

Author's response to the reviews of the revised manuscript "Subaqueous speleothems (Hells Bells) formed by the interplay of pelagic redoxcline biogeochemistry and specific hydraulic conditions in the El Zapote sinkhole, Yucatán Peninsula, Mexico" by Ritter et al., submitted to Biogeosciences

5 Dear Dr. Wajih Naqvi,

On behalf of all Co-authors, I would like to thank you and the anonymous referees for the excellent editorial handling and reviewing of our manuscript. Please find attached our point-by-point response to the comments of anonymous referees #1 and #2 to our revised manuscript and a list summarizing the relevant changes made to the revised manuscript.

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Sincerely,
Simon Ritter

15 **Author's response to the comments of anonymous referee #1:**

Referee:

I'd like to reiterate from my previous review that I think the authors have done an excellent job examining the processes controlling Hell's Bells formation over a series of manuscripts. I'm looking forward to reading the additional manuscripts they mention are in preparation.

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I have only a few general criticisms and hope to see the paper published in due course. First, the paper is now entirely too long. Second, and related to the first, at least some of the length of the manuscript can be reduced by removing the detailed descriptions of all of the different ways that the halocline can oscillate (see comments related to Page 26). I briefly expand on these points in my line-by-line comments, below.

25 *Page 1, Line 30 – I recommend adding annual tidal variability to the list of possible causes for halocline oscillation. The authors mention this in their rebuttal, but do not include it in the manuscript.*

Response: We added "annual tidal variability" as suggested by the referee.

Page 2, Line 5 – Stalactites, by definition, only form by dripping water. They cannot form underwater. Please change to "speleothems" or "secondary precipitates."

30 Response: Corrected as suggested. We also deleted the previous sentence as the information of how stalactites form is no longer needed.

Page 2, line 12 – meter should be plural.

Response: Corrected as suggested.

Page 2, line 15 – *if they are above the anoxic/sulfidic zone, how is the environment “toxic”?*

Response: That’s correct, “toxic” is now deleted in the manuscript.

Page 3, line 4 – *there is no plausible method of estimating unmapped cave lengths. I recommend removing the suggestion that there is more than 7,000 km of possible cave.*

Response: Corrected as suggested.

Page 3, line 6 – *Karst cave is probably redundant in this context.*

Response: Corrected as suggested. We deleted “karst”.

Page 3, line 10 – *recommend changing “seawater” to “saline water” as the geochemical composition of some of this water is distinct from seawater.*

Response: Corrected as suggested.

Page 3, line 11-14 – *the halocline has been proposed to be an area of carbonate mineral undersaturation on the basis of numerical models and geomorphology, however, geochemical studies rarely find undersaturation due to mixing (which is the point of the Gulley et al. 2016 paper that is cited here as evidence of undersaturation in haloclines).*

Response: We are aware that the reason of the undersaturation with respect to calcite in the halocline is discussed controversial in the literature. Therefore, we avoided going into detail on the origin of the undersaturation (e.g. mixing dissolution or dissolution due to elevated PCO_2 derived from microbial decomposition of Corg) and only state that the halocline “...is usually characterized by undersaturation with respect to CaCO_3 , leading to cave formation and conduit enlargement in the coastal carbonate aquifer (Back et al., 1986; Gulley et al., 2016; Mylroie and Carew, 1990; Smart et al., 2006)” in lines 12–14 on page 3 in the revised manuscript.

Thus, in our opinion, there is no need of a change to the manuscript due to the issue raised by the referee.

Page 6, line 10 – *change to tourist.*

Response: Corrected as suggested.

Page 6 – line 23. *The hysteresis observed in pH is clearly shown in the graphs, but I hesitate to conclude that the authors can interpret those results to indicate pH values should be lower in the saline water. Their pH profiles show pH values of ~6.8. Lower pH values of 6.6 seem unlikely considering the buffering capacity of seawater (Ben-Yaakov, 1973)*

Response: We measured lower pH Values of ~6.6 at El Zapote in June 2018 with a new sensor which supports the interpretation that the pH values in the halocline are likely to be lower than shown in the manuscript (shown in Fig. S2). Such low pH values (down to 6.0) were also reported for other deep stratified cenotes of the YP with dominant sulfate reduction like cenote Xcolac (Sacki et. al, 2002). Additionally, the geochemical conditions in the water body of the cenote, especially in the halocline are different to that of porewater in sea sediments.

Therefore, we are convinced that the presented pH values in the revised manuscript are reliable.

Page 6 – line 33-34 – how were the glass vials used to sample water underwater? Were the vials empty and then opened (pressure makes this unlikely). Were the vials pre-filled with water and then purged with breathing gas? If so, how many purges were used?

Response: The vials were transported open and filled with water in bags by the divers. At the desired sampling depth, the divers replaced the water inside the vials with the surrounding water at each water depth by shaky motions and sealed the vials. We addressed this issue by adding this information to the method section in lines 32–34 on page 6 as follows: “The containers were carried open and water-filled by the divers. At the desired sample depth, the water in the containers was exchanged by surrounding water via shaky motions, sealed underwater and the water depth was noted for each sample.”

Page 23 – Fig 9 - DOC appears to spike in the turbid layer and no₃ concentrations are high in freshwater above the halocline. Oxidation of organic material in sediments is important, but the data in figure 3 suggest that organic matter accumulation at the density interface in the turbid zone is also important. Is there any way to include these processes in Fig 9 and in the text referencing Fig 9?

Response: We agree that the organic matter accumulation at the density boundary could play a role in the biogeochemical cycle. The DOC spike in the turbid layer, however, is not indicative for accumulating organic matter as DOC could also reflect metabolites in the complex chemolithotrophic biogeochemical cycle within the redoxcline or turbid layer. Thus, we are not able to determine the origin of DOC and tie it to a specific process in the turbid layer, e.g. accumulation of fine-grained organic matter at the density boundary that could be metabolized by heterotrophic bacteria. Nevertheless, we tried to represent this in Fig. 9 by the top box “minor aerobic and anaerobic heterotrophy” to suggest the possibility of these processes. However, we did not include these processes in the text referencing Fig. 9 as we intend to focus the discussion around Fig. 9 only on the processes involved in Hells Bells formation, hence the processes that lead to the observed alkaline pH-shift.

Page 25, line 21 – replace “wide vertical zone” with “thick vertical zone” to make it more clear that the reference is to the vertical dimension.

Response: Corrected as suggested.

Page 26, entire section on oscillation of the halocline - I think the authors can reduce the length of this entire section by simply stating that the elevation of the halocline oscillates over multiple timescales in response to droughts, hurricanes, and annual tidal fluctuations that are superimposed upon on a longer term increase in sea level. This would allow them to dodge the complicated (and not particularly well-explained) hydraulics associated with recharge elevation of the halocline (which really doesn't add much besides length to this paper). If the authors feel this is super important, they should follow it up with a second paper with the data they mention having collected.

Response: We agree that some parts of this section were not well explained and deleted these parts of the section as suggested by the referee. We also followed the referee's suggestion by adding the following sentence in lines 30–31 on page 25 in the revised manuscript “Therefore, the halocline elevation can vary on multiple timescales in response

to droughts, recharge events and annual tidal fluctuations that are superimposed upon on a longer term sea level change.” We then go on and briefly explain the effects of droughts and recharge events on the halocline elevation. After our opinion, the discussion on the variable halocline elevation is essential for understanding the mechanism of Hells Bells formation. Thus, we intended to make this section shorter and more focused to account for the referee’s concerns about the manuscript length.

Page 30, Section 4.4 – In the rebuttal, the authors indicate that Hell’s Bells are also found in caves with light, so the requirement for a lightless environment here is confusing. Further, I think the authors should specifically mention the requirement for a thick mixing zone. I would then go on to explain how other mixing zones that have been studied in the Yucatan differ from these requirements.

Response: We followed the referee’s suggestion by specifically mentioning a thick halocline as a prerequisite for Hells Bells formation and deleting “lightless environment”. We changed the according sentences in lines 10–11 on page 30 in the revised manuscript to “A meromictic stagnant water body indicated by a thick halocline is needed that allows for the formation of a redoxcline in which anaerobic chemolithoautotrophy prevails.” and in the abstract in line 32 on page 1 to “Finally, we infer that highly stagnant conditions, i.e. a thick halocline, a low-light environment and sufficient input of organic material into a deep meromictic cenote are apparent prerequisites for Hells Bells formation.”

In order to keep the manuscript focused we do not wish to include further comparisons concerning halocline thicknesses of other cenotes of the YP. We will address this particular topic in our next manuscript where we will compare several cenotes with Hells Bells to cenote Angelita, which is devoid of Hells Bells.

Author’s response to the comments of anonymous referee #2:

Referee:

General comment:

I thank the authors for their detailed response to my comments.

My major concern (and obviously of the other reviewer as well) was the hypothesis of the authors that the growth of the Hells Bells is controlled by the depth of the halocline and the redoxcline, which, in turn, are related to recharge. Based on this hypothesis, the authors suggested episodic rather than continuous growth of the Hells Bells (at different depth) and even proposed a relationship with extreme events, such as hurricanes.

In my comment, I stated: “Considering the enormous size of at least some of the bells, it is hard to believe that these should have developed due to seasonal or episodic changes in the depth of the halocline. I would rather believe that this requires a long-term shift in the depth of the halocline, for instance over several thousand years during the Holocene. I had a quick look at the previous paper of the same group (Stinnesbeck et al., 2017b), which presented a few U-series data and reported growth rates of ca. 10-100 $\mu\text{m/a}$. In case of such slow growth rates, it is hard to believe that a short-term decrease in the depth of the halocline due to a recharge event would have a visible effect. In contrast, growth of a really large bell, requires slow and

progressive changes in the (mean) depth of the halocline. The U-series ages ranging from a few hundred to a few thousand years reported by the previous paper, actually seem to confirm this.”

