsible for sustained upward migration of the mixing zone necessary for Hells Bells to precipitate in areas above the modern redoxcline. The authors might explore droughts, which are well-documented in the Holocene (see Hodell et al., 2001, reference included in comments below), as an alternative process. Droughts would thin the freshwater lens and elevate the mixing zone for prolonged periods of time needed for the slow precipitation of calcite mentioned in the paper to form large Hells Bells features.

Response: We agree that the well documented droughts for the Yucatán Peninsula and a subsequently thinner fresh water layer brought up the halocline and certainly had its effect on Hells Bells formation. We added this point in the MS on page 24 lines 16–18.

We are currently working on the Hells Bells from different angles and produced an amount of data that cannot be shown in one manuscript only. The hypothesis of a recharge-driven upward migration of the halocline was developed integrating all information available, including some that is not shown here. Manuscripts presenting this data are currently in preparation.

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The main argument why we did not consider droughts as a mechanism of halocline elevation is U/Th age-dating on Hells Bells specimens of different water depths (MS currently in preparation) show approximately identical young ages (~ 150 a) at the lowermost crystal tips (1-2 mm) of the Bells. There is even a weak trend of the youngest samples found in the lowest water depths and the oldest samples found in greater water depths. This makes droughts or prolonged periods of time with an elevated halocline as the sole mechanism for the elevation of the halocline unlikely because this should be reflected in an age-zonation of the Hells Bells.

We are aware of the fact that the described behavior of the halocline in response to recharge is counter-intuitive and that the cited reference (Escolero et al., 2007) does not explain sufficiently why the halocline should be elevated as a response to recharge events. Yet, such a behavior is well documented by their results and additionally we discovered an online available thesis investigating dynamics of the coastal Karst aquifer around Merida that reports a similar halocline elevation in response to recharge events (Heise, 2013). In general, however, there is a scarcity of studies addressing aquifer responses to recharge events in the area and there is only one other that we know of, which is Kovacs et al., (2017) and this study does not have information about the halocline elevation as it only captured data from the fresh water layer.

We try to fill this gap of data and knowledge as we are currently gathering continuous data on the halocline elevation at El Zapote cenote. Preliminary results of one year of measurement are very interesting and reveal a complex hydraulic behavior. Generally, the halocline elevation follows the elevation of the fresh water lens which seems to be controlled by the mean tidal sea level throughout the year which leads to a variation of ~ 0.5 m of the halocline elevation. Furthermore, in general, if the freshwater lens gets thinner (water level of the fresh water drops), the halocline rises (determined by increasing conductivity of the data logger in the halocline) and vice versa. However, we detected several short-termed events (~ 1 day) of an abrupt elevation of the fresh water accompanied by an elevation of the halocline as a response to precipitation/recharge

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(Fig .1). Subsequently, the halocline slowly drops as expected from an increase of the fresh water lens due to precipitation. So far, there has not been a major precipitation event (hurricane) within the recorded available data (from Dec 2017- Dec 2018), but the observed effect of an initial halocline rise in response to recharge should increase in scale and time with an increase of the precipitation/recharge event.

Caption to Fig. 1: Preliminary results of data loggers from El Zapote cenote showing two examples of halocline elevation events. Water level was measured in the fresh water layer in 15 min intervals with a temperature and depth Logger (TD-Logger) in ~6 m water depth. In the halocline conductivity, temperature and depth was measured in ~38 m water depth also in 15 min intervals (CTD-Logger). The plots show the deviation of the respective elevation with respect to the start of measurement in December 2017. The halocline elevation is calculated from the conductivity values through a linear regression of the electrical conductivity within the interval of 37–40 m water depth. Note the periodic data noise in the halocline elevation, which always occur between 10–16h almost every day. This is most likely caused by diving and or by jumping activities into the cenote. Both cause small turbulences in the stratified water column.

Referee's minor comments:

Referee: Abstract Line 20: Given that this is a sinkhole, I'm uncomfortable with the author's description of this chamber as lightless in the absence of measurements of light intensity over annual timescales. While it may appear dark to divers, there may be some "light" that still makes it to the chamber during some portions of the year.

Response: You are right, we cannot rule out that there might be some portions of light at some time of the year. We even tried to measure the amount of light with water depth at the cenote around noon in December 2017. We found no measurable light below ~16 m water depth (<11 Lux). However, we cannot exclude that small portions of light may penetrate deeper into the water body for some daytime and season.

To meet the referee's comment we replaced "lightless" with "dark" in Abstract Line 20.

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Referee: Abstract Line 32: It would be nice if the authors could include information about why Hell's Bells are not more widely distributed in coastal aquifers in general (and the Yucatan more specifically) in the abstract. Improved information and word economy in the preceding sections should create space for this insertion.

Response: To meet the referee's suggestion we restructured the sentences and added some Information about the prerequisites for Hells Bells formation.

Additionally, we are addressing that issue in a MS (in prep.) in which we compare several cenotes with Hells Bells and cenote Angelita. One of the aspects of this study will be the spatial distribution of cenotes with Hells Bells

Referee: Page 2 Line 1: I've been working on karst research problems for almost 20 years and have never heard the term "pending speleothems." That doesn't mean the term does not exist, but it probably means it is not common.

Response: "Pending" was used to refer to the group of pendant speleothemes in this context, however we followed the referee's concerns and deleted "Pending" and changed the according sentence into "Speleothems, such as stalactites or dripstones, result from physicochemical...".

Referee: Page 3 lines 18-19: I don't understand why a recharge event would elevate the freshwater lens. It should depress it, as recharge would increase the thickness of the lens:. The manuscript that was cited wasn't much help in explaining the process either.

Response: We try to be more cautious around the halocline dynamics throughout the MS and therefore change the lines 18-19 on Page 3 to "Although Moore et al. (1992) and Stoessell et al. (1993) report that the thickness of the freshwater lens does not vary significantly between seasons or on a yearly basis, local and short-termed variations are possible and were reported by Escolero et al. (2007), who documented a significant halocline elevation of up to 17.5 m in between two measurements in the years 2000

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and 2003."

Referee: Page 10, Line 3: I suggest "oxygenated" or "aerobic" as an alternative to "oxygenized"

Response: Changed to "oxygenated".

Referee: Page 11, Figure 3: The symbols change in panel C but there is no corresponding legend. I'm guessing that the circles are still winch collected and the squares are diver collected, but it would be helpful if the symbols remained constant across each panel.

Response: We agree that these symbols are misguiding. In panel c the squares and circles stand for different parameters and it is not distinguished between samples taken by divers and those taken with the Niskin bottle. We will address this issue in the final revision once we received all referee comments.

Referee: Page 17 line 9: In this context, "though" is being used to contrast ideas and should not begin a sentence. I suggest restructuring this sentence and the previous one.

Response: Both sentences were restructured as suggested.

Referee: Page 18 line 9: Bells should be plural.

Response: Corrected as suggested.

Referee: Page 18 lines 13-15: I can agree that sulfide oxidation drives calcite precipitation but I'm still clear on why these features are only found here and not in any of the other sulfidic halocline caves in the Yucatan (or elsewhere), but hopefully this is discussed later in the text: : ...

Response: It is discussed later in the MS.

Referee: Page 19 lines 23-24: I find these sentences super confusing. Can you please

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rephrase them?

Response: The passage was rephrased as recommended.

Referee: Page 19 line 28: change to “debris mound”

Response: Changed as recommended to debris mound throughout the text and also in the legend of Fig. 3.

Referee: Page 20 lines 6-8: I think the authors might want to make this section more clear by saying that methanogenesis in the presence of sulfate isn't expected, but that the methanogenesis could be occurring in the sediments where sulfate is reasonably expected to be depleted. I would then recommend they cite a manuscript or two from the marine sediment pore water research community to demonstrate this depletion is common. The authors' interpretation is also supported by the concentration profile, a point which should be emphasized in the discussion.

Response: We added “...and methanogenesis is not expected in the presence of sulfate, ...” in order to be more clear in this section. Also we followed the referee's advice and cited supporting literature of pore water profiles.

It is not clear for the authors what the referee meant with “The authors' interpretation is also supported by the concentration profile, a point which should be emphasized in the discussion”. We do not think that the concentration profile in the water column reflects the concentration profile in the sediments, if that's what was meant by the referee. Instead, the concentration profile of the species that are released from organic matter decay are a consequence of the 2 sinks of the system, which are the redoxcline above and an advective depletion through a hydraulic connection somewhere below the cenote (<60 m water depth).

Referee: Page 20, line 13: It might be more clear to say that Fe is only present in low concentrations in limestone and there are limited/no sources of siliciclastic materials that could contribute iron in this part of the Yucatan. Fe-oxides are common in many

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karst areas where the Fe is derived from weathering of siliciclastics.

BGD

Response: This part was changed as recommended.

Referee: Page 24, Section 4.3.1 (entire section) – I find the role of Hurricanes in elevating the halocline to be a highly implausible explanation for the vertical zonation of the Hells Bells features. Hurricanes are infrequent and the “piston induced oscillation” effect described by the authors would be short lived. I find it hard to believe that any potential upward displacement by this mechanism would be of sufficient duration to allow for the precipitation of large masses of carbonate photographed above the modern redoxcline. I suggest that a more plausible scenario is that periods of drought (which are well-documented over the Holocene) lowered the water table and increased the elevation of the redoxcline for sustained periods of time (see Hodell, D.A., Brenner, M., Curtis, J.H. and Guilderson, T., 2001. Solar forcing of drought frequency in the Maya lowlands. *Science*, 292(5520), pp.1367-1370.)

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Response: We already referred to this comment in detail above.

Referee: Page 28: Section 4.4 prerequisites for formation of Hells Bells. Many cenotes in the Yucatan meet these criteria yet lack Hells Bells. I wonder if one of the more unique aspects of El Zapote is the relatively thick mixing zone, which the authors highlight in their discussion. Many of the haloclines are much thinner, typically due to high flux of fresh or saline water (though I was surprised by how thin the halocline at Cenote Angelita was in the supporting literature – I wouldn’t have guessed there was much flow in that location). I suppose many of the haloclines with high fluxes of organic material also have much greater light penetration (such as El Pit and Cenote Angelita), so the location of the mixing zone in a dark cenote (I still hesitate to say that this is completely dark) is also likely important.

Response: Yes, we are also convinced that the thickness of the halocline is an important factor as a thick halocline results from highly stagnant conditions. These have to be met in order to create an oxygen deficient redoxcline where anaerobic proton-

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consuming sulfide oxidation can occur. We investigated 3 cenotes with Hells Bells and found that all show a very thick halocline, some even thicker than that at El Zapote, and found that the characteristics of the redoxcline are very different to those of cenotes with a thinner halocline, which matches most of the studied cenotes at the Yucatán Peninsula (MS in preparation).

It seems however, that light intensity is not an important factor because Hells Bells occurrence ranges from close to absolute dark cenotes (i.e. cenote Holbox, which has only one tiny opening and is occupied by bats, that probably provide the organic material) to potentially “bright” cenotes (Cenote Tortugas, which has an opening of approximately the same size as cenote Angelita).

Referee: Page 27, line 19-20: I still wonder if the general orientation of all of the speleothems towards the center of the cenote indicates some phototoxic behavior. It would certainly be the simplest explanation and I would encourage the authors to measure light intensity over an annual cycle to see if perhaps small amounts of light are penetrating (not having visited this particular site, I’m unsure how likely this is, but the Cenote profile included in Fig 3, coupled with the fact that the cenote is sufficiently large in diameter to permit cavern tours, suggests that the entrance may be sufficiently large to allow some small amounts of light during some parts of the day/year).

Response: Yes, we cannot rule out that some phototoxic behavior influences Hells Bells formation because small amounts of light may penetrate the cenote. We agree that this would be the simplest explanation. However, we tried to find an explanation for the observed orientation of the speleothems that is consistent with our results and the hypothesis on the biologically-induced mechanism for calcite precipitation. The chemolithoautotrophy within the redoxcline/turbid layer (which is another barrier for light penetration) is the driving agent for Hells Bells formation, thus we concluded that Hells Bells orientate accordingly to the availability of the “food” of those chemolithoautotrophic organisms. To measure the amount of light over time within the cenote would be a good approach though, in order to clarify this particular problem.

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References: Escolero, O., Marin, L. E., Domínguez-Mariani, E. and Torres-Onofre, S.: Dynamic of the freshwater – saltwater interface in a karstic aquifer under extraordinary recharge action: the Merida Yucatan case study, Environ. Geol., 51, 719–723, doi:10.1007/s00254-006-0383-1, 2007. Heise, L.: Dynamics of the coastal Karst aquifer in northern Yucatán Peninsula. [online] Available from: <http://nинive.uaslp.mx/jspui/handle/i/3692>, 2013. Kovacs, S. E., Reinhardt, E. G., Stastna, M., Coutino, A., Werner, C., Collins, S. V., Devos, F. and Le Maillot, C.: Hurricane Ingrid and Tropical Storm Hanna's effects on the salinity of the coastal aquifer, Quintana Roo, Mexico, J. Hydrol., 551, 703–714, doi:10.1016/j.jhydrol.2017.02.024, 2017. Moore, Y. H., Stoessell, R. K. and Easley, D. H.: Fresh-Water/Sea-Water relationship within a Ground-Water flow system northeastern coast of the yucatan, Ground Water, 30(3), 343–350, 1992. Stoessell, R. K., Moore, Y. H. and Coke, J. G.: The Occurrence and Effect of Sulfate Reduction and Sulfide Oxidation on Coastal Limestone Dissolution in Yucatan Cenotes, Ground Water, 31(4), 566–575, 1993.

Interactive comment on Biogeosciences Discuss., <https://doi.org/10.5194/bg-2018-520>, 2019.

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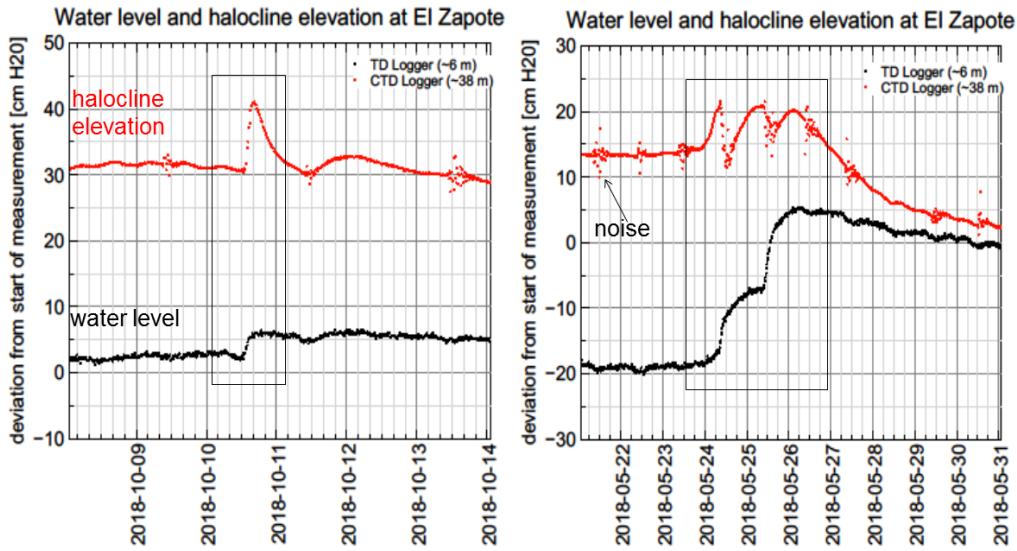


Fig. 1. Preliminary results of data loggers from El Zapote cenote showing two examples of halocline elevation events. (Full caption is given on page C4)

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Author's Response to the Interactive comment on “Subaqueous speleothems (Hells Bells) formed by the interplay of pelagic redoxcline biogeochemistry and specific hydraulic conditions in the El Zapote sinkhole, Yucatán Peninsula, Mexico” by Simon Michael Ritter et al.

The authors are very grateful for the time and effort taken for reviewing the manuscript

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in detail and appreciate the constructive comments and suggestions by anonymous referee #2. We addressed all issues raised by the referee in a point by point response below.

Referee's general comment: This paper presents detailed hydrogeochemical and geochemical analyses of the water column and the so-called Hells Bells formed in a cenote on the Yucatan Peninsula, Mexico, in order to determine the processes leading to the development of the Hells Bells. This is an interesting topic because these submerged speleothems are a unique feature suggesting that their growth is only possible in case of very specific conditions.

The paper is, in general, well written, the results are clearly presented and the developed hypotheses are sound and justified by the data. The text is here and there a bit lengthy (in particular, the conditions leading to the formation of the Hells Bells are repeated several times), but this is not a major issue.