I still think that the only way to test this hypothesis would be systematic U-series dating of several bells from various depths, which requires an enormous number of U-series ages. As far as I understand, this work is currently in progress, and the authors do not want to include these data in the current MS. This is OK, in particular in times where each PhD thesis must consist of several papers. It is still a bit unfortunate, however, for the reader of the current MS because the authors’ response to the reviews makes clear that their interpretations are – at least partly – based on these additional data.

In their revised MS, the authors estimate the potential growth rates of the Hells Bells (I very much appreciate that) and obtain results in agreement with their preliminary U-series data. This shows that the potential growth rates of the bells are in the range of a few hundred $\mu\text{m/a}$. In addition, they show data from their loggers suggesting changes in the range of 10-20 cm to recharge events. Finally, they state in their reply that “... The main argument why we did not consider droughts as a mechanism of halocline elevation is U/TH age-dating on Hells Bells specimens of different water depths (MS currently in preparation) show approximately identical young ages (~ 150 a) at the lowermost crystal tips (1-2 mm) of the Bells. There is even a weak trend of the youngest samples found in the lowest water depths and the oldest samples found in greater water depths. This makes droughts or prolonged periods of time with an elevated halocline as the sole mechanism for the elevation of the halocline unlikely because this should be reflected in an age-zonation of the Hells Bells.”

Considering all these points, I tend to agree that short-term recharge events may have the potential to change the depth of the halocline (even if not in the range of several meters) and result in episodic growth of Hells Bells at different depths. Even if the growth rates were much lower, this could still eventually result in large bells because you have a lot of time (thousands of years). This is comparable to a speleothem that is fed by a very slow and maybe episodic drip rate or only during a specific season of the year (e.g., winter). Growth is episodic, but you may still receive a large speleothem after sufficient time, which appears to have been grown continuously.

In summary, I tend to generally agree with the authors’ hypothesis, now that I have seen the additional data. However, there are two important points, which should be added to the MS prior to publication to make the suggested process clear to reader:

1. Due to the slow growth rates of $\mu\text{m/a}$, it will not be possible to reconstruct the detailed episodic growth history of a single bell on the seasonal or even annual time scale by U-series dating. If the water level and the depth of the halocline fluctuates on a daily, weekly, seasonal or even annual time-scale, you will always have growth at different depths in the cenote throughout the year. Thus, a reconstruction of individual extreme events (e.g., hurricanes) by precise U-series dating will not be possible using the Hells Bells. It is, thus, misleading to state that extreme events, such as hurricanes, can be recorded by the bells. Therefore, I suggest to delete the reference to hurricanes.

Response: We did not intend to suggest that individual hurricanes could be identified in the Hells Bells speleothem records and apologize if it seemed so in the manuscript. We therefore followed the referee’s suggestion and deleted

the references to hurricanes from the abstract, conclusions and most parts of the manuscript and minimized referencing hurricanes to a minimum of two text passages in the revised manuscript.

2. *As far as I understand, growth of the bells in the cenote is restricted to a relatively large range in depth of ca. 10 m. This seems to be a lot compared to the fluctuations of 10-20 cm observed in the logger data (Fig. 1 in the reply of the authors).*

5 *Thus, larger changes in depth may be related to long-term processes (centennial to millennial scale) and minor changes to short-term events. This should be made clear in the text even if the logger data are not shown.*

Response: Right, so far we only recorded recharge-driven halocline fluctuations of 10–20 cm. However, there was no major recharge event since we started logging. It is likely that a heavy recharge event will lead to significantly higher fluctuations of the halocline.

10 We accounted for the referee's concerns by altering the respective text passage in lines 11–12 on page 28 to "This range could solely depend on the hydraulic conditions, e.g. Hells Bells formation reflecting maximum and minimum elevations of the halocline as a result of droughts, recharge events and long-term sea level changes." With these changes and the changes that we introduced following the comments of anonymous referee #1, the matter of halocline elevation is now presented in a more general way in the revised manuscript. This means that
15 instead of focusing on the recharge-driven halocline elevation as it was the case in the original manuscript, several causes for a halocline elevation on variable time-scales are now always discussed in the revised manuscript.

In summary, I recommend to accept the MS after the suggested changes have been made.

References:

20 Socki, R. A., Perry, E. C. and Romanek, C. S.: Stable isotope systematics of two cenotes from the northern Yucatan Peninsula, Mexico, *Limnol. Oceanogr.*, 47(6), 1808–1818, doi:10.4319/lo.2002.47.6.1808, 2002.

List of relevant changes made in the revised manuscript

1. Several text passages were deleted in section "4.3.2. The role of halocline elevation in Hells Bells formation" in order
25 to meet the concerns about the length of the MS by anonymous referee #1.
2. For the same reason another text passage of this section concerning the time scale of frequent recharge-driven halocline elevation was deleted.
3. Several minor changes were made to the MS following the referee's suggestions. Details are given in the author's
30 responses to the referee comments.

All changes to the revised manuscript can be tracked in the marked up version of the revised manuscript.

Subaqueous speleothems (Hells Bells) formed by the interplay of pelagic redoxcline biogeochemistry and specific hydraulic conditions in the El Zapote sinkhole, Yucatán Peninsula, Mexico

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Abstract

Unique bell-shaped underwater speleothems were recently reported from the deep (~55 m) meromictic El Zapote sinkhole (cenote) on the Yucatán Peninsula, Mexico. The local diving community has termed these speleothems as Hells Bells because of their shape and appearance in a dark environment in ~28–38 m water depth above a sulfidic halocline. It was also suggested that Hells Bells form under water, yet the mystery of their formation remained unresolved. Therefore, we conducted detailed hydrogeochemical and geochemical analyses of the water column and Hells Bells speleothems including stable carbon isotopes. Based on the comprehensive results presented in this study we deduce that both, biogeochemical processes in the pelagic redoxcline and a dynamic halocline elevation of El Zapote cenote, are essential for Hells Bells formation. Hells Bells most likely form in the redoxcline, a narrow 1–2 m thick water layer immediately above the halocline where a pelagic chemolithoautotrophic microbial community thrives from the upward diffusion of reduced carbon, nitrogen and sulfur species released from organic matter degradation in organic-rich debris. We hypothesize that chemolithoautotrophy, in particular the proton consuming nitrate-driven anaerobic sulfide oxidation, favors calcite precipitation in the redoxcline and hence Hells Bells formation. A dynamic elevation of the halocline as a hydraulic response to droughts, annual tidal variability and recharge events ,e.g. hurricanes, is further discussed, which might explain the shape of Hells Bells as well as their occurrence over a range of 10 m water depth. Finally, we infer that highly stagnant conditions, i.e. a thick halocline, a low-light environment and sufficient input of organic material into a deep meromictic cenote are

apparent prerequisites for Hells Bells formation. This might explain their exclusivity to only a few cenotes in a restricted area of the northeastern Yucatán Peninsula.

1. Introduction

Speleothems, such as stalactites or dripstones, result from physicochemical processes under subaerial conditions in a cave atmosphere. Calcite usually precipitates due to CO₂-degassing and evaporation of water enriched in dissolved carbonate dripping into the cave. ~~Normally, stalactites rejuvenate and form a tip at the lower end from which the water drips to the cave floor. Nevertheless, in~~ recent years, however, researchers have identified a small group of ~~stalactites-speleothems~~ that appear to have calcified underwater. For these formations, interactions between physicochemical and biological calcite precipitation processes are interpreted (Barton and Northup, 2007; Bontognali et al., 2016; Gradzinski et al., 2012; Guido et al., 2013; Holmes et al., 2001; Jones et al., 2008, 2012; Macalady et al., 2007; Macintyre, 1984; Melim et al., 2001; Queen and Melim, 2006; Tredici et al., 2018). We recently presented a spectacular example for these subaqueous speleothems termed Hells Bells from El Zapote sinkhole about 26 km west of Puerto Morelos on the Yucatán Península of southern Mexico (Fig. 1) (Stinnesbeck et al., 2017b).

These bell-shaped structures consist of calcite and reach lengths of up to two meters. Hells Bells conically expand downward with strictly horizontal ring-like concentric swellings and neckings on the surface (Fig. 2). Apparently, they form in a lightless environment in freshwater above the anoxic and sulfidic halocline (Stinnesbeck et al., 2017b). Because of these ~~toxic~~ environmental conditions in complete darkness, the local diving community has termed the El Zapote speleothem formations as Hells Bells. They grow from the cavern ceiling and wall. Additionally, small individuals also cover a tree that has fallen into the sinkhole around ~3.5 cal kyr BP, which indicates that Hells Bells must have formed during the Holocene to at least historical times, and thus at periods when the deep sections of the cenote had already been submerged for thousands of years (Stinnesbeck et al., 2017b). Thus, the conditions for the formation of the biggest underwater speleothems worldwide must have existed consistently throughout the past thousands of years at El Zapote cenote.

The internal structure of Hells Bells is characterized by laminar fabrics of alternating units of elongated dogtooth spar calcite and microcrystalline spar calcite. Microspar layers and corroded lobes of dogtooth spar crystals indicate either discontinuous growth of Hells Bells and/or intermittent dissolution. Phylogenetic analyses of Hells Bells speleothem surfaces from specimens of 33 and 34 m water depth indicate that microorganisms inhabiting the Hells Bells potentially support a full nitrogen-cycle and autotrophic growth (Stinnesbeck et al., 2017b). The growth of Hells Bells may thus be mediated by specific physical and biogeochemical conditions above and in the halocline, while formation of Hells Bells was likely restricted to the lowermost part of the freshwater body. However, due to the limited available data including geochemical parameters, the suggested processes for Hells Bells formation were regarded as highly speculative.

Therefore, in this study we conducted detailed geochemical analysis including stable carbon isotopes of the water body and Hells Bells speleothems of El Zapote cenote. Based on the results we present a hypothesis on the subaqueous growth of these

exceptional structures. We deduce that both, biogeochemical processes in the pelagic redoxcline and a dynamic halocline elevation of El Zapote cenote, are essential for Hells Bells formation.

1.1 Geological background

The northeastern Yucatán Peninsula (~~YP~~) consists of horizontally layered shallow-water carbonates of Mio-, Plio- and Pleistocene ages (Lefticariu et al., 2006; Weidie, 1985) and probably hosts the largest network of underwater caves in the world. The Mexican state of Quintana Roo alone counts more than 370 underwater caves ~~with, with a total estimated length of >7000 km and~~ a confirmed total length of ~1460 km and individual cave systems reaching up to >350 km in length (QRSS, 2018). These ~~karst~~ cave systems developed predominantly by the interaction of glacioeustasy, littoral processes and mixing-zone hydrology during glacial periods of the Pleistocene (Smart et al., 2006; Weidie, 1985). Precipitation rapidly infiltrates through the porous limestone into the underlying coastal aquifer consisting of a meteoric water mass, the fresh water lens above a ~~marine-saline~~ water mass intruding from the coast (e.g. Kovacs et al., 2017a). The thickness of the freshwater lens varies between ~10–100 m and is generally thinner towards the coast (Beddows et al., 2002), while seawater intrudes up to 100 km inland (Beddows et al., 2007). The halocline separates the meteoric and marine water bodies and is usually characterized by undersaturation with respect to CaCO₃, leading to cave formation and conduit enlargement in the coastal carbonate aquifer (Back et al., 1986; Gulley et al., 2016; Mylroie and Carew, 1990; Smart et al., 2006). The depth of the halocline increases with distance from the coast (Bauer-Gottwein et al., 2011); areas closer to the coast show a higher salinity of the freshwater lens than inland areas (Kovacs et al., 2017b). The position of the halocline is also dependent on global sea level and the thickness of the freshwater lens. Hydraulic gradients are generally very low with values of 1–10 cm km⁻¹ (Bauer-Gottwein et al., 2011 and references therein). Although Moore et al. (1992) and Stoessell et al. (1993) report that the thickness of the freshwater lens does not vary significantly between seasons or on a yearly basis, local and short-termed variations are possible and were reported by Escolero et al. (2007), who documented a significant halocline elevation of up to 17.5 m in between two measurements in the years 2000 and 2003.

Sinkholes (locally called *cenotes*) were formed by dissolution and collapse of the carbonate rock. They are common throughout the ~~Yucatán Peninsula~~YP, connecting the subterranean cave system with the surface (Bauer-Gottwein et al., 2011). For more detailed information about the formation and occurrence of cenotes on the ~~Yucatán Peninsula~~YP we refer the readers to Torres-Talamante et al. (2011) and Schmitter-Soto et al. (2002).

1.2 El Zapote cenote

El Zapote cenote is located 26 km west of Puerto Morelos on the ~~Yucatán Peninsula~~YP of southeast Mexico (20°51'27.78" N 87°07'35.93" W; Fig. 1). In cross-section the cenote is bottle-shaped with a deep vertical water-filled shaft that opens at 28 m water depth to a wide cavern of 60 to >100 m in diameter that reaches to about 54 m water depth with a 20 m high debris mound in the center (Fig. 3a). A fallen tree stands on top of the debris mound and small Hells Bells cover the stem. There are no apparent passages or conduits that connect El Zapote cenote to a cave system. Additional details on El Zapote

cenote are given in Stinnesbeck et al. (2017b) and in Stinnesbeck et al. (2017a) who described the new genus and species of a giant ground sloth, *Xibalbaonyx oviceps*, from an individual that was found on the floor of El Zapote cenote.

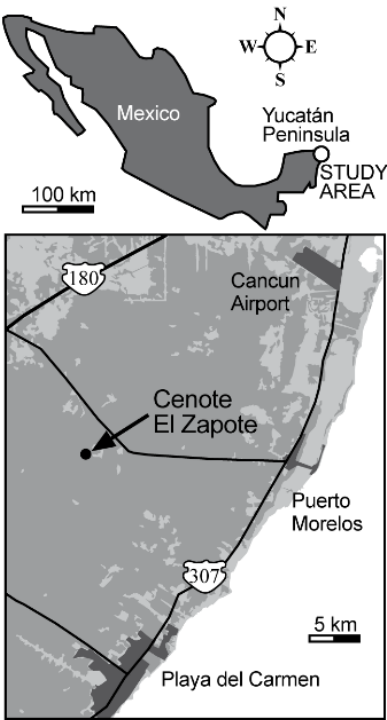


Fig. 1: Location of the studied El Zapote cenote with respect to Mexico and the Yucatán Peninsula.

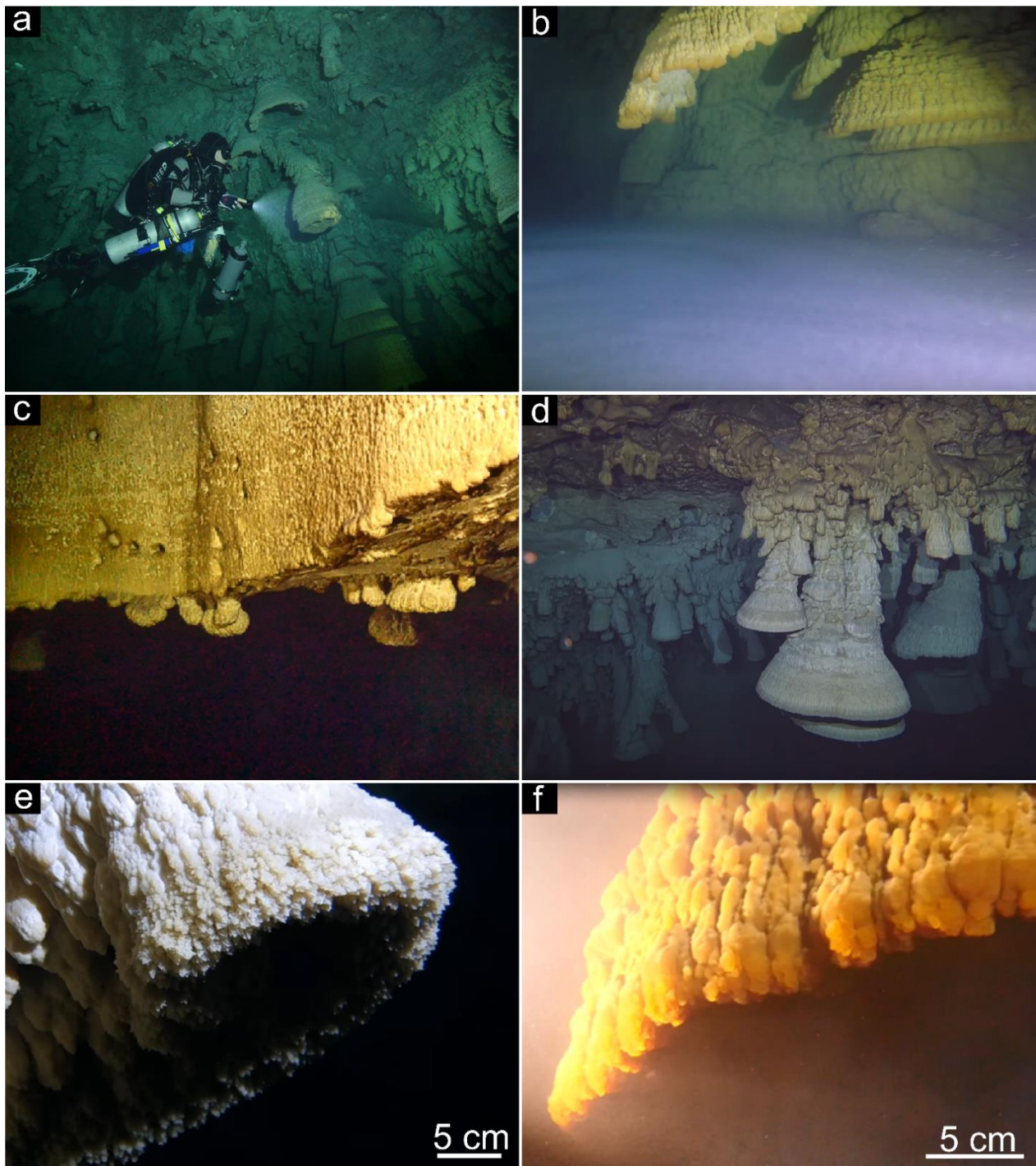


Fig. 2: Technical diver in El Zapote cenote during a sample collecting dive carrying the Niskin bottle, sample containers and the multiparameter sonde attached to the sidemount gas bottle (a). Turbid layer immediately above the halocline forming a distinct horizontal white cloud at around 36 m water depth (b). Transition of cenote shaft to the open dome-shaped cavern at 28 m water depth (c), where the vertical wall of the cenote shaft is devoid of Hells Bells speleothems (upper part of picture c), whereas small

specimens of Hells Bells grow down from the horizontal ceiling below (lower part of c). Brown-colored Mn-oxide coatings on host rock carbonates and Hells Bells speleothems reach down from 28 m water depth to around 30 m water depth at the transition of the cenote shaft to the wide dome-like cavern (c and d). Below around 30 m water depth, Hells Bells speleothem and host rock carbonate surfaces are devoid of brownish Mn-oxide coatings. They are white to light-grey colored revealing a distinct horizontal boundary (d). Close-up shot of the lowermost calcite rim of a Hells Bell speleothem at around 32–35 m water depth showing mm-sized calcite crystals (e).