I have two (moderate to major) general comments that should be addressed by the authors prior to publication. The first – and more serious - comment is related to the changes in the depth of the halocline, which are considered as the reason resulting in growth of Hells Bells at different water depths or of large bells over a longer time. The authors develop a hypothesis that the depth of the halocline is related to fresh water recharge at the surface and even suggest a potential relationship with the occurrence of hurricanes. However, the time scale of the growth of the Hells Bells is not discussed with sufficient detail in the paper. Considering the enormous size of at least some of the bells, it is hard to believe that these should have developed due to seasonal or episodic changes in the depth of the halocline. I would rather believe that this requires a longterm shift in the depth of the halocline, for instance over several thousand years during the Holocene. I had a quick look at the previous paper of the same group (Stinnesbeck et al., 2017b), which presented a few U-series data and reported growth rates of ca. 10-100 $\mu\text{m/a}$. In case of such slow growth rates, it is hard to believe that a short-term decrease in the depth of the halocline due to a recharge event



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would have a visible effect. In contrast, growth of a really large bell, requires slow and progressive changes in the (mean) depth of the halocline. The U-series ages ranging from a few hundred to a few thousand years reported by the previous paper, actually seem to confirm this. Thus, the authors should ideally present many more U-series data trying to resolve the timing and duration of the growth of the Hells Bells. If this is technically impossible or beyond the scope of the paper, they could also present data from a second campaign, probably shortly after a major recharge event. As far as I understand, these measurements are currently conducted (p. 26, line 30ff.) and could easily be included in a revised paper. If the authors do not want to include additional data (neither U-series, nor elevation data of the halocline), they must clearly address this issue in the revised paper and critically discuss the time-scales of the dynamics of the halocline and the growth rates of the Hells Bells.

Response: We fully agree that growth rates should be discussed in the MS and followed the referee's suggestion by critically discussing and estimating Hells Bells growth rates. We address this issue in a new chapter in the revised MS (Section: "4.3.1 Calcite precipitation rates"). Concerning the presentation of additional data, we are currently working on the Hells Bells from different angles and produced an amount of data that cannot be shown in one manuscript only. Manuscripts, also addressing age-dating of Hells Bells, are currently in preparation. We conducted a second campaign which confirmed the data presented in the MS, but the major goal of second campaign was to study other cenotes showing Hells Bells and compare them to a cenote without Hells Bells. We are currently preparing a manuscript presenting these data. Additionally, as the dynamic halocline elevation was also of major concern for Anonymous referee #1 and we already addressed a lot of the raised issues and we would therefore also like to refer to the Author's response AC1 (<https://www.biogeosciences-discuss.net/bg-2018-520/#discussion>).

Referee: My second point concerns the typical bell-shape of the Hells Bells. In section 4.3.2, the authors discuss several processes, but none of the them – as far as I

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understood – explains the “conically divergent” (p. 27, line 14) shape of the bells. This issue should either be addressed more clearly, or it should be stated in the MS that the processes discussed in the text cannot completely explain the typical shape of the bells.

Response: Absolutely correct, we did not comment on the conical divergent shape of Hells Bells but of course we thought about that, too. We think that the conical shape of Hells Bells and their downward divergence could be the macroscopic expression of the calcite crystal's microscopic features. We observed two major phases of calcite crystals, blocky or mosaic and elongated doog-dooth calcites. The latter are arranged in a botryoidal structure which strongly resembles the shape of the larger Hells Bells structures. We added this observation into the discussion section 4.3.2 and added a thin section photograph showing such features to the supplement (Fig. S4 in the revised Supplement). Preferential or faster growth on the outer edges of Bells, especially the edges pointing towards the center of the cenotes (this is described in detail in section 4.3.2) may lead to the hollowness because once these outer edges of a specimen have grown slightly deeper than the rest of the specimen, then the net growth of these parts will be higher because it is more frequently reached by the zone of calcite precipitation. This would ultimately lead into the observed hollowness of the larger specimen of Hells Bells. We also integrated these thoughts on the Hells Bells features into the manuscript in section 4.3.2

Referee: In summary, although the paper is generally very interesting and well written, and the data clearly deserve publication, I can only recommend publication in Biogeosciences after revision. Below, I list a few additional, more detailed comments.

Referee's detailed comments:

Referee: Page 6, line 9: Why are only data from a single campaign reported? In particular considering the important aspect of the dynamics of the halocline (see above), it would be much better to provide at least a few data from an additional campaign

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conducted shortly after a major recharge event (hurricane).

Response: We installed data loggers in the cenote for permanent observation of the halocline elevation and water level. This data is so far (~1 year) very promising, however there was no major precipitation event like a since the start of the measurement in December 2017. We elaborated this issue in our response to Referee comment #1 (AC1 on pages C2–C3 <https://www.biogeosciences-discuss.net/bg-2018-520/#discussion>). Furthermore, the main goal of this study is to present detailed hydrogeochemical results in order to develop a hypothesis on the mechanism for subaqueous calcite precipitation. The aspect a halocline elevation is a direct consequence of the conclusion that calcite most likely precipitates within the narrow redoxcline. In the revised MS we are more careful in discussing the halocline dynamics and clearly state that this is not supported by data, yet.

Referee: P. 6, line 20ff: "Due to increasing sulfide concentrations in water depths below the turbid layer and interaction of sulfide with the Ag/Cl pH electrode, a shift of pH of up to 0.2 pH units towards higher values was observed when comparing the pH logs of the way down with the pH logs of the way up (Fig. S1). This shift is dependent on the exposure time of the electrode and the respective sulfide concentrations and could not be quantified nor corrected for." May it be possible to quantify the effect in the laboratory by increasing the sulfide content of a test-solution?

Response: Please see the following response below.

Referee: P. 12, line 19ff.: "However, SI values calculated for the halocline suffer from the overestimated pH readings in the extremely sulfidic water of the halocline and are therefore not considered." This is a pity because supersaturation with respect to calcite within the halocline is the hypothesis presented to explain the growth of the Hells Bells. Thus, it would be really good to estimate the effect on the pH values (see above) and the resulting effect on SI. This information is essential for the validity of the presented hypothesis.

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Response: Unfortunately, quantifying the pH shift is not possible because the sonde that we used was a rental sonde. However, we did this campaign again at El Zapote cenote in 2018 with the exact same type of sonde but a new one. The pH values were identical to the ones presented in the MS except for the values below the redoxcline which were significantly lower at its minimum values in around 42 m water depth (up to ~0.25 pH units). We fully agree that this is essential for the validity of our hypothesis and therefore addressed this issue by including pH and SI calcite data from the sampling campaign in June 2018 into Supplement Fig. S2. We refer to this information in the method section and section “3.1.1 Calcite Saturation”. This allows now to clearly show that calcite is oversaturated in the redoxcline and undersaturated in the halocline indicating calcite dissolution.

Referee: P. 13, line 5: “: : : from X to Y : : :” Is this an artefact from a previous version of the paper?

Response: Yes, thank you. It was corrected in the MS with the according values.

Referee: P. 13, line 5: Fig. 5 should be Fig. 4.

Response: Corrected as suggested.

Referee: P. 14, line 6: Figs 6 should be Figs. 5.

Response: Corrected as suggested.

Referee: P. 16, line 6ff.: Please state here that the samples were collected from “several” specimens. This information is important. In addition, it is (again) problematic that no dating is provided. Then, the data would not be related to the “presumably youngest part” of the bells, but the age of the samples could be precisely determined.

Response: Corrected as suggested. We are currently working on a Manuscript with the focus on age-dating of the Hells Bells.

Referee: P. 16, line 11: “soluble” should be “insoluble”?

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Response: Yes, of course. Thank you. Corrected as suggested.

Referee: P. 16, line 21: d13Ccalcite values show a strong negative (not a positive) correlation with Sr/Ca and Ba/Ca if I correctly read Fig. 6.

Response: You are right. Corrected as suggested.

Referee: P. 16, line 25: In my opinion, the offset between the calculated and the measured d13C value of the HCO₃ (2-5 permille) is substantial. Thus, speaking of "slightly lower" values is not correct.

Response: We agree and deleted "slightly", now saying that the s13C value of the measured HCO₃ is lower than the calculated d13Ce_q HCO₃ vlaue.

Referee: Section 4: The introductory section could be deleted to make the paper more concise.

Response: We followed the referee's suggestions and deleted this section.

Referee: Section 4.2: See above. The introductory section could be deleted to make the paper more concise.

Response: We followed the referee's suggestions and deleted this section.

Referee: P. 22, line 7: Please provide a reference for reaction (R1).

Response: Reaction R1 was developed by the authors. However, in order to meet the referee's comment we referenced to supporting literature that shows the process of anaerobic sulfide oxidation via nitrate ("...could be obtained with nitrate as terminal electron acceptor producing elemental sulfur and nitrogen under the consumption of protons (e.g. Bailey et al., 2009).")

Referee: P. 22, line 21: Please define "SD-OM".

Response: SD-OM is already defined in the MS above ("...could be obtained with nitrate as terminal electron acceptor producing elemental sulfur and nitrogen under the

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consumption of protons (e.g. Bailey et al., 2009)."

Referee: P. 22, line 27: "As organisms usually prefer to metabolize ^{12}C (it takes less energy to break the ^{12}C bond instead of ^{13}C) they effectively consume HCO_3^- with lower $\delta^{13}\text{C}$ values, which subsequently results in higher $\delta^{13}\text{CHCO}_3^-$ values in the remaining dissolved inorganic carbon" It is true that the organisms preferentially metabolise $^{12}\text{CO}_2$, but they do not "effectively consume HCO_3^- with lower $\delta^{13}\text{C}$ values". The preferential consume of $^{12}\text{CO}_2$ (and the related increase in the $\delta^{13}\text{C}$ value of the CO_2) leads to chemical and isotopic reactions resulting in conversion of HCO_3^- into CO_2 and an increase in the $\delta^{13}\text{C}$ value of the HCO_3^- reservoir.

Response: We agree and deleted "effectively consume HCO_3^- with lower $\delta^{13}\text{C}$ values" . We changed the according sentence to "Organisms usually prefer to metabolize ^{12}C (it takes less energy to break the ^{12}C bond instead of ^{13}C), which results in higher $\delta^{13}\text{C-HCO}_3^-$ values in the remaining dissolved inorganic carbon."

Referee: P. 23, line 2: "It was shown before that Hells Bells form within the freshwater indicated by $\text{d}^{234}\text{U}_{\text{initial}}$ values of 16–25 ‰ of the Hells Bells calcite (Stinnesbeck et al., 2017b)." Please explain this statement. Why do these values suggest precipitation within freshwater? Due to the non-marine d^{234}U value (lower than 150 permille)? Or has the d^{234}U value of the water in the cenote been determined (at different depths)? Actually, freshwater often has higher d^{234}U values than seawater

Response: We agree that we cannot state this without showing the d^{234}U values over the depth range of the water body. We removed the argument of d^{234}U and changed the sentence to "It was suspected before that Hells Bells form within the freshwater body of El Zapote cenote (Stinnesbeck et al., 2017b)."

Referee: P. 24, line 8: "Hells Bells formed in modern to historic times : : :" How do you know that? Is this statement based on the few U-series ages reported by Stinnesbeck et al. (2017b)? It may very well be possible that there are much older specimens in the same cenote.

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Interactive
comment

Response: This statement is based on the U-series measurements and the 14C age of the tree stem which is covered by small specimen of Hells Bells. We agree that there should be much older specimen considering their enormous size. We changed the sentence to "Hells Bells formed in modern to at least historic times and occur..." by inserting "to at least historic times" in order to meet the referee's critique.

Referee: P. 24, line 13ff.: "Therefore, we propose that growth of Hells Bells is a non-permanent episodic process which majorly depends on a highly variable halocline elevation in the cenote (Fig 10)." See my major comment above. The probably very different growth rate of the bells and the seasonal to episodic dynamics of the depth of the halocline need to be discussed in detail.

Response: We fully agree and added a detailed discussion on growth rates to the revised MS. However, although we are convinced that the halocline dynamic is an interesting and upcoming research topic we believe that a detailed discussion on the dynamic elevation of the depth of the halocline is beyond the scope of the MS.

Referee: P. 24, line 16ff.: "Extraordinary recharge events (e.g. hurricanes) must have a significant effect on the depth position of this layer : : :" See above. Even if this is the case, it is not clear whether these episodic changes would be recorded by the slowly growing Hells Bells.

Response: The detailed discussion on growth rates in the revised MS demonstrates that short-term episodic events could be recorded in Hells Bells growth.

Interactive comment on Biogeosciences Discuss., <https://doi.org/10.5194/bg-2018-520>, 2019.

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List of relevant changes to the manuscript

Page numbers refer to the revised manuscript (MS).

1. Pages 23–25: A new section “4.3.1 Calcite precipitation rates” was included in the revised MS where calcite precipitation rates of the proposed biogeochemical mechanisms are estimated and discussed. For this purpose, an additional Figure (Fig. S4) showing the upward flux of DIC was added to the Supplement. Also, Table S5 was added to the Supplement comprising the results and used parameters of the calculations.

2. Page 18: Introductory section of section “4. Discussion” was deleted.

- 10 3. Page 19: Introductory section to section was deleted (4.2. Hydrogeochemical processes in El Zapote)

4. Pages 25-27: Section “4.3.2 The role of halocline elevation in Hells Bells formation”

Droughts were included as a potential mechanism for an elevated the halocline. This was also included in the Abstract (Page 1) and section “5. Conclusion” on Page 29. The time scales of both episodic halocline elevations and calcite precipitation rates or Hells Bells growth are now discussed in this section.

- 15 5. Page 28: In section “4.3.3 Shape of Hells Bells” a paragraph was added discussing the hollow interior of Hells Bells. Additionally, to support this discussion Fig. S5 was added to the Supplement, which shows a thin section of Hells Bells.

20

6. Pages 6 and 12: Figure S2 in the Supplement was changed by including pH and SI calcite data from a second sampling campaign in order to show calcite dissolving conditions in the halocline. We refer to Fig. S2 in the method section on page 6 and in the result section “3.1.1 Calcite Saturation” on page 12.

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7. Figure 3c was modified, the horizontal axis titles are now in the same color (grey and Black as the respective data points).

8. Figure 8: The order of magnitude of the flux in the caption of Fig. 8 was wrong. It was changed from $10^{-8} \mu\text{mol m}^{-2} \text{s}^{-1}$ to $10^{-5} \mu\text{mol m}^{-2} \text{s}^{-1}$.

30

Furthermore, the MS was subject to several minor changes mostly following the referee’s suggestions including spelling and wording. These changes are referred to in detail in the point-by-point responses.

Subaqueous speleothems (Hells Bells) formed by the interplay of pelagic redoxcline biogeochemistry and specific hydraulic conditions in the El Zapote sinkhole, Yucatán Peninsula, Mexico

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Abstract

Unique bell-shaped underwater speleothems were recently reported from the deep (~55 m) meromictic El Zapote sinkhole (cenote) on the Yucatán Peninsula, Mexico. The local diving community has termed these speleothems as Hells Bells

because of their shape and appearance in a lightlessdark environment in ~28–38 m water depth above a sulfidic halocline. It was also suggested that Hells Bells form under water, yet the mystery of their formation remained unresolved. Therefore, we conducted detailed hydrogeochemical and geochemical analyses of the water column and Hells Bells speleothems including stable carbon isotopes. Based on the comprehensive results presented in this study we deduce that both, biogeochemical processes in the pelagic redoxcline and a dynamic halocline elevation of El Zapote cenote, are essential for Hells Bells formation.

Hells Bells most likely form in the redoxcline, a narrow 1–2 m thick water layer immediately above the halocline where a pelagic chemolithoautotrophic microbial community thrives from the upward diffusion of reduced carbon, nitrogen and sulfur species released from organic matter degradation in organic-rich debris. We hypothesize that chemolithoautotrophy, in particular the proton consuming nitrate-driven anaerobic sulfide oxidation, favors calcite precipitation in the redoxcline and hence Hells Bells formation. A dynamic elevation of the halocline as a hydraulic response

to droughts and recharge events, e.g. hurricanes, is further discussed, which might explain the shape of Hells Bells as well as their occurrence over a range of 10 m water depth. Finally, we infer that highly stagnant conditions, a low light environment and sufficient input of organic material into a deep meromictic cenote are apparent prerequisites for Hells Bells formation

~~considering the. This might explain their~~ exclusivity ~~of these underwater speleothems~~ to only a few cenotes ~~of~~in a restricted area of the northeastern Yucatán Peninsula.

1. Introduction

~~Pending speleothems, also called Speleothems, such as~~ stalactites or dripstones, result from physicochemical processes under subaerial conditions in a cave atmosphere. Calcite precipitates due to CO₂-degassing and evaporation of water enriched in dissolved carbonate dripping into the cave. Normally, stalactites rejuvenate and form a tip at the lower end from which the water drips to the cave floor. Nevertheless, in recent years, researchers have identified a small group of stalactites that appear to have calcified underwater. For these formations, interactions between physicochemical and biological calcite precipitation processes are interpreted (~~Barton and Northup, 2007; Bontognali et al., 2016; Gradzinski et al., 2012; Guido et al., 2013; Holmes et al., 2001; Jones et al., 2008, 2012; Macalady et al., 2007; Macintyre, 1984; Melim et al., 2001; Queen and Melim, 2006~~).~~(Barton and Northup, 2007; Bontognali et al., 2016; Gradzinski et al., 2012; Guido et al., 2013; Holmes et al., 2001; Jones et al., 2008, 2012; Macalady et al., 2007; Macintyre, 1984; Melim et al., 2001; Queen and Melim, 2006; Tredici et al., 2018)~~. We recently presented a spectacular example for these subaqueous speleothems termed Hells Bells from El Zapote sinkhole about 26 km west of Puerto Morelos on the Yucatán Peninsula of southern Mexico (Fig. 1) (Stinnesbeck et al., 2017b).