2. Methods

2.1 Sampling

Sampling at El Zapote Cenote was carried out between December 10th and 15th 2017. Water samples were taken early in the morning prior to any tourist diving group activities to ensure sampling of an undisturbed water column. Water sample recovery and the recording of the in-situ parameters were carried out with a winch from the surface down to the top of the debris mound (0–36 m) and by technical divers from the top of the debris mound down to the cenote floor following the slope of the debris mound (Fig. 2a).

In-situ parameters pH (± 0.1), eH (± 20 mV), dissolved oxygen (± 0.1 mg l⁻¹, detection limit 0.2 mg/l), electrical conductivity (± 0.05 % of value), temperature (± 0.01 °C) and turbidity (± 2 % of value) were determined with a multiparameter water sonde EXO-1 (Xylem Analytics, Norway). All parameters, including water depth via pressure measurement, were internally logged by the sonde (± 0.04 m). Water depths were corrected to the ambient air pressure of the respective day of sampling. In order to reach the greatest possible water depth, a total of four winch-operated profiles were run within 2 days, with laterally shifting starting points of the profile at the surface. In order to complete the measurement in the whole water column, technical divers carried the EXO-1 Probe with them during sampling. Due to increasing sulfide concentrations in water depths below the turbid layer and interaction of sulfide with the Ag/Cl pH electrode, a shift of pH of up to 0.2 pH units towards higher values was observed when comparing the pH logs of the way down with the pH logs of the way up (Fig. S1). This shift is dependent on the exposure time of the electrode and the respective sulfide concentrations and could not be quantified nor corrected for. However, the sensor recovers to initial pH values after a certain time in non-sulfide water. Therefore, the pH values presented in this study are representative for the water column from 0 to 37 m water depth and are overestimated in water depths from 37–50 m where the actual absolute pH values are most likely lower, i.e. more acidic. Repeated measurement with a new sonde of the same type in June 2018 confirmed this assumption and showed lower pH values below 37 m water depth (Fig. S2).

Water samples from 15–35 m water depth were retrieved using a winch and a 5 L polyethylene Niskin bottle (Hydrobios, Kiel, Germany). Sampling depths represent the center of the 0.6 m tall sampling bottle and were determined by cable length with a depth counter attached to the winch. Water samples from 35.2–45 m water depth were retrieved by technical divers (Fig. 2a). Water samples collected by the technical divers were taken with 120 ml PE-containers. The containers were carried open and water-filled by the divers. At the desired sample depth, the water in the containers was exchanged by surrounding water via shaky motions, sealed underwater and the water depth was noted for each sample. Water samples for the analysis

of dissolved gases (CO_2 , CH_4) were taken in 24 ml glass vials and sealed underwater at the respective depth (four samples at each depth level). The EXO-1 sonde was attached to a side mounted compressed air bottle pointing towards the front of the technical diver in order to record the in-situ parameters of each water sample (Fig. 2a). The depth of the water samples taken by technical divers was corrected to the depth of the attached logging device (EXO-1). For four samples between 35 and 37 m, depth was interpreted from the increase of sodium and chloride contents correlated to the electrical conductivity in this interval.

Water samples were treated on-site immediately after the water samples were retrieved. Samples for determination of dissolved ions were taken with 20 ml sterile polypropylene syringes and then filtered through a cellulose acetate filter (0.45 μm). Samples for cation determination were acidified with 50 μl of 65 % HNO_3 (A.G.) to adjust a pH <2; they were stored in 15 ml Falcon polypropylene centrifuge tubes. Samples for anion determination were taken equivalent, but not acidified, and stored cool in 15 ml Falcon polypropylene centrifuge tubes. Samples for the determination of dissolved inorganic carbon (DIC) and dissolved organic carbon (DOC) were filtered through a cellulose acetate filter (0.45 μm), stored in 24 ml glass vials and sealed gas-tight. Samples for the determination of content and isotopic ratios of the dissolved gases CH_4 and CO_2 were filled in 24 ml glass vials; subsequently 100 μl 60 % HgCl_2 solution was added via a syringe pierced through the septum to sterilize the samples.

A large volume sample (5 l) of the turbid layer water at around 36 m water depth was taken with a Niskin bottle by technical divers and subsequently filtered through a 0.45 μm cellulose acetate filter with a vacuum pump. The filter was air-dried; back in the laboratory a small piece of the filter was coated with carbon for subsequent secondary electron (SE) imaging and analyses.

Sulfide and nitrite were determined on-site by photometric analysis with a Photometer (Hach Lange DR200) using Merck Spectroquant spectrometric methods

Technical divers collected several Hells Bells samples grown on the tree trunk from 7 water depth levels between 32.7 and 37.3 m. To obtain the youngest part of individual Hells Bells growth, samples were first studied under the microscope. Only samples with apparently fresh, well accentuated crystal tips were chosen for analysis geochemical and stable isotope analysis.

2.2 Analytical measurements

2.2.1 Major and trace element analysis

Major cation concentrations (Ca, Mg, Sr, Ba, K, Na, Si) of water samples were determined by optical emission spectroscopy with an Agilent 720 ICP-OES. Quality control was performed using the reference materials SPS-SW1, SPS-SW2 and TMDA 70.2. Recovery rates were in the range of 97–104 % for the analyzed elements. Measurement precision for each element was <2 % (RSD, n=3).

Concentrations of anions Cl^- , SO_4^{2-} and NO_3^- were determined with ion-chromatography (Dionex ICS-1100) with a RSD of <3 % derived from long-term repeated analysis of reference material SPS-WW1 NUTR. The concentrations of DOC were determined with a Total Carbon Analyzer (Shimadzu TOC-CPH) with a RSD of <2 % derived from repeated analysis of an in-house standard water.

5 Orthophosphate was determined by the photometric molybdenum phenyl blue method on 880 nm light extinction with a UV/VIS photometer (Specord 50, Analytic Jena).

Trace element concentrations (Fe, Mn, Mo, P, S, U) of selected water samples were determined by HR-ICP-MS (Thermo Finnigan Element 2). Analyses were normalized by an internal Indium-standard. Calibration solutions were prepared with the MERCK VI Multi Element Standard. The recovery rates of SLRS5 reference material were 95 % (Fe), 93 % (Mn), 82 %

10 (Mo) and 133 % (U) with respect to the referenced values and P and S were within the range of reported uncertified values. The precision was <3.7 % (RSD) derived from repeated (n=5) measurements of the reference material in the measurement run.

Around 3 mg of the powdered speleothem samples were digested in 2ml 10% HNO_3 for major and trace cation analyses. Subsequently, concentrations of Ca, Mg, Sr, Ba, P, S, Fe and Mn of the aliquots were determined by ICP-OES. Quality control of the measurement was performed using reference materials SPS-SW1 and SPS-SW2 with recovery rates ranging from 99 to 111 % for the analyzed elements. Quality control for digestion of the carbonate material was performed with limestone reference material ECRM 752-1 with recovery rates between 106–110 % for the elements Ca, Mg, Ba, Sr and Mn and 82 % for the element Fe.

20 Calcite saturation and HS^- activity was calculated with Phreeqc (Parkhurst and Appelo, 1999) using phreeqc.dat. The diffusion J was calculated with the first Fick's law with diffusion coefficients of D_{O_2} , $\text{D}_{\text{NO}_3^-}$, $\text{D}_{\text{HCO}_3^-}$ and D_{HS^-} of 2.1, 1.9, 1.2 and $1.8 \cdot 10^{-9} \text{ m}^2 \text{ s}^{-1}$, respectively taken from phreeqc.dat (Parkhurst and Appelo, 1999).

2.2.2 Stable carbon isotope and concentration measurements of CH_4 and CO_2

For the determination of dissolved gases, a 5 ml headspace with nitrogen gas (N_2 99.999 %) was created in each of the four samples of the respective water depth. Samples were taken for the analysis of dissolved gases at ambient laboratory temperatures of 23 °C. After equilibration (~24h), the headspace of the four samples was transferred and combined in one 12 ml evacuated exetainer vial. To ensure a pressureless transfer of the gas phase from the headspace to the exetainer, a brine solution of 200 g l^{-1} NaCl was introduced at the bottom of the vial and the gas-phase was simultaneously removed and subsequently transferred to evacuated exetainer vials. Concentrations CH_4 and CO_2 in the gas samples were measured as follows: Headspace samples (50 μl) were injected in a flow of 1 mL min^{-1} of helium with a split ratio of 5:1 to a Shin Carbon ST column (80/100 mesh, 2m x 0.53mm i.d., Restek Corporation) quantified by a gas chromatograph (GC-2010 Plus, Shimadzu Corporation, Kyoto, Japan) coupled to a Barrier Ionization Discharge (BID) detector (BID-2010 Plus, Shimadzu Corporation, Kyoto, Japan). The GC oven was initially held at 30°C for 1 min and then ramped at 10°C/min to 200°C. Quantification of CH_4 and CO_2 was carried out by comparison of the integrals of the peaks eluting at the same retention time

as that of the authentic standard with calibration curves. The dissolved concentrations of CH₄ in the water were then calculated from the measured mixing ratio using Henry's law (Wiesenburg and Guinasso, 1979) and solubility coefficients for CH₄ according to Weiss (1974) and Yamamoto et al. (1976).

Stable carbon isotope ratios of CO₂ ($\delta^{13}\text{C}$ -CO₂ values) were analyzed by gas chromatography stable isotope ratio mass spectrometry (GC-IRMS) by a HP 6890N gas chromatograph, coupled to a 253 Plus™ isotope ratio mass spectrometer (ThermoQuest Finnigan, Bremen, Germany) with average analytical uncertainties of 0.2‰ for $\delta^{13}\text{C}$ -CO₂ values. 2 σ uncertainties were derived from 5 replicates. All $^{13}\text{C}/^{12}\text{C}$ isotope ratios are expressed in the conventional δ notation in per mil versus VPDB, defined in ~~equation~~ (Eq. (1):

$$\delta^{13}C_{V-PDB} = \left[\left(^{13}\text{C}/^{12}\text{C}_{\text{sample}} \right) / \left(^{13}\text{C}/^{12}\text{C}_{\text{standard}} \right) \right] - 1 \quad (\text{Eq. 1})$$

For details of the $\delta^{13}\text{C}$ -CO₂ measurements by GC-IRMS we would like to refer to previous studies by Keppler et al. (2010) and Laukenmann et al. (2010).

Stable carbon isotope ratios of CH₄ ($\delta^{13}\text{C}$ -CH₄ values) were determined (GC-IRMS). In brief, CH₄ of the sample was trapped on Hayesep D and then transferred to the IRMS system (ThermoFinnigan Delta^{plus} XL, Thermo Finnigan, Bremen, Germany). The working reference gas was carbon dioxide of high purity (carbon dioxide 4.5, Messer Griesheim, Frankfurt, Germany) with a known $\delta^{13}\text{C}$ -CH₄ value of -23.634 ‰ \pm 0.006 ‰ versus V-PDB. All $\delta^{13}\text{C}$ -CH₄ values were corrected using two CH₄ working standards (isometric instruments, Victoria, Canada) and normalized by two-scale anchor calibration according to Paul et al. (2007). The average standard deviation of the analytical measurements was in the range of 0.1 ‰ to 0.3 ‰.

The $\delta^{13}\text{C}$ -HCO₃⁻ values were calculated from the measured $\delta^{13}\text{C}$ -CO₂ of the headspace of the water samples that was generated in the laboratory as equilibrium fractionation at 23° C ($\delta^{13}\text{C}$ -CO₂ + 8.16 ‰ = $\delta^{13}\text{C}$ -HCO₃⁻) after Mook (2000).

For stable carbon isotope analyses of carbonates, approximately 50 μg of powdered speleothem subsamples was analyzed using a ThermoFinnigan MAT253Plus gas source mass spectrometer equipped with a Thermo Fisher Scientific Kiel IV Carbonate Device at Heidelberg University. Values are reported relative to VPDB (Eq. 1) through the analysis of an in-house standard (Solnhofen limestone) calibrated to IAEA-603. The precision of the $\delta^{13}\text{C}$ analyses is better than 0.08 ‰ and 0.06 ‰ (at 1 σ level), respectively.

2.2.3 Optical methods

Hells Bells specimen ZPT 7, described in Stinnesbeck et al. (2017b) was vertically cut in half and thin sections were prepared from one half of the specimen. Photographs of the thin sections were taken with a Keyence VHX-6000.

Polished counterparts of the thin sections and small pieces of Hells Bells were coated with carbon for secondary electron (SE) imaging and energy dispersive X-ray (EDX) analyses. SE-imaging and element mapping was performed with a Leo 440 at 20 kV with a X-Max 80 mm² detector.

3. Results

3.1 Hydrogeochemistry

The water column of the El Zapote cenote is stratified into an oxygenated fresh water body overlying an anoxic transition zone of increasing electrical conductivity (EC), the halocline, and an anoxic salt water body below (Fig. 3 a). Water temperatures vary little between 0 to 30 m water depth (24.37–24.42 °C); a steep increase is identified in a narrow zone from 30–32 m water depth (24.42–24.55 °C), followed by almost invariable temperatures from 32 m water depth (24.55 °C) down to the bottom of the cenote (25.22°C) (Figs. 3a and b). A distinct density boundary, the top of the halocline, is identified at 36.6 m water depth by a steep increase in EC. Sea water-like salinity is reached at around 46 m water depth indicating a thick halocline layer of around 10 m thickness (Figs. 3a and b). Low turbidity readings indicate clear water throughout the water column, except for a ~1.6 m thick layer of increased turbidity immediately above the halocline from 35.0–36.6 m water depth, with a peak of 8.0 FNU detected at 35.8 m water depth (Figs. 3a and b; Table S1). This turbid layer is also easily detected macroscopically in the water column as a white cloudy layer (Fig. 2b) and coincides with a distinct redoxcline from ~35–37 m water depth, in which the redox potential (EH) decreases from ~250 to ~140 mV (Figs. 3a and b). Dissolved oxygen (DO) decreases nearly linear from 30 m to concentrations below detection limit at ~35 m water depth just above the turbid layer. Below, DO was below detection limit (Figs. 3a and b). The pH shows neutral values from 0–30 m water depth and slightly decreases to 6.90 at the top of the turbid layer (Fig. 3a). Within the turbid layer pH values increase to more alkaline values of around 6.94 at 35.8 m water depth. The pH values decrease again below the turbid layer to 6.73 at 40 m and invariably remain at about this value down to 48 m. From there, values increase to about neutral (6.95) close to the cenote bottom at 49 m water depth (Figs. 3a and b).

Concentrations of the major dissolved ions Na^+ , Cl^- , Ca^{2+} , Mg^{2+} and SO_4^{2-} reflect the stratification of the water column in the cenote, with generally low concentrations in the fresh water body from 0–30 m water depth and slightly increasing concentrations from 30 m water depth to the top of the turbid layer at 35 m water depth, a stronger increase within the turbid layer from 35–36.6 m water depth, and an even stronger increase from the top of the halocline at 36.6 m water depth down to the cenote bottom (Fig. 3c and Table S2). Mg/Ca ratios strongly increase from the top of the turbid layer at 35 m water depth downwards, due to higher Mg concentrations in the salt water body (Fig. 3c). Although sulfate concentrations increase downwards from the top of the halocline, a relative decrease of SO_4^{2-} ions is detected, compared to the chemically conservative ion Cl^- , by a decrease in $\text{SO}_4^{2-}/\text{Cl}^-$ within the turbid layer and below in the halocline (Fig. 3c).

chemical parameters determined in the water column between 31–41 m water depth. Na^+ and Cl^- concentrations are only shown in the range of 0–80 mmol l^{-1} to highlight the concentration pattern above and within the halocline.

Concentrations of DIC are about 7.8 mmol l^{-1} in the fresh water body. They increase in the turbid layer and show a peak at 40 m water depth with concentrations increasing to 14.5 mmol l^{-1} ; below, they decrease towards the cenote bottom (Fig. 3c and Table S2). The dissolved organic carbon (DOC) concentrations are low in the fresh water body and show a distinct peak within the turbid layer, coinciding with the peak in turbidity at 35.7 m water depth (Fig. 3c). Below the turbid layer DOC concentrations slightly increase and peak at 39–40 m water depth, decreasing from there towards the cenote bottom (Fig. 3c and Table S2). Nitrate concentrations are $\sim 50 \mu\text{mol l}^{-1}$ in the fresh water unit of the cenote shaft (Table S2). They decrease from 30 m water depth towards the top of the turbid layer and rapidly fall below detectable concentrations within this layer (Fig. 3c). Nitrite peaks in a narrow zone at the top of the turbid layer with concentrations of up to $0.8 \mu\text{mol l}^{-1}$ (Fig. 3c). High total sulfide (S(-II)) concentrations of up to 5.6 mmol l^{-1} were detected in 40 m water depth. Concentrations decrease upwards, fading in the lower part of the turbid layer at 36 m water depth (Fig. 3c). Below the 40 m depth level, S(-II) concentrations decrease to values around 3 mmol l^{-1} down to 45 m water depth (Fig. 3c and Table S2). Concentrations of dissolved CH_4 ($\text{CH}_4(\text{aq})$) are low in the fresh water body with values of about $0.09 \mu\text{mol l}^{-1}$. Methane concentrations increase from the turbid layer at 36 m water depth downwards to values of $25 \mu\text{mol l}^{-1}$ at 39 m water depth (Fig. 3c).

3.1.1 Calcite Saturation

The calculated saturation index (SI) of calcite shows calcite saturation in the fresh water body and the uppermost part of the halocline with values from 0.03–0.07 (Fig. S2). The SI closely follows the pH in the fresh water body revealing a distinct peak of slightly higher values of $\text{SI} = 0.1$ in the turbid layer at ~ 36 m water depth. The water body below the redoxcline is undersaturated with respect to calcite indicating calcite dissolution in the halocline (Fig. S2).