These bell-shaped structures consist of calcite and reach lengths of up to two meter. Hells Bells conically expand downward with strictly horizontal ring-like concentric swellings and neckings on the surface (Fig. 2). Apparently, they form in a lightless environment in freshwater above the anoxic and sulfidic halocline (Stinnesbeck et al., 2017b). Because of these toxic environmental conditions in complete darkness, the local diving community has termed the El Zapote speleothem formations as Hells Bells. They grow from the cavern ceiling and wall. Additionally, small individuals also cover a tree that has fallen into the sinkhole around ~3.5 cal kyr BP, which indicates that Hells Bells must have formed during the Holocene to at least historical times, and thus at periods when the deep sections of the cenote had already been submerged for thousands of years (Stinnesbeck et al., 2017b). Thus, the conditions for the formation of the biggest underwater speleothems worldwide must have existed consistently throughout the past thousands of years at El Zapote cenote.

The internal structure of Hells Bells is characterized by laminar fabrics of alternating units of elongated dogtooth spar calcite and microcrystalline spar calcite. Microspar layers and corroded lobes of dogtooth spar crystals indicate either discontinuous growth of Hells Bells and/or intermittent dissolution. Phylogenetic analyses of Hells Bells speleothem surfaces from specimens of 33 and 34 m water depth indicate that microorganisms inhabiting the Hells Bells potentially support a full nitrogen-circle and autotrophic growth (Stinnesbeck et al., 2017b). The growth of Hells Bells may thus be mediated by specific physical and biogeochemical conditions above and in the halocline, while formation of Hells Bells was likely restricted to the lowermost part of the freshwater body. However, due to the limited available data including geochemical parameters, the suggested processes for Hells Bells formation were regarded as highly speculative.

Therefore, in this study we conducted detailed geochemical analysis including stable carbon isotopes of the water body and Hells Bells speleothems of El Zapote cenote. Based on the results we present a hypothesis on the subaqueous growth of these exceptional structures. We deduce that both, biogeochemical processes in the pelagic redoxcline and a dynamic halocline elevation of El Zapote cenote, are essential for Hells Bells formation.

5 1.1 Geological background

The northeastern Yucatán Peninsula consists of horizontally layered shallow-water carbonates of Mio-, Plio- and Pleistocene ages (Lefticariu et al., 2006; Weidie, 1985) and probably hosts the largest network of underwater caves in the world. The Mexican state of Quintana Roo alone counts more than 370 underwater caves, with a total estimated length of >7000 km and a confirmed total length of ~1460 km and individual cave systems reaching up to >350 km in length (QRSS, 2018). These

10 karst cave systems developed predominantly by the interaction of glacioeustacy, littoral processes and mixing-zone hydrology during glacial periods of the Pleistocene (Smart et al., 2006; Weidie, 1985). Precipitation rapidly infiltrates through the porous limestone into the underlying coastal aquifer consisting of a meteoric water mass, the fresh water lens above a marine water mass intruding from the coast (e.g. Kovacs et al., 2017a). The thickness of the freshwater lens varies between ~10–100 m and is generally thinner towards the coast (Beddows et al., 2002), while seawater intrudes up to 100 km

15 inland (Beddows et al., 2007). The halocline separates the meteoric and marine water bodies and is usually characterized by undersaturation with respect to CaCO₃, leading to cave formation and conduit enlargement in the coastal carbonate aquifer (Back et al., 1986; Gulley et al., 2016; Mylroie and Carew, 1990; Smart et al., 2006). The depth of the halocline increases with distance from the coast (Bauer-Gottwein et al., 2011); areas closer to the coast show a higher salinity of the freshwater lens than inland areas (Kovacs et al., 2017b). The position of the halocline is also dependent on global sea level and the

20 thickness of the freshwater lens. Hydraulic gradients are generally very low with values of 1–10 cm km⁻¹ (Bauer-Gottwein et al., 2011 and references therein). Although Moore et al. (1992) and Stoessell et al. (1993) report that the thickness of the freshwater lens does not vary significantly between seasons or on a yearly basis, local and short-termed variations are possible and were reported by Escolero et al. (2007), who documented a significant halocline elevation of up to 17.5 m in response to great recharge events, such as hurricanes, between two measurements in the years 2000 and 2003.

25 Sinkholes (locally called *cenotes*) were formed by dissolution and collapse of the carbonate rock. They are common throughout the Yucatán Peninsula, connecting the subterranean cave system with the surface (Bauer-Gottwein et al., 2011). For more detailed information about the formation and occurrence of cenotes on the Yucatán Peninsula we refer the readers to Torres-Talamente et al. (2011) and Schmitter-Soto et al. (2002).

1.2 El Zapote cenote

30 El Zapote cenote is located 26 km west of Puerto Morelos on the Yucatán Peninsula of southeast Mexico (20°51'27.78" N 87°07'35.93" W; Fig. 1). In cross-section the cenote is bottle-shaped with a deep vertical water-filled shaft that opens at 28 m water depth to a wide cavern of 60 to >100 m in diameter that reaches to about 54 m water depth with a 20 m high debris

mountmound in the center (Fig. 3-a3a). A fallen tree stands on top of the debris mountmound and small Hells Bells cover the stem. There are no apparent passages or conduits that connect El Zapote cenote to a cave system. Additional details on El Zapote cenote are given in Stinnesbeck et al. (2017b) and in Stinnesbeck et al. (2017a) who described the new genus and species of a giant ground sloth, *Xibalbaonyx oviceps*, from an individual that was found on the floor of El Zapote cenote.

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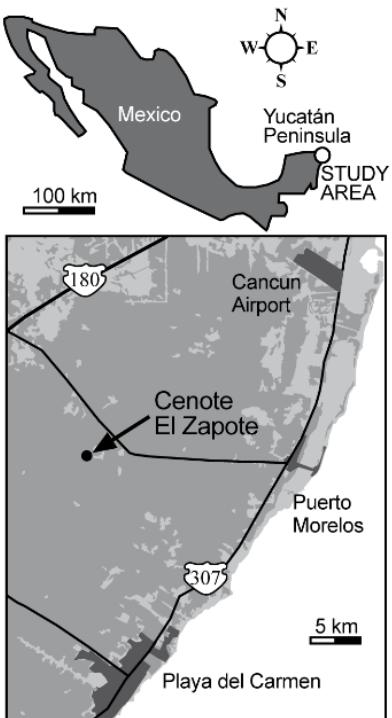


Fig. 1: Location of the studied El Zapote cenote with respect to Mexico and the Yucatán Peninsula.

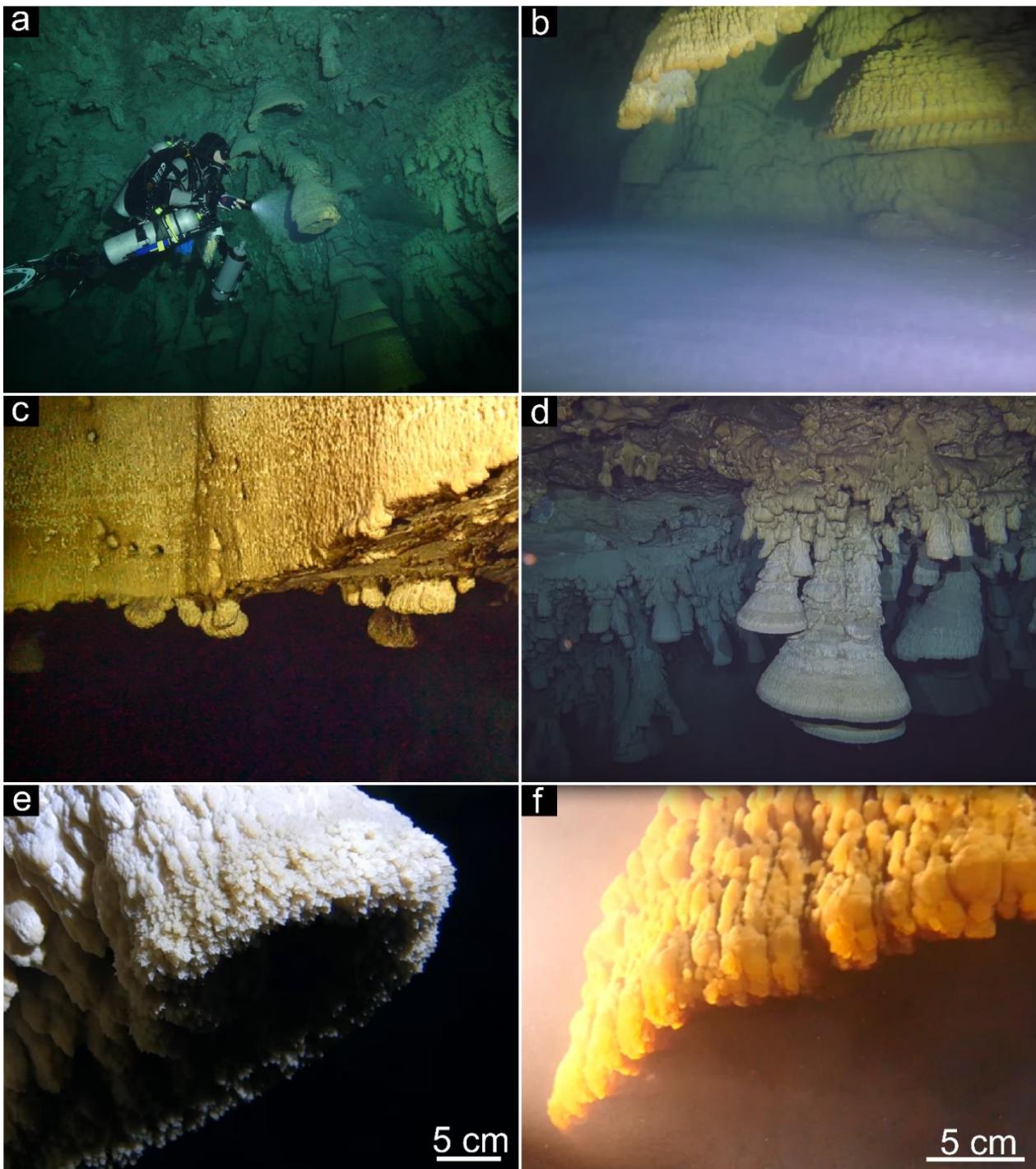


Fig. 2: Technical diver in El Zapote cenote during a sample collecting dive carrying the Niskin bottle, sample containers and the multiparameter sonde attached to the sidemount gas bottle (a). Turbid layer immediately above the halocline forming a distinct horizontal white cloud at around 36 m water depth (b). Transition of cenote shaft to the open dome-shaped cavern at 28 m water depth (c), where the vertical wall of the cenote shaft is devoid of Hells Bells speleothems (upper part of picture c), whereas small

specimens of Hells Bells grow down from the horizontal ceiling below (lower part of c). Brown-colored Mn-oxide coatings on host rock carbonates and Hells Bells speleothems reach down from 28 m water depth to around 30 m water depth at the transition of the cenote shaft to the wide dome-like cavern (c and d). Below around 30 m water depth, Hells Bells speleothem and host rock carbonate surfaces are devoid of brownish Mn-oxide coatings. They are white to light-grey colored revealing a distinct horizontal boundary (d). Close-up shot of the lowermost calcite rim of a Hells Bell speleothem at around 32–35 m water depth showing mm-sized calcite crystals (e).

2. Methods

2.1 Sampling

Sampling at El Zapote Cenote was carried out between December 10th and 15th 2017. Water samples were taken early in the morning prior to any touristic diving group activities to ensure sampling of an undisturbed water column. Water sample recovery and the recording of the in-situ parameters were carried out with a winch from the surface down to the top of the debris ~~mountmound~~ (0–36 m) and by technical divers from the top of the debris ~~mountmound~~ down to the cenote floor following the slope of the debris ~~mountmound~~ (Fig. 2-a2a).

In-situ parameters pH (± 0.1), eH (± 20 mV), dissolved oxygen (± 0.1 mg l⁻¹, detection limit 0.2 mg/l), electrical conductivity (± 0.05 % of value), temperature (± 0.01 °C) and turbidity (± 2 % of value) were determined with a multiparameter water sonde EXO-1 (Xylem Analytics, Norway). All parameters, including water depth via pressure measurement, were internally logged by the sonde (± 0.04 m). Water depths were corrected to the ambient air pressure of the respective day of sampling. In order to reach the greatest possible water depth, a total of four winch-operated profiles were run within 2 days, with laterally shifting starting points of the profile at the surface. In order to complete the measurement in the whole water column, technical divers carried the EXO-1 Probe with them during sampling. Due to increasing sulfide concentrations in water depths below the turbid layer and interaction of sulfide with the Ag/Cl pH electrode, a shift of pH of up to 0.2 pH units towards higher values was observed when comparing the pH logs of the way down with the pH logs of the way up (Fig. S1). This shift is dependent on the exposure time of the electrode and the respective sulfide concentrations and could not be quantified nor corrected for. However, the sensor recovers to initial pH values after a certain time in non-sulfide water. Therefore, the pH values presented in this study are representative for the water column from 0 to 37 m water depth and are overestimated in water depths from 37–50 m where the actual absolute pH values are most likely lower, i.e. more acidic.

Repeated measurement with a new sonde of the same type in June 2018 confirmed this assumption and showed lower pH values below 37 m water depth (Fig. S2).

Water samples from 15–35 m water depth were retrieved using a winch and a 5 L polyethylene Niskin bottle (Hydrobios, Kiel, Germany). Sampling depths represent the center of the 0.6 m tall sampling bottle and were determined by cable length with a depth counter attached to the winch. Water samples from 35.2–45 m water depth were retrieved by technical divers (Fig. 2-a2a). Water samples collected by the technical divers were taken with 120 ml PE-containers and the water depth was noted for each sample. Water samples for the analysis of dissolved gases (CO₂, CH₄) were taken in 24 ml glass vials and sealed underwater at the respective depth (four samples at each depth level). The EXO-1 sonde was attached to a side

mounted compressed air bottle pointing towards the front of the technical diver in order to record the in-situ parameters of each water sample (Fig. 2-a2a). The depth of the water samples taken by technical divers was corrected to the depth of the attached logging device (EXO-1). For four samples between 35 and 37 m, depth was interpreted from the increase of sodium and chloride contents correlated to the electrical conductivity in this interval.

- 5 Water samples were treated on-site immediately after the water samples were retrieved. Samples for determination of dissolved ions were taken with 20 ml sterile polypropylene syringes and then filtered through a cellulose acetate filter (0.45 µm). Samples for cation determination were acidified with 50 µl of 65 % HNO₃ (A.G.) to adjust a pH <2; they were stored in 15 ml Falcon polypropylene centrifuge tubes. Samples for anion determination were taken equivalent, but not acidified, and stored cool in 15 ml Falcon polypropylene centrifuge tubes. Samples for the determination of dissolved inorganic carbon 10 (DIC) and dissolved organic carbon (DOC) were filtered through a cellulose acetate filter (0.45 µm), stored in 24 ml glass vials and sealed gas-tight. Samples for the determination of content and isotopic ratios of the dissolved gases CH₄ and CO₂ were filled in 24 ml glass vials; subsequently 100 µl 60 % HgCl₂ solution was added via a syringe pierced through the septum to sterilize the samples.

A large volume sample (5 l) of the turbid layer water at around 36 m water depth was taken with a Niskin bottle by technical 15 divers and subsequently filtered through a 0.45 µm cellulose acetate filter with a vacuum pump. The filter was air-dried; back in the laboratory a small piece of the filter was coated with carbon for subsequent secondary electron (SE) imaging and analyses.

Sulfide and nitrite were determined on-site by photometric analysis with a Photometer (Hach Lange DR200) using Merck Spectroquant spectrometric methods
20 Technical divers collected several Hells Bells samples grown on the tree trunk from 7 water depth levels between 32.7 and 37.3 m. To obtain the youngest part of individual Hells Bells growth, samples were first studied under the microscope. Only samples with apparently fresh, well accentuated crystal tips were chosen for analysis geochemical and stable isotope analysis.

2.2 Analytical measurements

25 2.2.1 Major and trace element analysis

Major cation concentrations (Ca, Mg, Sr, Ba, K, Na, Si) of water samples were determined by optical emission spectroscopy with an Agilent 720 ICP-OES. Quality control was performed using the reference materials SPS-SW1, SPS-SW2 and TMDA 70.2. Recovery rates were in the range of 97–104 % for the analyzed elements. Measurement precision for each element was <2 % (RSD, n=3).

30 Concentrations of anions Cl⁻, SO₄²⁻ and NO₃⁻ were determined with ion-chromatography (Dionex ICS-1100) with a RSD of <3 % derived from long-term repeated analysis of reference material SPS-WW1 NUTR. The concentrations of DOC were

determined with a Total Carbon Analyzer (Shimadzu TOC-CPH) with a RSD of <2 % derived from repeated analysis of an in-house standard water.

Orthophosphate was determined by the photometric molybdenum phenyl blue method on 880 nm light extinction with a UV/VIS photometer (Specord 50, Analytic Jena).

- 5 Trace element concentrations (Fe, Mn, Mo, P, S, U) of selected water samples were determined by HR-ICP-MS (Thermo Finnigan Element 2). Analyses were normalized by an internal Indium-standard. Calibration solutions were prepared with the MERCK VI Multi Element Standard. The recovery rates of SLRS5 reference material were 95 % (Fe), 93 % (Mn), 82 % (Mo) and 133 % (U) with respect to the referenced values and P and S were within the range of reported uncertified values. The precision was <3.7 % (RSD) derived from repeated (n=5) measurements of the reference material in the measurement
10 run.

Around 3 mg of the powdered speleothem samples were digested in 2ml 10% HNO₃ for major and trace cation analyses. Subsequently, concentrations of Ca, Mg, Sr, Ba, P, S, Fe and Mn of the aliquots were determined by ICP-OES. Quality control of the measurement was performed using reference materials SPS-SW1 and SPS-SW2 with recovery rates ranging from 99 to 111 % for the analyzed elements. Quality control for digestion of the carbonate material was performed with
15 limestone reference material ECRM 752-1 with recovery rates between 106–110 % for the elements Ca, Mg, Ba, Sr and Mn and 82 % for the element Fe.

Calcite saturation and HS⁻ activity was calculated with PhreeqC (Parkhurst and Appelo, 1999) using phreeqc.dat. The diffusion J was calculated with the first Fick's law with diffusion coefficients of D_{O₂}, D_{NO₃}, D_{HCO₃} and D_{HS⁻} of 2.1, 1.9, 1.2 and 1.28 10⁻⁹ m² s⁻¹, respectively taken form phreeqc.dat (Parkhurst and Appelo, 1999).

20 2.2.2 Stable carbon isotope and concentration measurements of CH₄ and CO₂

For the determination of dissolved gases, a 5 ml headspace with nitrogen gas (N₂ 99.999 %) was created in each of the four samples of the respective water depth. Samples were taken for the analysis of dissolved gases at ambient laboratory temperatures of 23 °C. After equilibration (~24h), the headspace of the four samples was transferred and combined in one 12 ml evacuated exetainer vial. To ensure a pressureless transfer of the gas phase from the headspace to the exetainer, a brine
25 solution of 200 g l⁻¹ NaCl was introduced at the bottom of the vial and the gas-phase was simultaneously removed and subsequently transferred to evacuated exetainer vials. Concentrations CH₄ and CO₂ in the gas samples were measured as follows: Headspace samples (50µl) were injected in a flow of 1mL min⁻¹ of helium with a split ratio of 5:1 to a Shin Carbon ST column (80/100 mesh, 2m x 0.53mm i.d., Restek Corporation) quantified by a gas chromatograph (GC-2010 Plus, Shimadzu Corporation, Kyoto, Japan) coupled to a Barrier Ionization Discharge (BID) detector (BID-2010 Plus, Shimadzu Corporation, Kyoto, Japan). The GC oven was initially held at 30°C for 1 min and then ramped at 10°C/min to 200°C. Quantification of CH₄ and CO₂ was carried out by comparison of the integrals of the peaks eluting at the same retention time as that of the authentic standard with calibration curves. The dissolved concentrations of CH₄ in the water were then
30

calculated from the measured mixing ratio using Henry's law (Wiesenburg and Guinasso, 1979) and solubility coefficients for CH₄ according to Weiss (1974) and Yamamoto et al. (1976).

Stable carbon isotope ratios of CO₂ ($\delta^{13}\text{C}$ -CO₂ values) were analyzed by gas chromatography stable isotope ratio mass spectrometry (GC-IRMS) by a HP 6890N gas chromatograph, coupled to a 253 Plus™ isotope ratio mass spectrometer 5 (ThermoQuest Finnigan, Bremen, Germany) with average analytical uncertainties of 0.2‰ for $\delta^{13}\text{C}$ -CO₂ values. 2 σ uncertainties were derived from 5 replicates. All $^{13}\text{C}/^{12}\text{C}$ isotope ratios are expressed in the conventional δ notation in per mil versus VPDB, defined in [equation \(Eq. 1\)](#):

$$\delta^{13}\text{C}_{V-PDB} = [(^{13}\text{C}/^{12}\text{C}_{sample})/(^{13}\text{C}/^{12}\text{C}_{standard})] - 1 \quad (\text{Eq. 1})$$

For details of the $\delta^{13}\text{C}$ -CO₂ measurements by GC-IRMS we would like to refer to previous studies by Keppler et al. (2010) 10 and Laukenmann et al. (2010).

Stable carbon isotope ratios of CH₄ ($\delta^{13}\text{C}$ -CH₄ values) were determined (GC-IRMS). In brief, CH₄ of the sample was trapped on Hayesep D and then transferred to the IRMS system (ThermoFinnigan Delta^{plus} XL, Thermo Finnigan, Bremen, Germany). The working reference gas was carbon dioxide of high purity (carbon dioxide 4.5, Messer Griesheim, Frankfurt, Germany) with a known $\delta^{13}\text{C}$ -CH₄ value of -23.634 ‰ ± 0.006 ‰ versus V-PDB. All $\delta^{13}\text{C}$ -CH₄ values were corrected using 15 two CH₄ working standards (isometric instruments, Victoria, Canada) and normalized by two-scale anchor calibration according to Paul et al. (2007). The average standard deviation of the analytical measurements was in the range of 0.1 ‰ to 0.3 ‰.

The $\delta^{13}\text{C}$ -HCO₃⁻ values were calculated from the measured $\delta^{13}\text{C}$ -CO₂ of the headspace of the water samples that was generated in the laboratory as equilibrium fractionation at 23° C ($\delta^{13}\text{C}$ -CO₂ + 8.16 ‰ = $\delta^{13}\text{C}$ - HCO₃⁻) after Mook (2000).

20 For stable carbon isotope analyses of carbonates, approximately 50 µg of powdered speleothem subsamples was analyzed using a ThermoFinnigan MAT253Plus gas source mass spectrometer equipped with a Thermo Fisher Scientific Kiel IV Carbonate Device at Heidelberg University. Values are reported relative to VPDB (Eq. 1) through the analysis of an in-house standard (Solnhofen limestone) calibrated to IAEA-603. The precision of the $\delta^{13}\text{C}$ analyses is better than 0.08 ‰ and 0.06 ‰ (at 1σ level), respectively.

25 2.2.3 Optical methods

Hells Bell specimen ZPT 7, described in Stinnesbeck et al. (2017b) was vertically cut in half and thin sections were prepared from one half of the specimen. Photographs of the thin sections were taken with a Keyence VHX-6000.

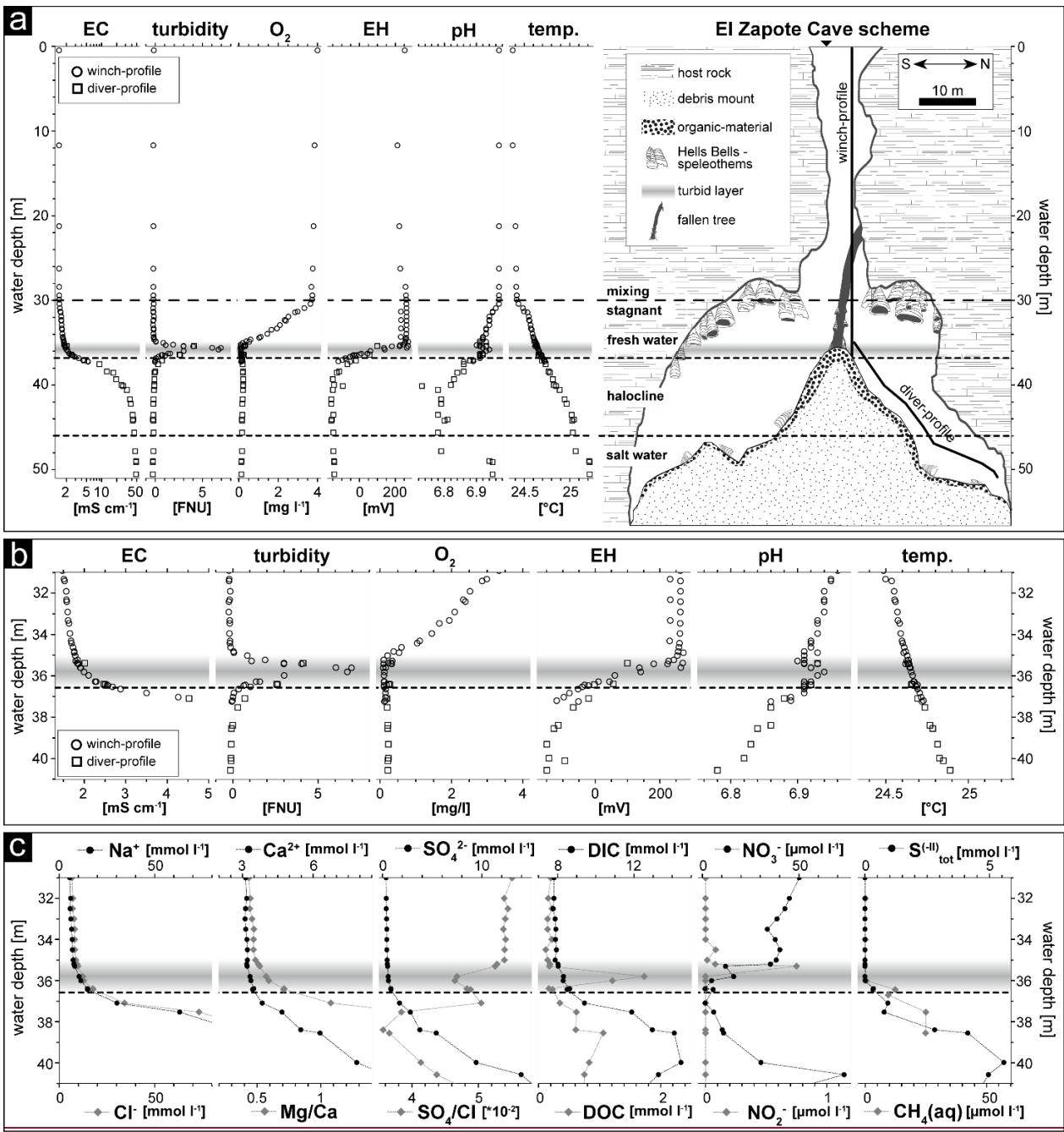
Polished counterparts of the thin sections and small pieces of Hells Bells were coated with carbon for secondary electron (SE) imaging and energy dispersive X-ray (EDX) analyses. SE-imaging and element mapping was performed with a Leo 440 30 at 20 kV with a X-Max 80 mm² detector.

3. Results

3.1 Hydrogeochemistry

The water column of the El Zapote cenote is stratified into an ~~oxygenized~~oxygenated fresh water body overlying an anoxic transition zone of increasing electrical conductivity (EC), the halocline, and an anoxic salt water body below (Fig. 3 a).

- 5 Water temperatures vary little between 0 to 30 m water depth (24.37–24.42 °C); a steep increase is identified in a narrow zone from 30–32 m water depth (24.42–24.55 °C), followed by almost invariable temperatures from 32 m water depth (24.55 °C) down to the bottom of the cenote (25.22°C) (Figs. 3-a3a and b). A distinct density boundary, the top of the halocline, is identified at 36.6 m water depth by a steep increase in EC. Sea water-like salinity is reached at around 46 m water depth indicating a thick halocline layer of around 10 m thickness (Figs. 3-a3a and b). Low turbidity readings indicate clear water 10 throughout the water column, except for a ~1.6 m thick layer of increased turbidity immediately above the halocline from 35.0–36.6 m water depth, with a peak of 8.0 FNU detected at 35.8 m water depth (Figs. 3-a3a and b; Table S1). This turbid layer is also easily detected macroscopically in the water column as a white cloudy layer (Fig. 2-b2b) and coincides with a distinct redoxcline from ~35–37 m water depth, in which the redox potential (EH) decreases from ~250 to ~140 mV (Figs. 3-a3a and b). Dissolved oxygen (DO) decreases nearly linear from 30 m to concentrations below detection limit at ~35 m water 15 depth just above the turbid layer. Below, DO was below detection limit (Figs. 3-a3a and b). The pH shows neutral values from 0–30 m water depth and slightly decreases to 6.90 at the top of the turbid layer (Fig. 3-a3a). Within the turbid layer pH values increase to more alkaline values of around 6.94 at 35.8 m water depth. The pH values decrease again below the turbid layer to 6.73 at 40 m and invariably remain at about this value down to 48 m. From there, values increase to about neutral (6.95) close to the cenote bottom at 49 m water depth (Figs. 3-a3a and b).
- 20 Concentrations of the major dissolved ions Na^+ , Cl^- , Ca^{2+} , Mg^{2+} and SO_4^{2-} reflect the stratification of the water column in the cenote, with generally low concentrations in the fresh water body from 0–30 m water depth and slightly increasing concentrations from 30 m water depth to the top of the turbid layer at 35 m water depth, a stronger increase within the turbid layer from 35–36.6 m water depth, and an even stronger increase from the top of the halocline at 36.6 m water depth down to the cenote bottom (Fig. 3-e3c and Table S2). Mg/Ca ratios strongly increase from the top of the turbid layer at 35 m water 25 downwards, due to higher Mg concentrations in the salt water body (Fig. 3-e3c). Although sulfate concentrations increase downwards from the top of the halocline, a relative decrease of SO_4^{2-} ions is detected, compared to the chemically conservative ion Cl^- , by a decrease in $\text{SO}_4^{2-}/\text{Cl}^-$ within the turbid layer and below in the halocline (Fig. 3-e3c).



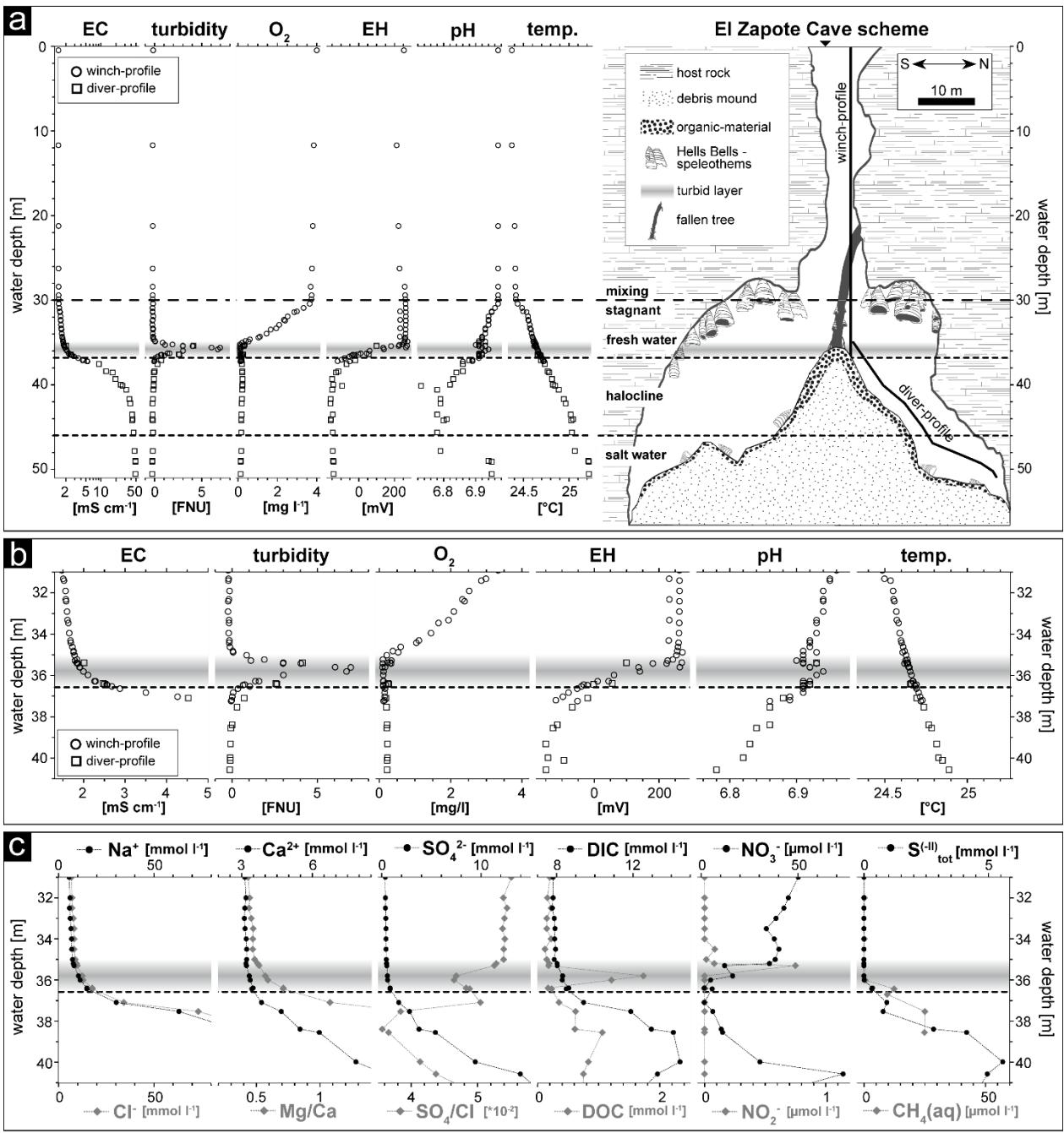


Fig. 3: Hydrogeochemistry of the water column of El Zapote cenote. The horizontal grey band indicates the depth position of the turbid layer, while the dashed line indicates the top of the halocline at 36.6 m water depth. a: Water in-situ parameters versus water depth (left) in relation to the El Zapote cenote cross section (right). In-situ parameters and samples were taken along a winch profile and a diver profile as shown in the cenote cross section. Note the logarithmic scale of the electrical conductivity (EC). b: Close-up of the water in-situ parameters in 31–41 m water depth. Note that the scale of EC is non-logarithmic and is only shown for the range between 1–5 $mS\ cm^{-1}$, in order to point out the increase in salinity at the beginning of the halocline. c: Water

chemical parameters determined in the water column between 31–41 m water depth. Na^+ and Cl^- concentrations are only shown in the range of 0–80 mmol l⁻¹ to highlight the concentration pattern above and within the halocline.