3.1.2 Trace elements

Dissolved iron and manganese concentrations are very low in the fresh water body with concentrations of 0.1 and $0.01 \mu\text{mol l}^{-1}$, respectively, and slightly increase within the turbid layer towards the salt water body, to concentrations of up to 0.47 (Fe) and $0.06 (\text{Mn}) \mu\text{mol l}^{-1}$ (Fig. S3). Phosphate and silica concentrations are invariably low in the fresh water body ($\text{P}_{\text{ortho}} \sim 0.25$ and $\text{Si} \sim 63 \mu\text{mol l}^{-1}$) and increase in the salt water body peaking at 40 m water depth with concentrations up to 10.3 (P_{ortho}) and 275 (Si) $\mu\text{mol l}^{-1}$ (Fig. S3). Uranium content correlates to the redox potential of the water as indicated by uniform contents of $\sim 0.012 \mu\text{mol l}^{-1}$ in the fresh water column and rapidly decreasing values in one order of magnitude in the turbid layer, to $0.0012 \mu\text{mol l}^{-1}$ at 40 m water depth (Fig. 4 and Table S2).

3.1.3 Stable carbon isotopes of DIC and CH_4

The $\delta^{13}\text{C-HCO}_3^-$ values at water depth from 28 to 42 are shown in Figure. 4. The average $\delta^{13}\text{C-HCO}_3^-$ value is -9.8 ‰ in the fresh water body where DIC content is about 8 mmol l^{-1} . In the turbid layer $\delta^{13}\text{C-HCO}_3^-$ values show a distinct peak towards

less negative values of up to -7.9 ‰ at slightly increasing DIC concentrations. Below the turbid layer $\delta^{13}\text{C-HCO}_3^-$ values rapidly decrease towards more negative values of -12.4 ‰ between 39 and 42 m water depth at increasing DIC concentrations (Fig. 4). A rather slight increase of $\delta^{13}\text{C-HCO}_3^-$ values (-11.6 ± 0.7 ‰) is observed towards the cenote bottom at 44 m water depth (Table S3).

- The $\delta^{13}\text{C-CH}_4$ values are shown alongside with the CH_4 concentrations in Figure 5. The pattern of $\delta^{13}\text{C-CH}_4$ within the water column is similar to that of $\delta^{13}\text{C-HCO}_3^-$. In the fresh water body values of $\delta^{13}\text{C-CH}_4$ are approximately constant at about -49 ‰ and CH_4 concentrations are very low, roughly corresponding to that of atmospheric equilibrium ($0.04\text{--}0.09 \mu\text{mol l}^{-1}$). $\delta^{13}\text{C-CH}_4$ increases to -28 ‰ within the turbid layer and again decreases to -61 ‰ below the turbid layer, while CH_4 concentrations increase within and below the turbid layer (Fig. 4).

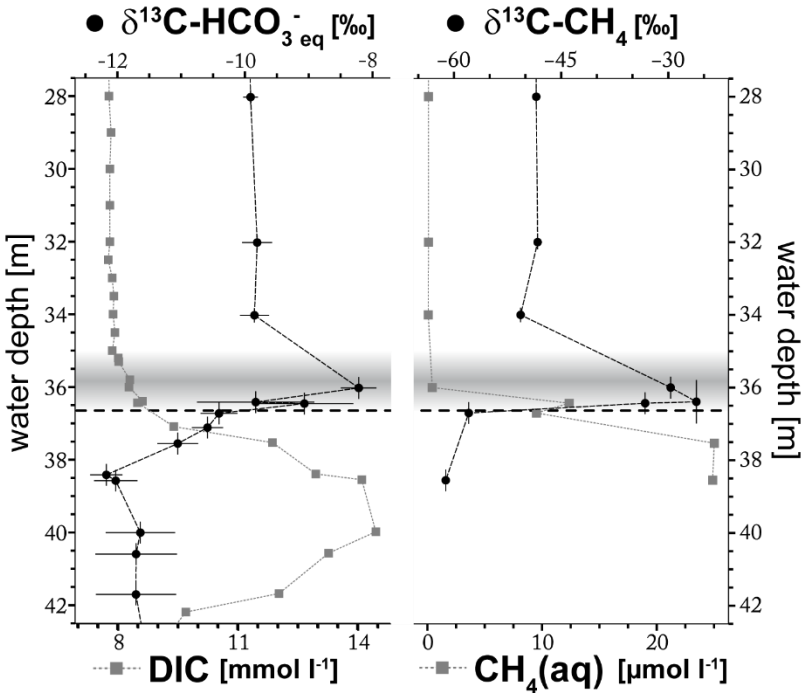


Fig. 4: Stable carbon isotope values $\delta^{13}\text{C-HCO}_3^-$ of the dissolved HCO_3^- in equilibrium with $\delta^{13}\text{C-CO}_2$ values measured in headspace and $\delta^{13}\text{C-CH}_4$ values of water samples alongside the concentrations of DIC and CH_4 of water samples. The grey band represents the turbid layer in 35–36.6 m water depth and the horizontal dashed line indicates the top of the halocline at 36.6 m water depth. Horizontal error bars represent 2σ uncertainties, vertical error bars indicate up to 0.6 m uncertainty of gas samples that were not taken from the sample used to determine chemical parameters (see [section Sect. 2.1](#)).

3.2. Petrography of Hells Bells speleothems

Petrographic characteristics of Hells Bells are shown in Fig. 6. The size of individual crystals of Hells Bells is varying from μm scale to several mm-sized crystals that are easily identified macroscopically. The latter are frequently dominant in the youngest calcite rims at the bottom of Hells Bells from water depths reaching from ~28 to ~35 m (Fig. 2e). Hells Bells from

greater water depths show rounded or globular calcite surfaces at the lowermost margin of the speleothems indicating dissolution (Fig. 2f). SEM-images of the lowermost part of Hells Bell surfaces are shown in Figs. 6a, b and c. The calcite morphology varies from bladed or book-like calcite crystals (Fig. 5 a), dogtooth-like calcite crystals (Fig. 5b) and blocky calcite rhombs (Fig. 5c). In thin sections of the specimen ZPT-7 (Fig. 5 d1) (see also Stinnesbeck et al., 2017b), these crystal morphologies are expressed as rather botryoidal (dog-tooth-like and bladed-shaped) and mosaic calcite phases (blocky calcite rhombs) (Fig. S5). Electron images of the polished counter pieces of the thin section are shown in Fig. 5d2. An element map of Mg shows that botryoidal calcite phases incorporated more Mg (appearing brighter in Fig. 5d3) than the mosaic calcite phases (appearing darker in Fig. 5d3).

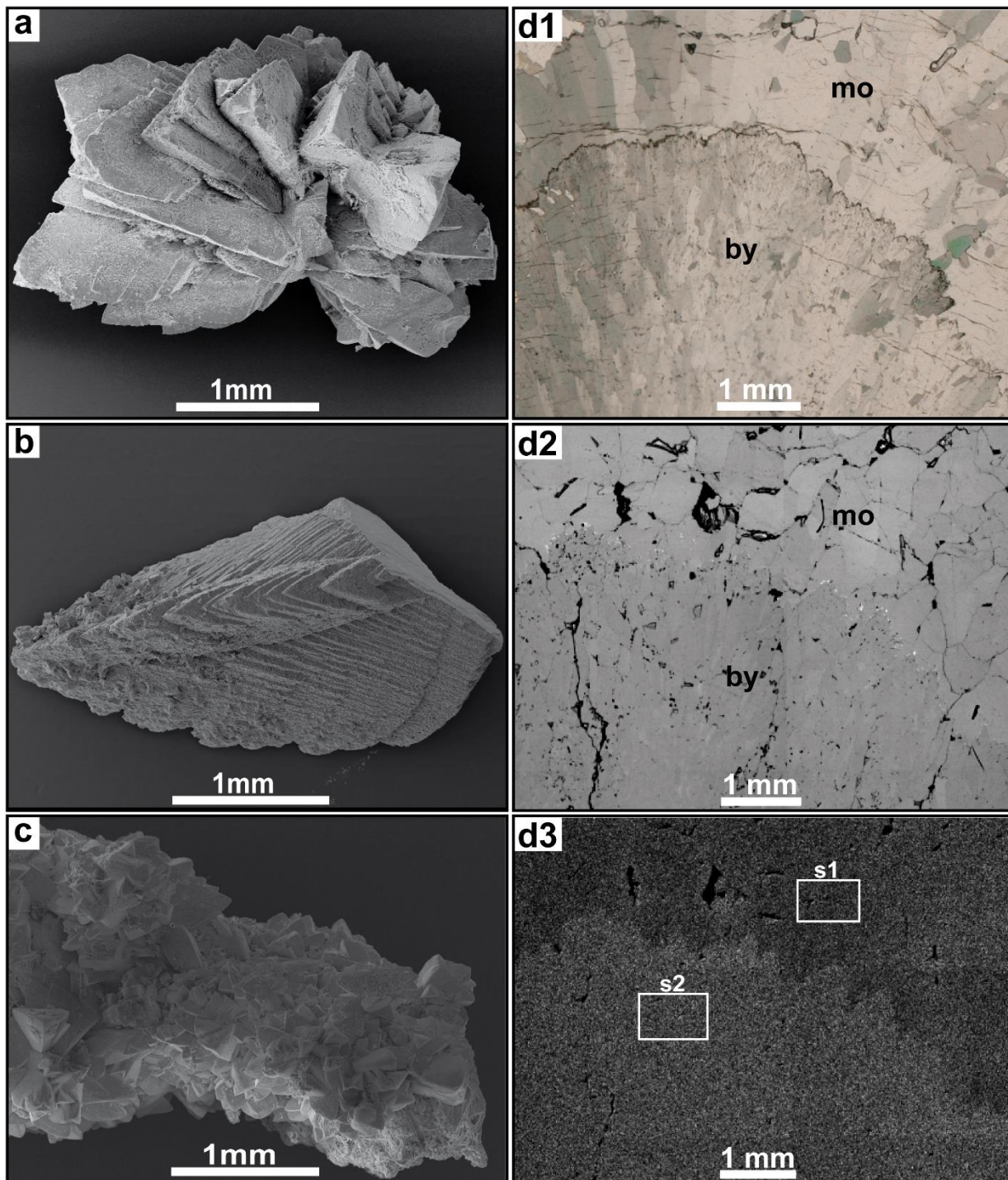


Fig. 5: Petrographic characteristics of Hells Bells speleothems. SE-images of Hells Bell samples Z17-8DC (a), Z17-18J (b), Z17-9J (c) of El Zapote cenote showing bladed (a), dogtooth-like (b) and blocky (c) calcite rhombs. Polarized transmitted light-microscopic images of a thin section from ZPT-7 (d1) (shown in Stinnesbeck et al., 2017b) showing different calcite fabrics of

angular coarse-grained mosaic calcite (mo) and fine grained elongated botryoidal calcite (by). The same detail is shown in the BSE-image of the polished counter slab that corresponds to the thin section (d2). The Mg-element map (d3), where higher abundances of Mg appear brighter, indicates a difference in Mg content between the botryoidal and mosaic calcite phases. The white rectangles represent areas of measured integrated element spectra.