Concentrations of DIC are about 7.8 mmol l⁻¹ in the fresh water body. They increase in the turbid layer and show a peak at 40 m water depth with concentrations increasing to 14.5 mmol l⁻¹; below, they decrease towards the cenote bottom (Fig. 3

5 [e3c](#) and Table S2). The dissolved organic carbon (DOC) concentrations are low in the fresh water body and show a distinct peak within the turbid layer, coinciding with the peak in turbidity at 35.7 m water depth (Fig. [3-e3c](#)). Below the turbid layer DOC concentrations slightly increase and peak at 39–40 m water depth, decreasing from there towards the cenote bottom (Fig. [3-e3c](#) and Table S2). Nitrate concentrations are ~50 $\mu\text{mol l}^{-1}$ in the fresh water unit of the cenote shaft (Table S2). They decrease from 30 m water depth towards the top of the turbid layer and rapidly fall below detectable concentrations within 10 this layer (Fig. [3-e3c](#)). Nitrite peaks in a narrow zone at the top of the turbid layer with concentrations of up to 0.8 $\mu\text{mol l}^{-1}$ (Fig. [3-e3c](#)). High total sulfide (S-II) concentrations of up to 5.6 mmol l⁻¹ were detected in 40 m water depth. Concentrations decrease upwards, fading in the lower part of the turbid layer at 36 m water depth (Fig. [3-e3c](#)). Below the 40 m depth level, S-II concentrations decrease to values around 3 mmol l⁻¹ down to 45 m water depth (Fig. [3-e3c](#) and Table S2). Concentrations of dissolved CH_4 ($\text{CH}_4(\text{aq})$) are low in the fresh water body with values of about 0.09 $\mu\text{mol l}^{-1}$. Methane 15 concentrations increase from the turbid layer at 36 m water depth downwards to values of 25 $\mu\text{mol l}^{-1}$ at 39 m water depth (Fig. [3-e3c](#)).

3.1.1 Calcite Saturation

The calculated saturation index (SI) of calcite shows calcite saturation in the fresh water body and the uppermost part of the halocline with values from 0.03–0.07 (Fig. S2). The SI closely follows the pH in the fresh water body revealing a distinct

20 peak of slightly higher values of $\text{SI} = 0.071$ in the turbid layer at ~36 m water depth. ~~However, SI values calculated for the water body below the redoxcline is undersaturated with respect to calcite indicating calcite dissolution in the halocline suffer from the overestimated pH readings in the extremely sulfidic water of the halocline and are therefore not considered. (Fig. S2).~~

3.1.2 Trace elements

25 Dissolved iron and manganese concentrations are very low in the fresh water body with concentrations of 0.1 and 0.01 $\mu\text{mol l}^{-1}$, respectively, and slightly increase within the turbid layer towards the salt water body, to concentrations of up to 0.47 (Fe) and 0.06 (Mn) $\mu\text{mol l}^{-1}$ (Fig. S3). Phosphate and silica concentrations are invariably low in the fresh water body ($\text{P}_{\text{ortho}} \sim 0.25$ and $\text{Si} \sim 63 \mu\text{mol l}^{-1}$) and increase in the salt water body peaking at 40 m water depth with concentrations up to 10.3 (P_{ortho}) and 275 (Si) $\mu\text{mol l}^{-1}$ (Fig. S3). Uranium content correlates to the redox potential of the water as indicated by uniform 30 contents of ~0.012 $\mu\text{mol l}^{-1}$ in the fresh water column and rapidly decreasing values in one order of magnitude in the turbid layer, to 0.0012 $\mu\text{mol l}^{-1}$ at 40 m water depth (Fig. 4 and Table S2).

3.1.3 Stable carbon isotopes of DIC and CH₄

The $\delta^{13}\text{C-HCO}_3^-$ values at water depth from 28 to 42 m are shown in Figure 5. The average $\delta^{13}\text{C-HCO}_3^-$ value is -9.8 ‰ in the fresh water body where DIC content is about 8 mmol l⁻¹. In the turbid layer $\delta^{13}\text{C-HCO}_3^-$ values show a distinct peak towards less negative values of up to -7.9 ‰ at slightly increasing DIC concentrations. Below the turbid layer $\delta^{13}\text{C-HCO}_3^-$ values rapidly decrease towards more negative values of -12.4 ‰ between 39 and 42 m water depth at increasing DIC concentrations (Fig. 4). A rather slight increase of $\delta^{13}\text{C-HCO}_3^-$ values (-11.6 ± 0.7 ‰) is observed towards the cenote bottom at 44 m water depth (Table S3).

The $\delta^{13}\text{C-CH}_4$ values are shown alongside with the CH₄ concentrations in Figure 5. The pattern of $\delta^{13}\text{C-CH}_4$ within the water column is similar to that of $\delta^{13}\text{C-HCO}_3^-$. In the fresh water body values of $\delta^{13}\text{C-CH}_4$ are approximately constant at about -49 ‰ and CH₄ concentrations are very low, roughly corresponding to that of atmospheric equilibrium (0.04–0.09 μmol l⁻¹). $\delta^{13}\text{C-CH}_4$ increases to -28 ‰ within the turbid layer and again decreases to -61 ‰ below the turbid layer, while CH₄ concentrations increase within and below the turbid layer (Fig. 4).

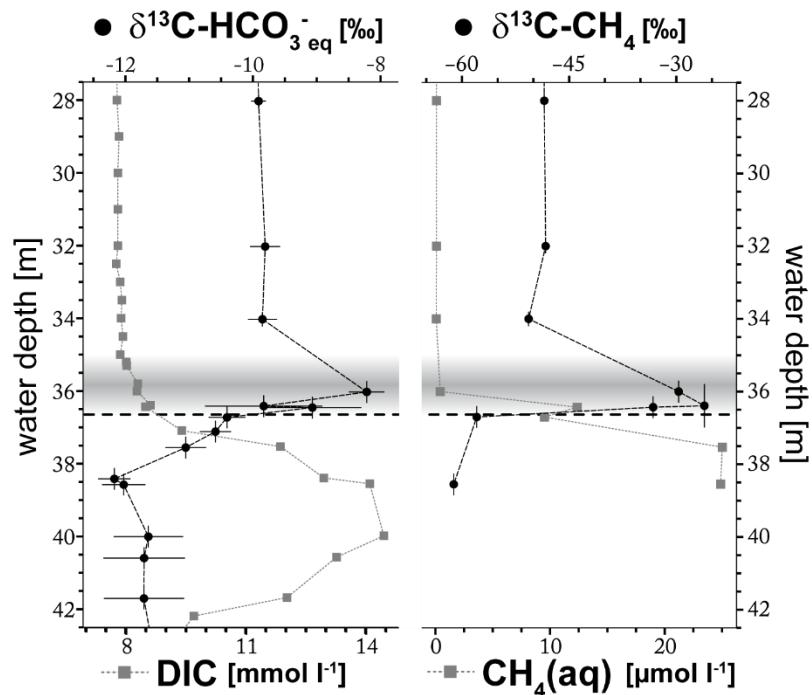


Fig. 4: Stable carbon isotope values $\delta^{13}\text{C-HCO}_3^-$ eq of the dissolved HCO₃⁻ in equilibrium with $\delta^{13}\text{C-CO}_2$ values measured in headspace and $\delta^{13}\text{C-CH}_4$ values of water samples alongside the concentrations of DIC and CH₄ of water samples. The grey band represents the turbid layer in 35–36.6 m water depth and the horizontal dashed line indicates the top of the halocline at 36.6 m water depth. Horizontal error bars represent 2σ uncertainties, vertical error bars indicate up to 0.6 m uncertainty of gas samples that were not taken from the sample used to determine chemical parameters (see section 2.1).

3.2. Petrography of Hells Bells speleothems

Petrographic characteristics of Hells Bells are shown in Fig. 56. The size of individual crystals of Hells Bells is varying from μm scale to several mm-sized crystals that are easily identified macroscopically. The latter are frequently dominant in the youngest calcite rims at the bottom of Hells Bells from water depths reaching from ~28 to ~35 m (Fig. 2-e2e).

5 Hells Bells from greater water depths show rounded or globular calcite surfaces at the lowermost margin of the speleothems indicating dissolution (Fig. 2-f2f). SEM-images of the lowermost part of Hells Bell surfaces are shown in Figs. 6-a6a, b and c. The calcite morphology varies from bladed or book-like calcite crystals (Fig. 5 a), dogtooth-like calcite crystals (Fig. 5-b5b) and blocky calcite rhombs (Fig. 5-e5c). In thin sections of the specimen ZPT-7 (Fig. 5 d1) (see also Stennesbeck et al., 2017b), these crystal morphologies are expressed as rather botryoidal (dog-tooth-like and bladed-shaped) and mosaic calcite phases
10 (blocky calcite rhombs) (Fig. S5). Electron images of the polished counter pieces of the thin section are shown in Fig. 5-d25d2. An element map of Mg shows that botryoidal calcite phases incorporated more Mg (appearing brighter in Fig. 5-d25d2) than the mosaic calcite phases (appearing darker in Fig. 5-d35d3).

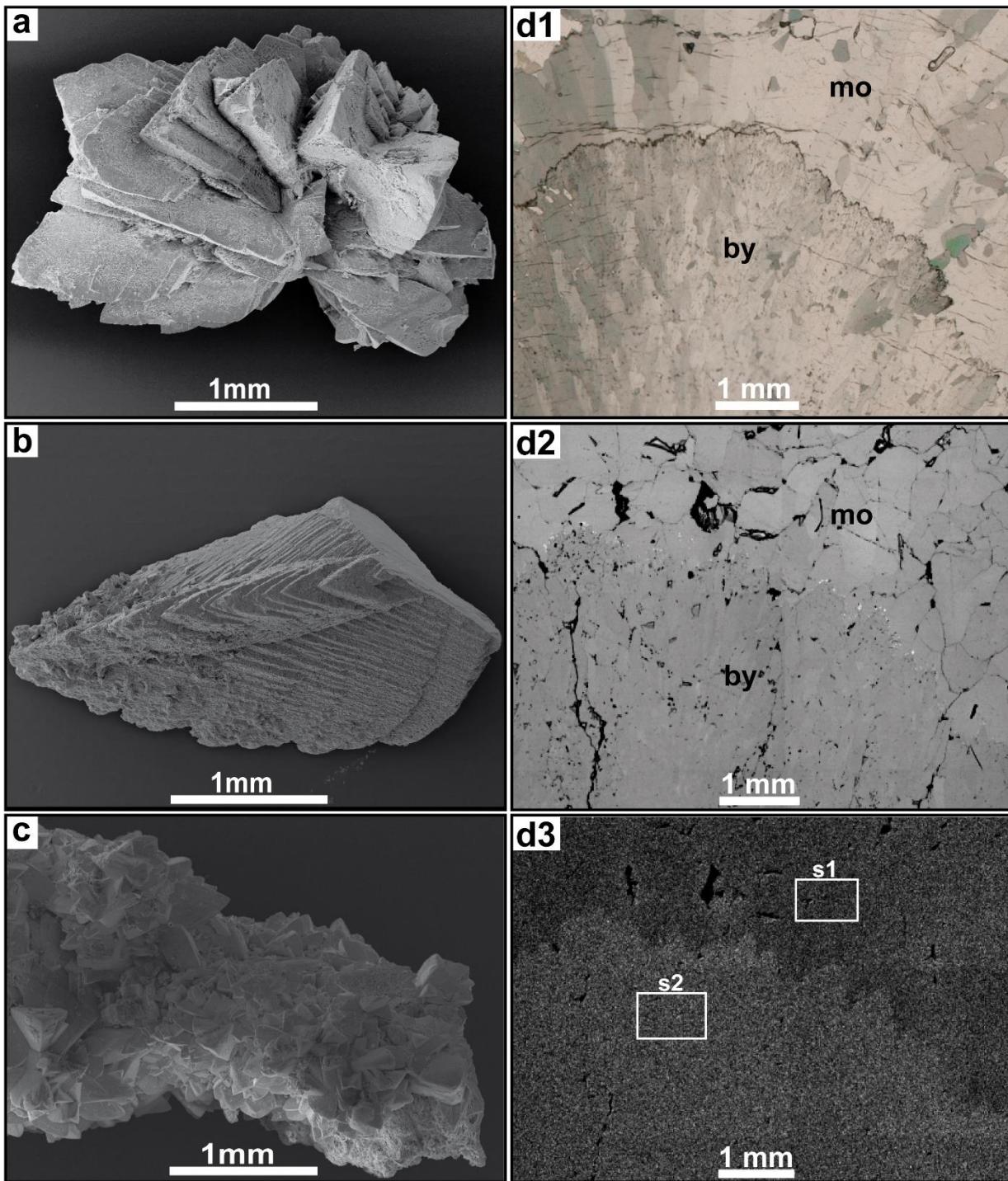


Fig. 5: Petrographic characteristics of Hells Bells speleothems. SE-images of Hells Bell samples Z17-8DC (a), Z17-18J (b), Z17-9J (c) of El Zapote cenote showing bladed (a), dogtooth-like (b) and blocky (c) calcite rhombs. Polarized transmitted light-microscopic images of a thin section from ZPT-7 (d1) (shown in Stinnesbeck et al., 2017b) showing different calcite fabrics of

angular coarse-grained mosaic calcite (mo) and fine grained elongated botryoidal calcite (by). The same detail is shown in the BSE-image of the polished counter slab that corresponds to the thin section (d2). The Mg-element map (d3), where higher abundances of Mg appear brighter, indicates a difference in Mg content between the botryoidal and mosaic calcite phases. The white rectangles represent areas of measured integrated element spectra.

5 3.3 Geochemistry of Hells Bells speleothems

Samples were collected from the lowermost and presumably youngest part of several Hells Bell specimens that grew on a ceiba tree fallen into the El Zapote cenote at about 3.5 cal kyr BP (Stinnesbeck et al., 2017b). They were analyzed for major and trace elements and stable carbon isotopes. The results are given in Table. 1.

3.3.1 Major and trace elements

10 The calcite of Hells Bell speleothems revealed no residues after digesting ~3 mg sample in 12 ml dilute 1 M Nitric indicating that Hells Bells calcite is devoid of acid solubleinsoluble impurities. The Mg/Ca, Sr/Ca and Ba/Ca ratios show narrow ranges with mean values of $22.5 \pm 2.9 \times 10^{-3}$, $38.6 \pm 5.9 \times 10^{-5}$ and $1.10 \pm 0.31 \times 10^{-5}$, respectively. They are closely related and positively correlate in each sample (Fig. 6-a6a). There is also a trend towards decreasing ratios with increasing water depth of the respective sample (Fig. 6-b6b). Iron and manganese show more variable concentrations with ratios of Fe/Ca and Mn/Ca 15 between $3.0 - 11.3 \times 10^{-5}$ and $16 - 39.3 \times 10^{-6}$, respectively. Iron and manganese show a weak positive correlation but no dependency on water depth. The content of sulfur in Hells Bells carbonate is constantly high with concentrations of $0.8 - 1.0$ g kg⁻¹ (Table S4) and mean S/Ca ratios of $2.87 \pm 0.42 \times 10^{-3}$, showing no dependency on water depth of the sample (Table 1).

3.3.2 Hells Bells stable carbon isotopes

Stable carbon isotope values of Hells Bells calcite samples ($\delta^{13}\text{C}_{\text{calcite}}$) from different water depth range from -12.85 to -13.82
20 ‰ with a mean value of -13.37 ± 0.70 ‰ (n=9, Table 1). There is a weak correlation of increasing $\delta^{13}\text{C}_{\text{calcite}}$ values with water depth of the samples (Fig. 6-b6b). Furthermore, $\delta^{13}\text{C}_{\text{calcite}}$ values show a strong positivenegative correlation with Sr/Ca and Ba/Ca with r^2 of 0.82 and 0.89 (Fig. 6-a6a).

The stable carbon isotope ratio of the HCO_3^- that is in equilibrium with the Hells Bells calcite ($\delta^{13}\text{C}_{\text{eqHCO}_3^-}$) at 25 °C water temperature ($\delta^{13}\text{C}_{\text{Calcite}} - 0.91$ ‰ = $\delta^{13}\text{C}_{\text{eqHCO}_3^-}$) was calculated after Mook (2000). The calculated $\delta^{13}\text{C}_{\text{eqHCO}_3^-}$ is -
25 14.28±0.70 ‰ which is slightly lower than the $\delta^{13}\text{C-HCO}_3^-$ determined for the water column with a range of -9.1 to -12.3 ‰.