5 3.3 Geochemistry of Hells Bells speleothems

Samples were collected from the lowermost and presumably youngest part of several Hells Bell specimens that grew on a ceiba tree fallen into the El Zapote cenote at about 3.5 cal kyr BP (Stinnesbeck et al., 2017b). They were analyzed for major and trace elements and stable carbon isotopes. The results are given in Table. 1.

3.3.1 Major and trace elements

- 10 The calcite of Hells Bell speleothems revealed no residues after digesting ~3 mg sample in 12 ml dilute 1 M Nitric indicating that Hells Bells calcite is devoid of acid insoluble impurities. The Mg/Ca, Sr/Ca and Ba/Ca ratios show narrow ranges with mean values of $22.5 \pm 2.9 \cdot 10^{-3}$, $38.6 \pm 5.9 \cdot 10^{-5}$ and $1.10 \pm 0.31 \cdot 10^{-5}$, respectively. They are closely related and positively correlate in each sample (Fig. 6a). There is also a trend towards decreasing ratios with increasing water depth of the respective sample (Fig. 6b). Iron and manganese show more variable concentrations with ratios of Fe/Ca and Mn/Ca
15 between $3.0\text{--}11.3 \cdot 10^{-5}$ and $16\text{--}39.3 \cdot 10^{-6}$, respectively. Iron and manganese show a weak positive correlation but no dependency on water depth. The content of sulfur in Hells Bells carbonate is constantly high with concentrations of 0.8–1.0 g kg⁻¹ (Table S4) and mean S/Ca ratios of $2.87 \pm 0.42 \cdot 10^{-3}$, showing no dependency on water depth of the sample (Table 1).

3.3.2 Hells Bells stable carbon isotopes

- Stable carbon isotope values of Hells Bells calcite samples ($\delta^{13}\text{C}_{\text{calcite}}$) from different water depth range from -12.85 to -13.82
20 ‰ with a mean value of -13.37 ± 0.70 ‰ (n=9, Table 1). There is a weak correlation of increasing $\delta^{13}\text{C}_{\text{calcite}}$ values with water depth of the samples (Fig. 6b). Furthermore, $\delta^{13}\text{C}_{\text{calcite}}$ values show a strong negative correlation with Sr/Ca and Ba/Ca with r^2 of 0.82 and 0.89 (Fig. 6a).

- The stable carbon isotope ratio of the HCO_3^- that is in equilibrium with the Hells Bells calcite ($\delta^{13}\text{C}_{\text{-eqHCO}_3^-}$) at 25 °C water temperature ($\delta^{13}\text{C}_{\text{Calcite}} - 0.91$ ‰ = $\delta^{13}\text{C}_{\text{-eqHCO}_3^-}$) was calculated after Mook (2000). The calculated $\delta^{13}\text{C}_{\text{-eqHCO}_3^-}$ is -
25 14.28 ± 0.70 ‰ which is lower than the $\delta^{13}\text{C}\text{-HCO}_3^-$ determined for the water column with a range of -9.1 to -12.3 ‰.

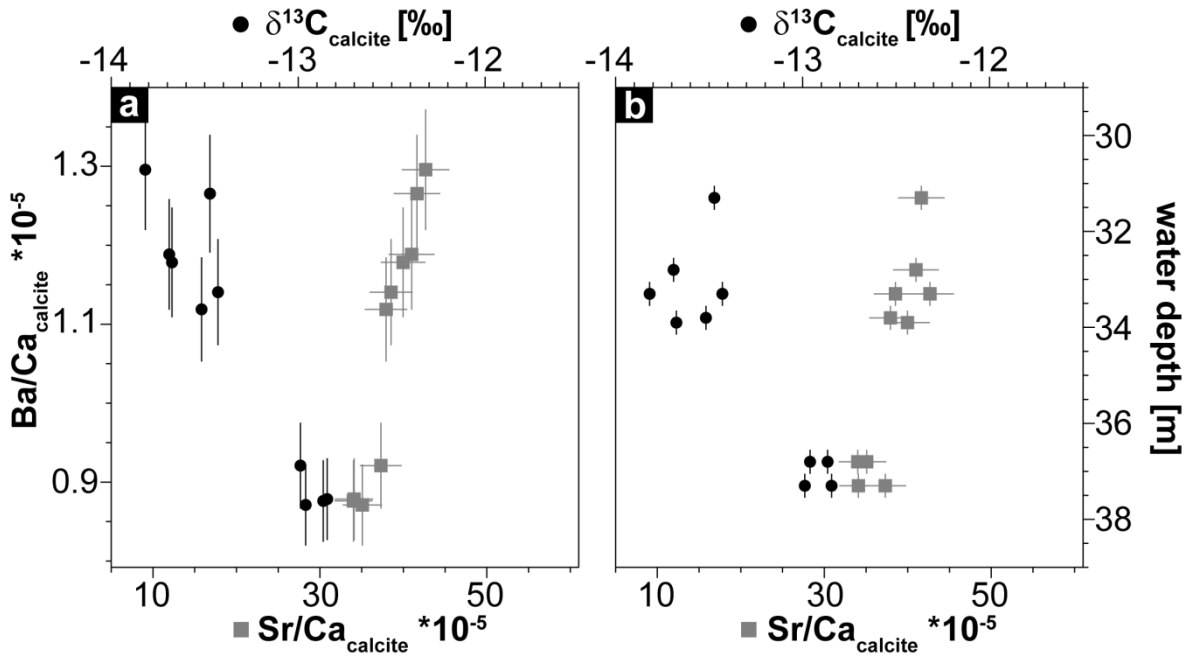


Fig. 6: Geochemical data of Hells Bell speleothems showing a strong correlation between Sr/Ca and Ba/Ca ratios ($r^2 = 0.91$) and between Ba/Ca and $\delta^{13}\text{C}_{\text{calcite}}$ ($r^2 = 0.89$) (a) and a trend of increasing $\delta^{13}\text{C}$ and decreasing Sr/Ca with increasing water depth of the samples (b). Given uncertainties represent 2σ standard deviations and ± 0.25 m is assumed as uncertainty for the water depth of the Hells Bell samples.

3.4 Turbid layer filtrate

Although the turbid layer appears dense in photographs taken during dives, the water sampled from the turbid layer was clear, with no visible turbidity during sample handling. Electron microscopy of the filter, however, reveals that abundant particles were extracted from the turbid layer (Fig. 7a). Particle sizes range between 1–100 μm , but most are in the range of 1–10 μm . They consist of Ca-carbonate crystals (Figs. 7c and c1), globular particles consisting of elemental sulfur (Figs. 7d and d1) and silicate particles of different compositions (Figs. 7e and e1). Also, numerous intact and broken shells of siliceous diatoms were found on the filter. Some calcite crystals incorporated broken parts of silica shells (Figs. 7c1).