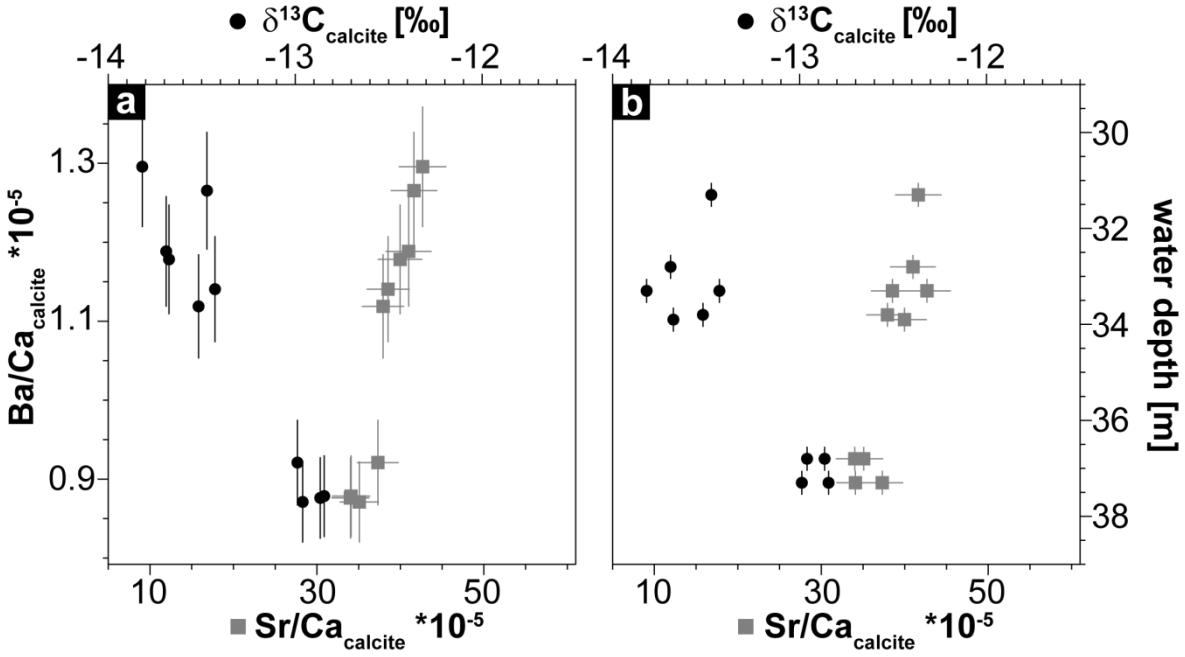
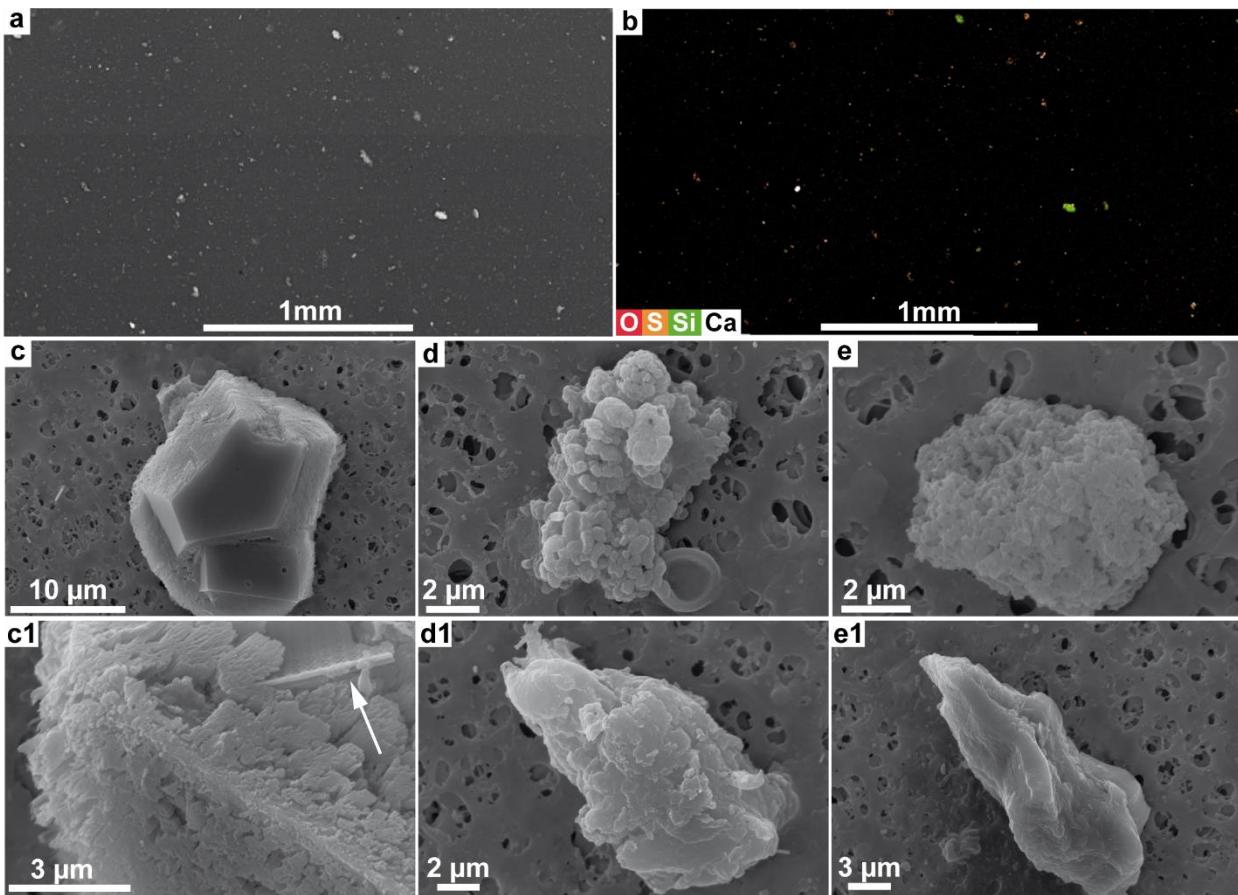


Fig. 6: Geochemical data of Hells Bell speleothems showing a strong correlation between Sr/Ca and Ba/Ca ratios ($r^2 = 0.91$) and between Ba/Ca and $\delta^{13}\text{C}_{\text{calcite}}$ ($r^2 = 0.89$) (a) and a trend of increasing $\delta^{13}\text{C}$ and decreasing Sr/Ca with increasing water depth of the samples (b). Given uncertainties represent 2σ standard deviations and ± 0.25 m is assumed as uncertainty for the water depth of the Hells Bell samples.

5

3.4 Turbid layer filtrate

Though Although the turbid layer appears dense in photographs taken during dives, the water sampled from the turbid layer was clear, with no visible turbidity during sample handling. Though, electron microscopy of the filter, however, reveals that abundant particles were extracted from the turbid layer (Fig. 7a-7a). Particle sizes range between 1–100 μm , but most are in the range of 1–10 μm . They consist of Ca-carbonate crystals (Figs. 7c-7c and c1), globular particles consisting of elemental sulfur (Figs. 7d-7d and d1) and silicate particles of different compositions (Figs. 7e-7e and e1). Also, numerous intact and broken shells of siliceous diatoms were found on the filter. Some calcite crystals incorporated broken parts of silica shells (Figs. 7e-7c1).



5 Fig. 7: SEM-analysis of turbid layer filtrate. Various particles on the filter are visible on the SEM picture of a larger area of the
 10 filter (a). An element map for O, S, Si and Ca of the same filter area of (a) is shown in (b) revealing that most particles consist of
 elemental sulfur, Ca-rich particles and silica particles. EDX analysis of individual particles on the filter verified the particles as
 Ca-carbonates (c, c1), elemental sulfur (d, d1) and silicate phases (e, e1). The white arrow in c1 points to a fragmentary silica shell.

4. Discussion

The detailed analysis of the water body and the geochemical analyses of Hells Bells presented in this study enables us to infer a comprehensive hypothesis on Hells Bells formation. In our previous study on Hells Bells we already argued that Hells Bells form under water in a lightless environment. We suspected the formation to take place in the lowermost fresh water body above the halocline (Stinnesbeck et al., 2017b). However, due to insufficient hydrogeochemical data we could only speculate on the processes leading to Hells Bells formation.

In the following discussion we deduce that the stagnancy of the water body favors the development of a pelagic redoxcline in which chemolithoautotrophic processes, mainly intense anaerobic sulfide oxidation, creates favorable conditions for calcite precipitation and hence, Hells Bells formation. Furthermore, we discuss the role of a dynamic halocline elevation and infer

~~the crucial factors for Hells Bells formation in El Zapote Cenote, which may give hints to the apparent exclusivity of these speleothems to only few cenotes in a very restricted area.~~

4.1 Limnological and hydrological conditions in El Zapote Cenote

The water temperature profile (Fig. 3-a3a) offers valuable clues on the hydrological conditions in the El Zapote cenote.

5 Mixing of the water in the narrow cenote shaft from 0–30 m water depth is indicated by constant temperatures and oxygenation, whereas linearly increasing temperatures in the wide dome-shaped cenote from 30–55 m water depth and linearly decreasing dissolved oxygen concentrations indicate conductive heat transport and oxygen diffusion, respectively (Fig. 3-a3a). This suggests that the water body from 0–30 m water depth is mixing-dominated and diffusion-dominated from 30–>50 m water depth. This interpretation is also supported by constant EC values in the cenote shaft and constantly 10 increasing EC values from 30 m water depth down to the top of the halocline at 36.8 m water depth (Fig. 3-a3a). Another indication for stagnant conditions of the water body is the shape of the halocline itself. Compared to other cenotes of the Yucatan Peninsula being deep enough to reach the halocline, El Zapote cenote particularly differs in the extent of the halocline, the transition zone from fresh to salt water. At El Zapote cenote, the halocline is about 10 m thick (Fig. 3a), as compared to a transition zone thickness of 1–5 m of other cenotes of Quintana Roo (Kovacs et al., 2017b); Stoessell et al., 15 1993).

The constant decrease of DIC, sulfide and orthophosphate below about 40 m water depth indicates a sink of these chemical species into depths greater than the cenote (>54 m water depth). This sink may result from advection of flowing water masses in conduits or zones of intensified hydraulic conductivity in a deeper cave system at around 60 m below the present sea level. Such deep cave systems could have developed during glacial sea level low stands (e.g. Smart et al., 2006).

20 In general, the water body of El Zapote cenote is stagnant from 30 m water depth down to the bottom of the cave where mass transfer is predominantly due to chemical diffusion. This is essential for the understanding of hydrogeochemistry and the ongoing biogeochemical processes in the El Zapote cenote.

4.2 Hydrogeochemical processes in El Zapote cenote

~~There might be main biogeochemical domains regulating the redox conditions of the meromictic El Zapote cenote, the organic matter rich sediments of the debris mount and the redoxcline. Therefore, we here discuss the biogeochemical processes occurring in the sediment and in the redoxcline of the water body.~~

4.2.1 Sedimentary biogeochemical processes

The anaerobic conditions and high concentrations of metabolites such as S(-II) and CH₄ can be attributed to anaerobic heterotrophic organic matter (OM) decay in the debris ~~mountmound~~ sediments. Both the debris ~~mountmound~~ and the cenote floor are covered with a relatively thick layer (~1 m) of OM, mostly leaves and other plant remains, according to the descriptions of the divers. As a consequence of stagnancy in the meromictic water body and oxygen deficiency on the cave

bottom, this OM is respiration by heterotrophic microorganisms in the sediment via anaerobic fermentative and respiratory pathways.

Anaerobic OM degradation by fermentation and sulfate reducing bacteria produce hydrogen and hydrogen sulfide (S(-II)), CO₂ (DIC) and acidity, thus lowering the pH. Elevated concentrations of DIC and S(-II) are found in the halocline (Fig. 3c), and low δ¹³C-HCO₃⁻ indicate a microbial origin of the hydrogen carbonate (e.g. Mook, 2000) (Fig. 4). Additionally, pH values are more acidic in the halocline (Fig. 3a and b) and sulfate reduction is further supported by decreasing SO₄²⁻/Cl⁻ ratios in the halocline of up to 32% compared to seawater ratio of 5.2 (Fig. 3-e3c) (Stoessell et al., 1993).

Methane-producing archaea (methanogens) metabolize degraded OM releasing CH₄ and DIC. Although this pathway is less energy efficient than sulfate reduction, and methanogenesis is not expected in the presence of sulfate, methanogens may dominate in deeper parts of the sediments where sulfate is already consumed. (e.g. (Whiticar, 1999)) Diffusion of CH₄ from the sediment into the water column leads to CH₄ concentrations of up to 25 μmol l⁻¹ identified in the halocline of El Zapote. Ammonium is likely released from organic matter degradation in the organic rich sediment and is also released to the water column at the halocline.

Other common anaerobic heterotrophic metabolic pathways in sediments, such as the reduction of iron, are subordinated processes, most likely due to the limitation low concentrations of iron oxides in the limestone karstic area and limited source of siliciclastic materials in this part of the Yucatán Peninsula. The elevated but still exceedingly low amounts of dissolved iron in the halocline as compared to the fresh water body (Fig. S3) are rather not indicative for the absence of these processes, as iron solubility is limited by the affinity to form iron sulfides in the presence of high amounts of S-II.

4.2.2 Water column biogeochemical processes

The redoxcline from 35 to 36.8 m water depth coincides with a peak in turbidity which is detectable both visually (Fig. 2-b2b) and geochemically (Figs. 3-a3a and b). Dissolved oxygen (DO) concentrations drop to undetectable levels at the top of the redoxcline, indicating that anaerobic biogeochemical processes prevail within the redoxcline (Fig. 8).

In our previous study we tentatively attributed these conditions to a full heterotrophic redox zonation due to organic matter decomposition (Stinnesbeck et al., 2017b). Fine organic matter accumulates along the density contrast at the top of the halocline and heterotrophic microbial communities thrive from the aerobic and anaerobic decomposition of this organic matter. This is also indicated in the results of this study by minor nitrification from ~34–35 m water depth (Fig. 3-b3b), non-linearly decreasing dissolved oxygen contents from ~34–35 m and by slightly more acidic pH values above and in the uppermost part of the turbid layer.

Nevertheless, the more detailed data presented in this study now underline the importance of planktonic chemolithoautotrophic processes in the pelagic redoxcline which are driven by the upward diffusion of reduced sulfur, carbon and nitrogen species released from the anaerobic degradation of organic material at the cenote floor. Pelagic redoxclines develop in density stratified marine (e.g Berg et al., 2015) as well as lake environments (e.g. Noguerola et al., 2015). In redoxclines below the photic zone the microbial community is dominated by chemolithoautotrophs, with a

considerable amount of chemoautotrophic production and dark carbon fixation (e.g. Grote et al., 2008; Jørgensen et al., 1991; Jost et al., 2010; Noguerola et al., 2015). The development of pelagic redoxclines was also reported for deep density stratified cenotes of the Yucatán Peninsula (e.g. Socki et al., 2002; Stoessell et al., 1993; Torres-Talamente et al., 2011).

In our previous study members β -proteobacteria *Hydrogeophilaceae* and the ε -proteobacteria genus *Sulfurovum* were 5 reported as dominant within the aqueous microbial community. Most members of these bacterial groups are chemolithotrophic or mixotrophic using reduced sulfur compounds or hydrogen as electron acceptors and oxygen or nitrogen-compounds as electron acceptors (Stinnesbeck et al., 2017b).

The white cloudy turbid layer could be the result of a dense accumulation of these microorganisms e.g. sulfur-oxidizing bacteria, analogous to that reported for Bundera sinkhole in Australia (Seymour et al., 2007). Elemental sulfur particles or 10 polysulfides were detected on the turbid layer filtrate and indicate sulfur oxidation in the turbid layer or redoxcline (sec 3.4 and Fig. 7); these particles are formed as intermediates in the microbial oxidation of sulfide (Findlay, 2016).

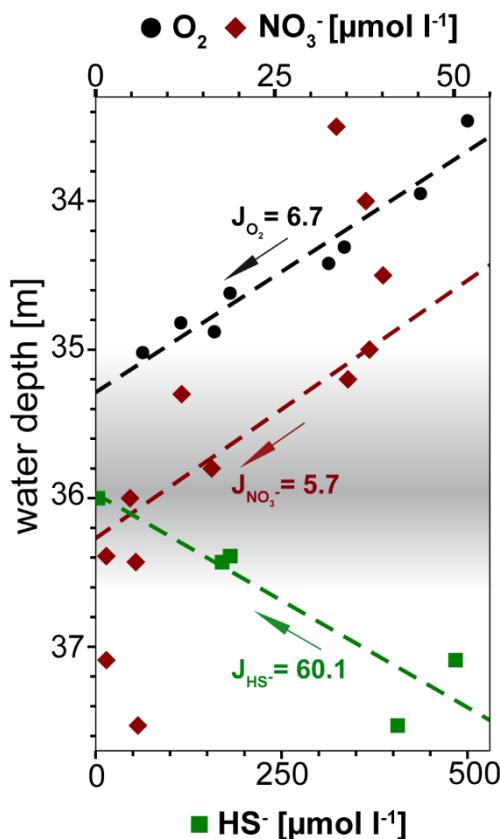
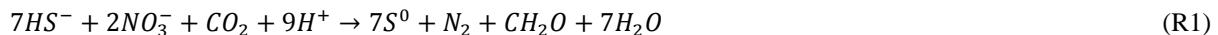


Fig. 8: Concentration profiles of dissolved O₂, NO₃⁻ and HS⁻ (calculated with PhreeqC) in water depths around the redoxcline. The fluxes J are given in $10^{-85} \mu\text{mol m}^{-2} \text{s}^{-1}$. The linear fit of O₂ and HS⁻ is calculated for the range of plotted values, while for NO₃⁻ it is calculated only for the range from 34.4–36.6 m water depth. Only O₂ values above detection limit ($6.3 \mu\text{mol l}^{-1}$) from winch profile 2 were considered for the calculation (Table S1).

The oxidation of sulfide in the redoxcline is likely anaerobic, as sulfide vanishes at around 36 m while dissolved oxygen is already at undetectable levels at 35 m water depth and both concentration profiles are not overlapping (Fig. 8). Furthermore, the oxygen flux towards the redoxcline is around one magnitude lower than the flux of the reduced sulfur species HS^- indicating that sulfide oxidation via aerobic pathways is minor (Fig. 8). Thus, sulfide oxidation within the redoxcline must be predominantly via anaerobic pathways. As the downward flux of NO_3^- towards the redoxcline intersects with the upward flux of HS^- (Fig. 8), assimilatory anaerobic sulfide oxidation could be obtained with nitrate as terminal electron acceptor producing elemental sulfur and nitrogen under the consumption of protons-[\(e.g. Bailey et al., 2009\)](#). The overall mass-balanced energy generating reaction for chemoautotrophic nitrate-driven anaerobic sulfide oxidation (ND-SO) is given in reaction (R1):



According to reaction (R1) ND-SO could account to one third of the HS^- oxidation, despite the flux of NO_3^- towards the redoxcline is around on order of magnitude lower than the HS^- flux (Fig. 8). Furthermore ND-SO is acid consuming and sulfide oxidation to elemental sulfur is more acid consuming than the full sulfide oxidation to sulfate (see also Visscher and Stoltz, 2005). The abundance of elemental sulfur particles found in the turbid layer filtrate (Fig. 7) indicates that sulfide oxidation to elemental sulfur is predominant. Full oxidation of sulfide to sulfate is less likely as no increase of sulfate is observed in the redoxcline (Fig. ~~3e~~^{3c}). Maxima in pH are known to occur when sulfide is oxidized to elemental sulfur with nitrate as electron acceptor (Kamp et al. 2006). In consequence, the minimum of nitrate in the redoxcline and slight alkaline pH shift, indicate that ND-SO is a relevant process in the redoxcline (Fig. ~~3b~~^{3b} and c, Fig. 8). Therefore, the proton consuming ND-SO could be the biogeochemical process in the redoxcline creating a disequilibrium in the carbonate dissolution-precipitation reaction, favoring calcite precipitation. This mechanism was recently reported for the formation of stromatolites below the photic zone of the Arabian Sea. There, a collective effect of proton-consuming ND-SO and alkalinity-producing sulfate driven-oxidation of CH_4 (SD-OM) leads to authigenic carbonate precipitation in microbial mats in the vicinity of CH_4 -seeps (Himmler et al., 2018).

Anaerobic SD-OM (e.g. Bailey et al. 2009) is likely to occur at the redoxcline, as dissolved CH_4 concentrations vanish at around the same depth of sulfide (~36.5 m) and $\delta^{13}\text{C}-\text{CH}_4$ values show a strong peak towards higher values at the same water depth (Fig. 4).

Autotrophy also supports calcite precipitation by taking up CO_2 for the synthesis of biomass (Castanier et al., 1999; Kosamu and Obst, 2009). Although a decrease of DIC is not observed at the redoxcline, chemolithoautotrophy is indicated by the $\delta^{13}\text{C}-\text{HCO}_3^-$ in the water body (Fig. 4). The peak of higher values in the redoxcline indicates inorganic carbon assimilation by microorganisms (dark- CO_2 uptake). ~~As organisms~~^{Organisms} usually prefer to metabolize ^{12}C (it takes less energy to break the ^{12}C bond instead of ^{13}C) ~~they effectively consume HCO_3^- with lower $\delta^{13}\text{C}$ values, which subsequently,~~, which results in higher $\delta^{13}\text{C}-\text{HCO}_3^-$ values in the remaining dissolved inorganic carbon. Hence, the peak towards more positive

$\delta^{13}\text{C}-\text{HCO}_3^-$ values identified in the redoxcline of El Zapote at ~36 m water depth, may be attributed to microbial CO_2 assimilation or dark CO_2 fixation.

4.3 Hypothesis on Hells Bells formation

It was ~~shown suspected~~ before that Hells Bells form within the freshwater ~~indicated by $\delta^{234}\text{U}_{\text{initial}}$ values~~ body of 16–25‰ of the Hells Bells calcite ~~El Zapote cenote~~ (Stinnesbeck et al., 2017b). The depth zone of Hells Bells formation within the fresh water layer can now be narrowed down by the application of given distribution coefficients $D(\text{Mg})$ of the temperature-dependent partitioning of Mg into calcite in equation (Eq. 2):

$$(Mg/Ca)_{\text{solution}} = \frac{(Mg/Ca)_{\text{solid}}}{D(Mg)} \quad (\text{Eq. 2})$$

Applying the mean value of $Mg/\text{Ca}_{\text{solid}}$ determined for Hells Bells calcites (Table 1) and $D(\text{Mg})$ at 25°C given by Huang and Fairchild (2001) and Rimstidt et al. (1998), the calculation of Mg/Ca of the solution from which the Hells Bells calcite precipitated yields a $Mg/\text{Ca}_{\text{solution}}$ of 0.73 and 1.06, respectively. $Mg/\text{Ca}_{\text{solution}}$ ratios in this range are found in the water of the redoxcline and the uppermost top of the halocline in 36–37 m water depth, thus supporting the interpretation that Hells Bells formation takes place in the redoxcline (Fig. 3-e3c and Table S4).

The calcite crystals found in the turbid layer filtrate give further hints on calcite precipitation in the redoxcline (Fig. 7-e7c). It is not yet known whether these particles represent autochthonous matter of the turbid layer. Nevertheless, formation of calcite crystals at the density boundary is likely, as fine particulate matter is accumulated there and may act as crystallization seeds. This process is indicated by calcite crystal formation around silica shells (Fig. 7-e7c1). The high sulfur contents found in Hells Bells calcite also supports this assumption as small sulfur particles are abundant in this water layer and are easily enclosed in calcite crystals growing there (Table 1).

Based on the indications of Hells Bells formation in the redoxcline and taking the biogeochemical processes discussed before into account we propose the following scenario illustrated in Figure 9. It summarizes the biogeochemical processes inducing calcite oversaturation and calcite precipitation in the turbid layer and the redoxcline of El Zapote cenote. Heterotrophic bacterial decomposition of organic matter in sediment of the debris ~~mount~~mound releases CO_2 (HCO_3^-), nutrients (P_{ortho}) and reduced species of sulfur (S-II) and nitrogen (NH_4^+). Due to the stagnant conditions in the cenote, these species are transported via diffusion, thereby allowing for the formation of a defined and stable redoxcline. Here, anaerobic chemolithoautotrophy, and especially proton consuming nitrate-driven sulfide oxidation, (ND-SO), increase alkalinity, thus favoring calcite precipitation (Fig. 9). The required Ca^{2+} ions for calcite precipitation are constantly supplied to the redoxcline by upward diffusion from the Ca-enriched saline water body (Stinnesbeck et al., 2017b).

Biogeochemical processes involved in Hells Bells formation

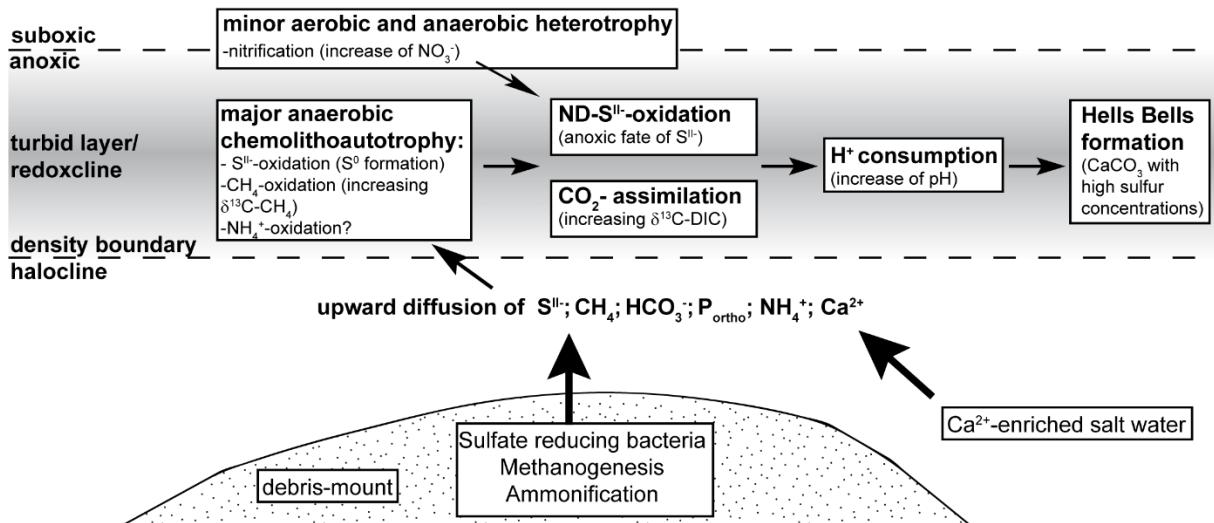


Fig. 9: Scheme of the biogeochemical processes involved in the sediment and redoxcline of the water column of the El Zapote cenote that lead to Hells Bells formation.

However, 4.3.1 Calcite precipitation rates

5 In order to test the plausibility of the hypothesis on Hells Bells formation within the redoxcline, calcite precipitation rates of both biogeochemical processes in the redoxcline alone, ND-SO and CO₂-assimilation are not sufficient to explain the shape of Hells Bells nor their occurrence over a wide vertical zone of around 10 m water depth. This implies assessed. All used parameters and results are given in Table S5.

The overall chemical reaction of the carbonate balance is given in R2 and the partial reactions of R2 are given in R2a–R2d.

10 Under equilibrium conditions, carbonate precipitation after R2 is acid producing because for each mole of precipitated calcite in reaction (R2) one proton is released to compensate for the abstracted carbonate ion (R2b) due to proton shift in the partial reactions (R2b–R2d).



Assuming that another essential up to one third of the hydrogen sulfide flux towards the redoxcline (Fig. 8) is oxidized by the proton consuming ND-SO after reaction (R1) and that the proton consumption of this process is buffered by both, the

observed increase in pH values of 0.04 within the redoxcline (Fig. 3b) and calcite precipitation after reaction (R2), allows for an estimation of the calcite precipitation rate R_{ND-SO} derived from ND-SO within the redoxcline after equation (Eq. 3).

$$R_{ND-SO} = \frac{1}{3} J_{HS^-} * \left(\frac{\alpha}{mHS^-} - \frac{\beta}{mHS^-} \right) \left[\frac{mol}{m^2 s} \right] \quad (\text{Eq. 3})$$

With $mHS^- = 7$ (moles of HS^- in R1); $\alpha = 9$ (moles of H^+ in R1); $\beta = 0.91$ (moles of H^+ consumed in pH increase of 0.04).

5 Equation 3 yields that 7.3 mmoles calcite $m^{-2} a^{-1}$ or 0.73 g calcite $m^{-2} a^{-1}$ could be precipitated within the redoxcline due to ND-SO. This adds up to a total calcite precipitation rate R_{ND-SO} of 2.2–6.2 kg calcite a^{-1} in the whole redoxcline of the circular 60–100 m wide El Zapote cenote.

10 Additionally, the second biogeochemical process in the redoxcline that might lead to calcite precipitation, the CO_2 -assimilation or dark CO_2 -fixation (Fig. 9), has to be taken into account. Under equilibrium conditions, for each mole of

10 abstacted or assimilated CO_2 one proton is consumed due to the carbonate balance (R2a–R2d) and this proton consumption is compensated by calcite precipitation (R2). The CO_2 -assimilation rate in the redoxcline can be approached by the upward flux of DIC ($J_{DIC(HCO_3^-)}$) towards the redoxcline (calculated as HCO_3^- and given in Fig. S4). The calcite precipitation rate $R_{CO_2-assim}$ derived from the CO_2 -assimilation is then limited to the half of the DIC flux towards the redoxcline as given in equation (Eq. 4).

$$15 R_{CO_2-assim.} = \frac{1}{2} J_{DIC(HCO_3^-)} \quad (\text{Eq. 4})$$

15 Equation 4 yields a calcite precipitation rate $R_{CO_2-assim.}$ of 40 mmoles calcite $m^{-2} a^{-1}$, which is 12–34 kg a^{-1} in the whole redoxcline. Such rates seem reasonable when compared to reported dark CO_2 -fixation rates from 0.2–2.7 $\mu\text{mol } CO_2 l^{-1} d^{-1}$ of deep marine pelagic redoxclines in the Black Sea (Jørgensen et al., 1991), the Baltic sea (Glaubitz et al., 2009; Jost et al., 2008) as well as for a deep karstic lake pelagic redoxcline (Noguerola et al., 2015) (Table S5). Applied to the redoxcline of 20 El Zapote cenote, these rates would yield calcite precipitation rates from 11–420 kg calcite a^{-1} in the 0.5 m thick and 60–100 m wide redoxcline of the circular El Zapote cenote.

20 The summed up calcite precipitation rates of R_{ND-SO} and $R_{CO_2-assim.}$ in the redoxcline can be converted to calcite growth rates from 0.27–1.46 $\mu\text{m } a^{-1} m^{-2}$, taking a calcite density of 2.7 g cm^{-3} into account. Calcite growth, however, is likely to be concentrated on a much smaller area, i.e. the crystal or substrate surfaces of Hells Bells hanging in the redoxcline, the cenote walls and the tree stem within the redoxcline (Figs. 2 and 3). This would result in a higher actual calcite growth rate, for example if the calcite growth is concentrated to 1 % of the area, this would result in actual growth rates of 27–146 $\mu\text{m } a^{-1}$.

These growth rates are close to the reported net growth rates of $12\text{--}90 \mu\text{m a}^{-1}$ for a U-series dated Hells Bells specimen from the tree (Stinnesbeck et al., 2017b) demonstrating the plausibility of Hells Bells formation by the biogeochemical mechanisms proposed in this study. However, the actual calcite growth rates must be involved in Hells Bells formation, which is significantly higher than the reported net growth rates as these rates also comprise repeated elevation of the haloeline-phases of calcite dissolution (see Secs. 3.1.1 and 3.2; Stinnesbeck et al., 2017b).

Eventually, the comparison of the estimated actual calcite precipitation rates in the redoxcline and the reported net growth rates is hindered by a lack of data of both, the actual area of calcite precipitation and the time and intensity of calcite dissolution.

4.3.12 The role of halocline elevation in Hells Bells formation

Hells Bells formed in modern to historic times and occur in a relatively wide vertical zone of about 10 m from 28–38 m water depth (Stinnesbeck et al., 2017b). Hells Bells formed in modern to at least historic times and occur in a relatively wide vertical zone of about 10 m from 28–38 m water depth (Stinnesbeck et al., 2017b). This indicates that their underwater growth occurred under environmental conditions similar to the ones detected by us, as modern sea water levels were already reached at about 4.5 ka BP (Hengstum et al., 2010) and thus significantly earlier. Modern Hells Bells therefore precipitate either permanently in the entire depth zone reaching from 28–38 m, or in the narrow 1–2 m wide redoxcline or turbid layer above the halocline (Fig. 9). According to the data presented here the latter hypothesis appears much more likely to us. Therefore, we propose that growth of Hells Bells is a non-permanent episodic process which majorly depends on a highly variable the halocline elevation in the cenote (Fig. 10). The halocline position is a function of hydrostatic pressure of the overlying fresh water layer, with a general trend of increasing depth with increasing distance to the coast (Fig. 10a, 10a) (e.g. Bauer-Gottwein et al., 2011). Extraordinary recharge events (e.g. hurricanes) must have a significant effect on the depth position of this layer (Fig. 10b). During Extensive droughts occurred repeatedly in the Holocene (Hodell et al., 2001) and could have led to a prolonged elevation of the halocline as the fresh water layer was thinner in consequence of decreased precipitation (Evans et al., 2018). Although on much shorter timescales, extraordinary recharge events (e.g. hurricanes) must have an effect on the depth position of the halocline (Fig. 10b). Generally, during these events of enormous precipitation, the halocline is temporarily pushed downwards by the amount of new the fresh water infiltrating into the Yucatán karst aquifer. However, the halocline can also be elevated as a response to precipitation events. Escolero et al. (2007) detected a observed a piston-like effect regarding the halocline rise of 17.5 melevation in a response to recharge in the Northern Yucatán aquifer well and ascribed this to a piston like effect. During recharge events, the water flow is predominantly vertical and salt water is temporarily pushed down. In the period after these events, the halocline bounces back and oscillates until a static equilibrium is again reached. Additionally, Escolero et al. (2007) detected a halocline elevation of 17.5 m in between measurements in the years 2000 and 2003 in an observation well indicating a highly dynamic halocline elevation within the Yucatán karst aquifer, that might be recharge-driven.

Regional and local vertical and lateral hydraulic transmissivities of both the epi- and the phreatic karst can ~~also~~ result in spatially variable hydraulic pressure of the fresh water lens (Williams, 1983), thus leading to a lowered halocline beneath areas of higher, and an elevated halocline beneath areas of less vertical hydraulic transmissivity. El Zapote cenote is located in the Holbox fracture zone that in the area is characterized by N-S trending lineaments of increased permeability (Bauer-Gottwein et al., 2011, and references therein). Hurricanes that pass the area frequently (Farfán et al., 2014) ~~will~~could therefore lead to episodic elevation of the halocline (Fig. 10).

5

Halocline dynamic at El Zapote Cenote

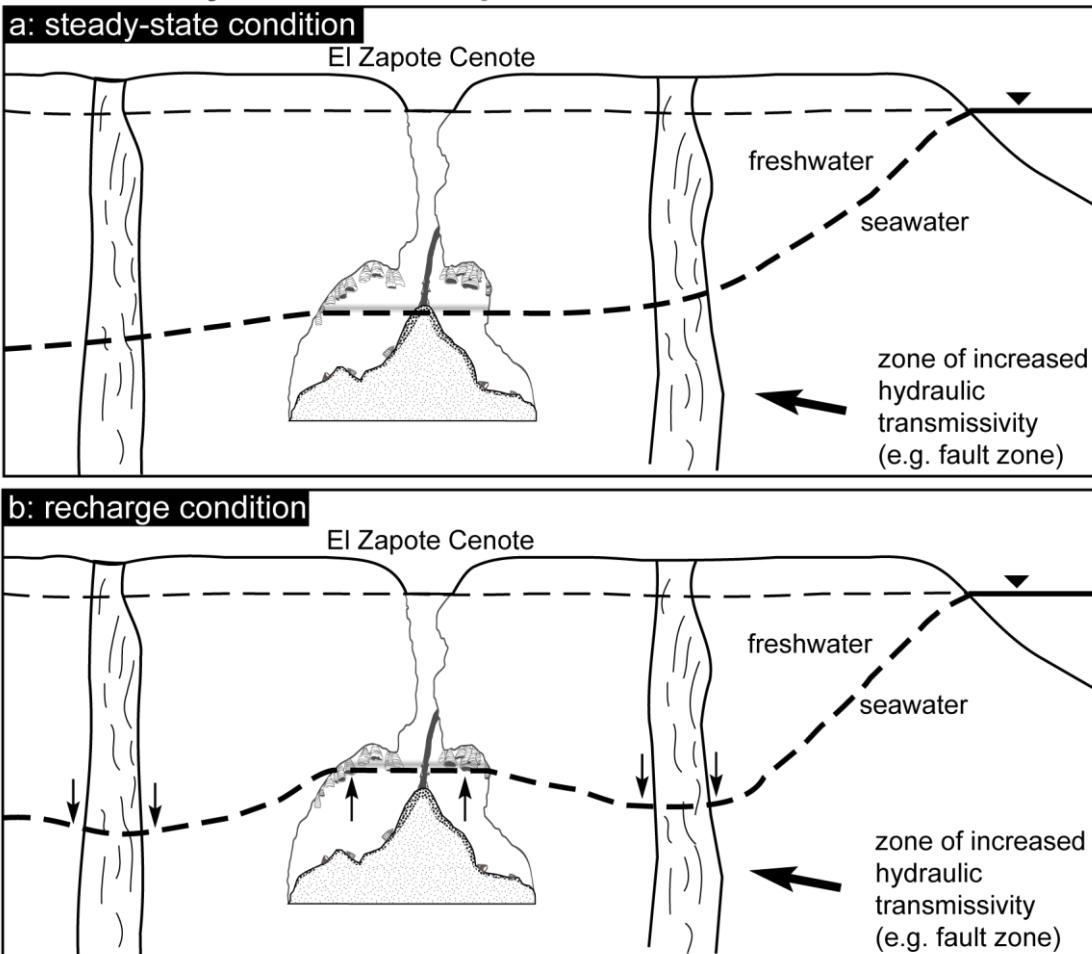


Fig. 10: Sketch of dynamic halocline elevation within the Yucatán Karst aquifer. Halocline depth increases with increasing distance from the coast in a steady-state condition (a). Recharge events (e.g. hurricanes) result in a lower halocline beneath areas of high vertical transmissivities and an elevated halocline in areas of low hydraulic transmissivity, e.g. El Zapote cenote (b).

10

~~Although there is yet Both, a lack of data to allow for a positive identification of El Zapote as a prolonged halocline elevation during droughts and short-term but frequent recharge-driven cenote, the process explains halocline elevations could result in~~

the presence of Hells Bells in a zone of 28–38 m water depth. ~~During an episodic rise of the halocline, the zone of Hells Bells formation (redoxcline) also rises, thus inducing calcite formation in a comparably shallow water depth (Fig. 10).~~ Our model also explains ~~Furthermore, repeated phases of precipitation and dissolution indicated by~~ the alternating ~~presence of~~ layers of dog-tooth calcite and microcrystalline calcite ~~indicating dissolution of Hells Bells thin sections~~ (Stinnesbeck et al., 2017b~~+~~) suggest an episodic halocline elevation. During episodes of an elevated halocline, ~~precipitation or~~ dissolution of Hells Bells may occur in ~~the lower depth range depths~~ due to the ~~intermittent increase of concurrent elevation of the calcite precipitating redoxcline and the underlying~~ sulfide-rich and carbonate-undersaturated water.

The time-scale of the frequent recharge-driven episodic halocline elevations is unclear to date, but could reach from a few days up to several weeks. Nevertheless, the estimations of calcite precipitation rates in the redoxcline suggest that Hells Bells growth on such short time-scales is possible (Sec. 4.3.1).

A ~~dynamic~~variable halocline depth position at El Zapote cenote is also supported by the positive correlation with water depth of Sr/Ca and a negative correlation of $\delta^{13}\text{C}$ of the Hells Bells calcite (Fig. 6). Hells Bells formed in lower water depths show slightly higher contents of the trace elements Sr, Mg and Ba and slightly lower $\delta^{13}\text{C}$ values than Hells Bells formed in greater water depths (Table 1 and Fig. 6). The higher incorporation of the trace elements Sr, Mg and Ba is either obtained by faster growth rates (Tesoriero and Pankow, 1996), or by elevated concentrations in the solution from which the calcite ~~is~~ precipitated. The latter process is more likely, as the amount of seawater increases in the turbid layer when the halocline is located at lower water depths. Lower $\delta^{13}\text{C}$ values support this assumption, as lowest $\delta^{13}\text{C-HCO}_3^-$ values are detected in the halocline of the modern El Zapote cenote (Fig. 4).

The increase in seawater in the lowermost fresh water and the turbid layer could result from ~~turbulence~~increased salinity of the freshwater body during droughts. Furthermore, turbulences induced by halocline elevation as a reaction to recharge events (Kovacs et al., 2017b). Minor mixing of the water bodies would be sufficient to increase the concentrations of Sr, Mg, Ba and decrease $\delta^{13}\text{C-HCO}_3^-$ in the turbid layer, as seen at El Zapote.

Following this model of a hydraulically-driven halocline elevation, the question remains why Hells Bells are restricted to a zone of 28–38 m water depth. This range could solely depend on the hydraulic conditions, e.g. Hells Bells formation reflecting maximum and minimum elevations of the halocline during recharge events. The lower (38 m) level of Hells Bells formation may represent the stable environmental conditions in the modern El Zapote cave, influenced only by the thickness of the freshwater body and the mean sea level. The upper range boundary, on the other hand, could well be given by the shape of the sinkhole and limnological conditions in the narrow cenote shaft reaching from 0–30 m water depth. In this latter unit the water body is mixing- rather than diffusion-dominated (Fig. 3. Section 4.1). A rise of the halocline to about 28 m water depth would therefore lead to an exposure to fast and convective oxygen supply from the mixed-in fresh water body above, and consequently to aerobic microbial sulfide oxidation, which is an acid-producing reaction (e.g. Jones et al., 2015). Hells Bells formation would then stop as it is tied to the anaerobic ND-SO. The occurrence of a zone of brown-colored manganese oxide coatings on Hells Bells and cave wall at and above 30 m water depth, indicates that the redoxcline must

temporarily have reached up to this level (Fig. 2-e2c, d and Fig. S4S6). Manganese dissolved in the halocline and turbid layer was then oxidized to manganese oxide precipitates (Fig. S4S6).

The dynamic history of halocline elevation at El Zapote cenote cannot be resolved to date but raises an interesting issue of further research on the dynamic hydraulic response of the Yucatán aquifer to extraordinary recharge events, especially as this process could be a key factor for the formation of Hells Bells. We are currently addressing this issue by logging the hydraulic head of the fresh water body and the electrical conductivity at a fixed position in the halocline of El Zapote cenote.

4.3.23 Shape of Hells Bells

Previously, we attributed the growth of Hells Bells to microbial mediation (Stinnesbeck et al., 2017b). We hypothesized that autotrophy (ammonia oxidation) and denitrification are the main factors that trigger calcite precipitation at the surface of the Hells Bells and that calcite precipitation could further be supported by the presence of negatively charged extracellular polymeric substances (EPS), leading to the accumulation of Ca^{2+} ions and to supersaturation of calcite within biofilms (e.g. Dupraz et al., 2009). However, the large size and form of the dog-tooth calcite crystals of Hells Bells rather resemble slow growing inorganic calcite crystals rather than biologically-mediated precipitates (Fig. 2-e2e and Fig. 5). This hypothesis is supported by Bosak & Newman (2005) who investigated microbial kinetic controls on calcite morphology and found that microbially mediated calcite precipitated at low calcite supersaturation shows more anhedral crystal morphologies, compared to the more euhedral abiotic ones. Although the microbial activity in the redoxcline induces calcite oversaturation, the hypothesis presented in this study is compatible with an inorganic calcite precipitation of the Hells Bells from a biologically mediated water layer.

Ultimately, the hypotheses of dynamic halocline elevation and biogeochemically induced calcite precipitation in the redoxcline can be integrated. to explain some morphological features of the Hells Bells. They grow downward and are conically divergent, with a strict horizontally lower margin and a hollow interior. Specimens also tend to be oriented towards the cenote center (Stinnesbeck et al., 2017b). The horizontal downward growth is indicative for a precipitation from a defined layer within the water column (i.e. the redoxcline). Also, an abrupt elevation of the redoxcline as a response to recharge events and a subsequent decelerated drop towards its original position serves to explain the downward growth of Hells Bells. This is indicated by the tendency towards downward orientation of the calcite crystal growth axis (Fig. 2-e2e). The fact that Hells Bells specimens growing on the inclined cave wall are always oriented towards the cenote center could result from a lateral gradient in the chemolithoautotrophic intensity. The “energy source” (i.e. sulfide) used by the chemolithotrophic microbial community in the redoxcline is the release of reduced carbon, sulfur and nitrogen species from anaerobic organic matter decay in the organic-rich sediments on the debris mountmound. Both the morphology of the cenote and the diffusive mass transport likely result in radial concentration gradients of upwards diffusing reduced species from sediment of the debris mountmound. These conditions limit the availability of reduced species in locations of the redoxcline distal to the debris mountmound, and vice versa. Consequently, the intensity of chemolithoautotrophy and hence calcite oversaturation is preferentially higher in the center proximal to the debris mountmound and decreases towards the cenote

walls. This accounts for both the inclined bells as well as for horse-shoe like horizontal openings of Hells Bells which always face towards the wall. Furthermore, the often observed hollow interior of Hells Bells could be due to preferential growth of the outer edges of Hells Bells, especially the parts facing towards the cenote center. Once such parts of Hells Bells grew slightly more, hence into deeper water depth, this will result in a higher net growth as these parts are likely to be reached more frequently by the narrow zone of calcite precipitation (redoxcline). The conical shape and downward divergence of Hells Bells could be a continuation from the microscopic to the macroscopic level as the angles of the botryoidal calcite phases in thin section (Fig. S5) strongly resemble those of the large specimens of Hells Bells (Fig. 2d).

Nevertheless, we can also not exclude the potential influence of microorganisms forming a biofilm community on the surface of Hells Bells. Stinnesbeck et al. (2017b) showed that this community does not resemble the planktonically growing microbial biocenosis but forms a distinct community that seems to thrive catalyzing the reduction and oxidation of different nitrogen species. However, it is not known to date to what percentage the activity of these organisms contributes to the shape of the speleothems.

4.4 Prerequisites for the formation of Hells Bells

Hells Bells have so far been identified in a few cenotes only of a restricted area of the North-Eastern Yucatán Peninsula (Stinnesbeck et al., 2017b), although the peninsula hosts many thousands of sinkholes (Bauer-Gottwein et al., 2011). The question thus arises which factors are needed for the generation of these underwater speleothems. The following apparent prerequisites for Hells Bells formation appear likely to us:

- The cenote or sinkhole must be deep enough to reach the halocline in order to have a density stratified water column (meromixis).
- Sufficient input of organic material to the cenote bottom is required to create anoxia in the halocline with a release of reduced sulfur, carbon and nitrogen species.
- A meromictic stagnant water body is needed that allows for the formation of a redoxcline in which anaerobic chemolithoautotrophy prevails in a lightless environment. This leads to a narrow zone of calcite oversaturation in the water body.
- Special hydraulic conditions are needed which allow the halocline to rise and fall in order to form subaqueous speleothems.

5. Conclusion

The unique underwater speleothems termed Hells Bells recently described from El Zapote west of Puerto Morelos on the northern Yucatan Peninsula, Mexico, are most likely formed in the redoxcline, a narrow layer in the lowermost fresh water body immediately overlying the halocline. We propose a biogeochemical mechanism for the formation of these structures, that induces calcite oversaturation favoring calcite precipitation within the redoxcline. The upward diffusion of reduced sulfur, carbon and nitrogen stimulates a chemolithoautotrophic microbial community thriving above the halocline at El Zapote cenote. Chemolithoautotrophy and proton-consuming nitrate-driven anaerobic sulfide oxidation lead to calcite

precipitation, and hence Hells Bells formation, in a narrow depth zone confined to the redoxcline, or turbid layer. We further postulate a dynamic elevation of the halocline as an episodic hydraulic response to both droughts and recharge events, e.g. hurricanes, that accounts for Hells Bells occurrence over a vertical range of 10 m water depth.

Video supplement

- 5 <https://doi.org/10.5446/39353>

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Author contribution

- 25 WS and AGG initialized the project and funding was acquired by WS, AGG and MIS. Sampling was planned and conducted by SR and CS. Underwater sampling was conducted by JAO, DP, CL and EAN, while underwater videographer TV was documenting part of the sampling producing the footage for the video supplement. Instrumentation and methodology was provided by MIS for hydrogeochemical analyses, FK for gas and stable carbon isotope analytics, AH and GA for light

microscope-imaging and AV for SEM-imaging and GA and. SR, LK and NS collected the data and CS, MIS, FK and AV validated it. SR, CS, MIS, LK, NS interpreted the results. SR developed the hypotheses equally supported by CS and MIS. SR visualized the data and prepared the original draft, MIS supervised. JG, WS and FK critically reviewed the manuscript.

Competing interests

- 5 The authors declare that they have no conflict of interest.

Disclaimer

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Table 1: Geochemistry of samples from the lowermost tips of Hells Bells growing on a subfossil ceiba tree that fell into the El Zapote cenote about 3500 cal yr BP. Individual tree bells were sampled at water depths from 31.3 to 37.3 m. The lack of samples in water depths from 34.3 to 36.8 m water depth is due to poor visibility in the turbid layer above the halocline (compare Fig. 1b). Given uncertainties represent 2 σ standard deviations.

Sample No.	water depth [m] ^a	Mg/Ca [$\times 10^{-3}$]	Sr/Ca [$\times 10^{-5}$]	Ba/Ca [$\times 10^{-5}$]	Fe/Ca [$\times 10^{-5}$]	Mn/Ca [$\times 10^{-6}$]	S/Ca [$\times 10^{-3}$]	$\delta^{13}\text{C}$ [‰VPDB]
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1	31.3							
2	31.3	21.4	41.6	1.27	4.3	21	2.78	-13.47 ± 0.01
3	32.8	25.6	41.0	1.19	5.2	26	2.69	-13.69 ± 0.01
4	32.8							
5	33.3	22.0	42.7	1.30	3.8	24	3.05	-13.82 ± 0.01
6	33.3	22.3	38.5	1.14	3.0	16	2.54	-13.43 ± 0.01
7	33.8							
8	33.8	20.8	37.9	1.12	3.5	24	2.72	-13.52 ± 0.01
9	33.9	23.2	40.0	1.18	3.9	22	2.76	-13.68 ± 0.01
10	33.9							
11	36.8	22.3	34.0	0.88	11.3	39	3.14	-12.87 ± 0.02
12	36.8	21.1	34.1	0.88	10.6	28	3.02	-12.85 ± 0.00
13	37.3	24.0	37.3	0.92	6.2	32	3.17	-12.99 ± 0.01
14	37.3							
	Mean	22.5	38.6	1.10	5.76	25.8	2.87	-13.37
	2σ	2.9	5.9	0.31	5.84	12.8	0.42	0.70