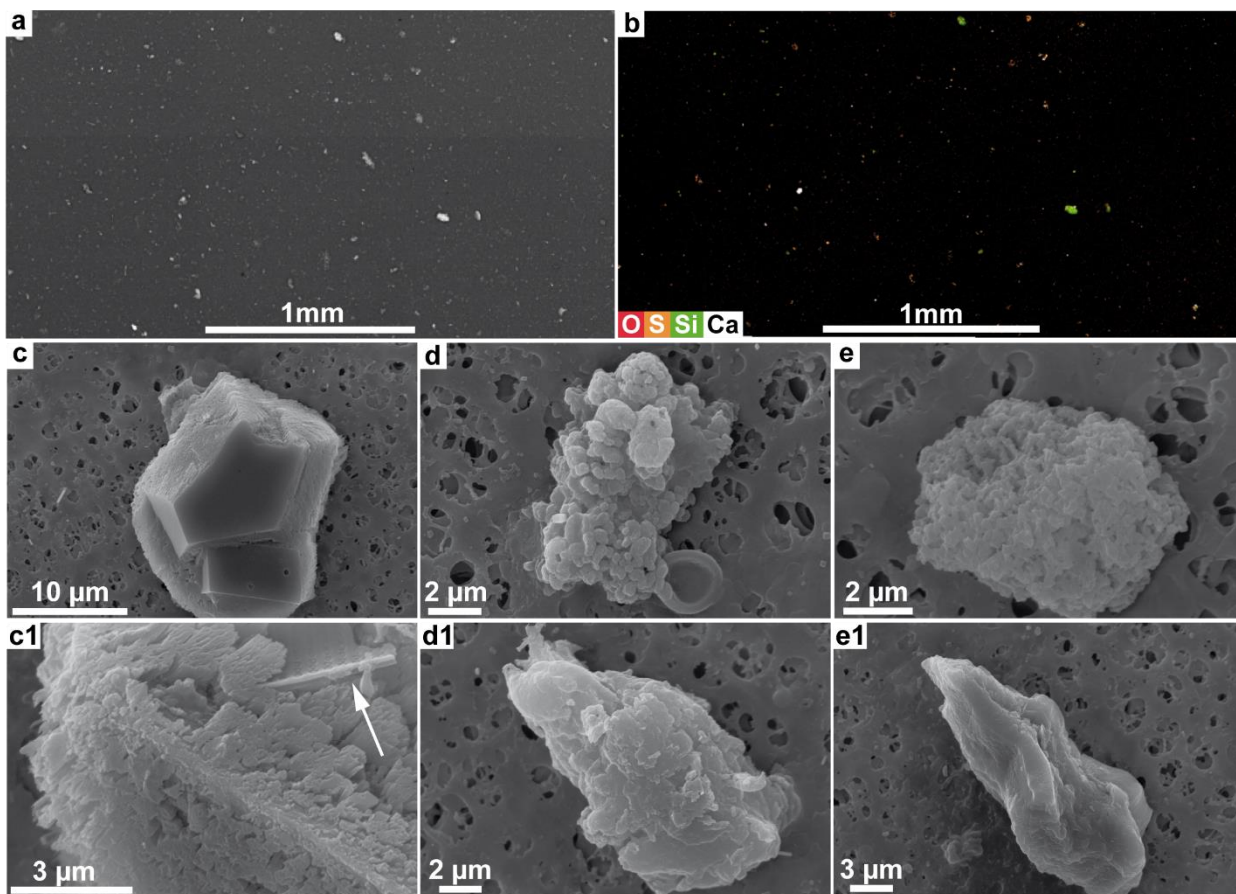


Fig. 7: SEM-analysis of turbid layer filtrate. Various particles on the filter are visible on the SEM picture of a larger area of the filter (a). An element map for O, S, Si and Ca of the same filter area of (a) is shown in (b) revealing that most particles consist of elemental sulfur, Ca-rich particles and silica particles. EDX analysis of individual particles on the filter verified the particles as Ca-carbonates (c, c1), elemental sulfur (d, d1) and silicate phases (e, e1). The white arrow in c1 points to a fragmentary silica shell.

4. Discussion

4.1 Limnological and hydrological conditions in El Zapote Cenote

The water temperature profile (Fig. 3a) offers valuable clues on the hydrological conditions in the El Zapote cenote. Mixing of the water in the narrow cenote shaft from 0–30 m water depth is indicated by constant temperatures and oxygenation, whereas linearly increasing temperatures in the wide dome-shaped cenote from 30–55 m water depth and linearly decreasing dissolved oxygen concentrations indicate conductive heat transport and oxygen diffusion, respectively (Fig. 3a). This suggests that the water body from 0–30 m water depth is mixing-dominated and diffusion-dominated from 30–>50 m water depth. This interpretation is also supported by constant EC values in the cenote shaft and constantly increasing EC values from 30 m water depth down to the top of the halocline at 36.8 m water depth (Fig. 3a). Another indication for stagnant

conditions of the water body is the shape of the halocline itself. Compared to other cenotes of the Yucatan Peninsula being deep enough to reach the halocline, El Zapote cenote particularly differs in the extent of the halocline, the transition zone from fresh to salt water. At El Zapote cenote, the halocline is about 10 m thick (Fig. 3a), as compared to a transition zone thickness of 1–5 m of other cenotes of Quintana Roo -(Kovacs et al., 2017b); Stoessell et al., 1993).

5 The constant decrease of DIC, sulfide and orthophosphate below about 40 m water depth indicates a sink of these chemical species into depths greater than the cenote (>54 m water depth). This sink may result from advection of flowing water masses in conduits or zones of intensified hydraulic conductivity in a deeper cave system at around 60 m below the present sea level. Such deep cave systems could have developed during glacial sea level low stands (e.g. Smart et al., 2006).

In general, the water body of El Zapote cenote is stagnant from 30 m water depth down to the bottom of the cave where mass transfer is predominantly due to chemical diffusion. This is essential for the understanding of hydrogeochemistry and the ongoing biogeochemical processes in the El Zapote cenote.

4.2 Hydrogeochemical processes in El Zapote cenote

4.2.1 Sedimentary biogeochemical processes

15 The anaerobic conditions and high concentrations of metabolites such as S(-II) and CH₄ can be attributed to anaerobic heterotrophic organic matter (OM) decay in the debris mound sediments. Both the debris mound and the cenote floor are covered with a relatively thick layer (~1 m) of OM, mostly leaves and other plant remains, according to the descriptions of the divers. As a consequence of stagnancy in the meromictic water body and oxygen deficiency on the cave bottom, this OM is respired by heterotrophic microorganisms in the sediment via anaerobic fermentative and respiratory pathways.

Anaerobic OM degradation by fermentation and sulfate reducing bacteria produce hydrogen and hydrogen sulfide (S(-II)), CO₂ (DIC) and acidity, thus lowering the pH. Elevated concentrations of DIC and S(-II) are found in the halocline (Fig. 3c), and low $\delta^{13}\text{C-HCO}_3^-$ indicate a microbial origin of the hydrogen carbonate (e.g. Mook, 2000) (Fig. 4). Additionally, pH values are more acidic in the halocline (Fig. 3a and b) and sulfate reduction is further supported by decreasing SO₄²⁻/Cl⁻ ratios in the halocline of up to 32% compared to seawater ratio of 5.2 (Fig. 3c) (Stoessell et al., 1993).

Methane-producing archaea (methanogens) metabolize degraded OM releasing CH₄ and DIC. Although this pathway is less energy efficient than sulfate reduction, and methanogenesis is not expected in the presence of sulfate, methanogens may dominate in deeper parts of the sediments where sulfate is already consumed (e.g. (Whiticar, 1999) Diffusion of CH₄ from the sediment into the water column leads to CH₄ concentrations of up to 25 $\mu\text{mol l}^{-1}$ identified in the halocline of El Zapote. Ammonium is likely released from organic matter degradation in the organic rich sediment and is also released to the water column at the halocline.

30 Other common anaerobic heterotrophic metabolic pathways in sediments, such as the reduction of iron, are subordinated processes, most likely due to low concentrations of iron in -limestone- and limited source of siliciclastic materials in this part of the Yucatán Peninsula YP. The elevated but still exceedingly low amounts of dissolved iron in the halocline as compared

to the fresh water body (Fig. S3) are rather not indicative for the absence of these processes, as iron solubility is limited by the affinity to form iron sulfides in the presence of high amounts of S^{-II}.

4.2.2 Water column biogeochemical processes

The redoxcline from 35 to 36.8 m water depth coincides with a peak in turbidity which is detectable both visually (Fig. 2b) and geochemically (Figs. 3a and b). Dissolved oxygen (DO) concentrations drop to undetectable levels at the top of the redoxcline, indicating that anaerobic biogeochemical processes prevail within the redoxcline (Fig. 8).

In our previous study we tentatively attributed these conditions to a full heterotrophic redox zonation due to organic matter decomposition (Stinnesbeck et al., 2017b). Fine organic matter accumulates along the density contrast at the top of the halocline and heterotrophic microbial communities thrive from the aerobic and anaerobic decomposition of this organic matter. This is also indicated in the results of this study by minor nitrification from ~34–35 m water depth (Fig. 3b), non-linearly decreasing dissolved oxygen contents from ~34–35 m and by slightly more acidic pH values above and in the uppermost part of the turbid layer.

Nevertheless, the more detailed data presented in this study now underline the importance of planktonic chemolithoautotrophic processes in the pelagic redoxcline which are driven by the upward diffusion of reduced sulfur, carbon and nitrogen species released from the anaerobic degradation of organic material at the cenote floor. Pelagic redoxclines develop in density stratified marine (e.g. Berg et al., 2015) as well as lake environments (e.g. Noguerola et al., 2015). In redoxclines below the photic zone the microbial community is dominated by chemolithoautotrophs, with a considerable amount of chemoautotrophic production and dark carbon fixation (e.g. Grote et al., 2008; Jørgensen et al., 1991; Jost et al., 2010; Noguerola et al., 2015). The development of pelagic redoxclines was also reported for deep density stratified cenotes of the Yucatán Peninsula (e.g. Socki et al., 2002; Stoessell et al., 1993; Torres-Talamente et al., 2011).

In our previous study members *β-proteobacteria Hydrogeophilaceae* and the *ε-proteobacteria* genus *Sulfurovum* were reported as dominant within the aqueous microbial community. Most members of these bacterial groups are chemolithotrophic or mixotrophic using reduced sulfur compounds or hydrogen as electron donors and oxygen or nitrogen-compounds as electron acceptors (Stinnesbeck et al., 2017b).

The white cloudy turbid layer could be the result of a dense accumulation of these microorganisms e.g. sulfur-oxidizing bacteria, analogous to that reported for Bundera sinkhole in Australia (Seymour et al., 2007). Elemental sulfur particles or polysulfides were detected on the turbid layer filtrate and indicate sulfur oxidation in the turbid layer or redoxcline (Sect. 3.4 and Fig. 7); these particles are formed as intermediates in the microbial oxidation of sulfide (Findlay, 2016).

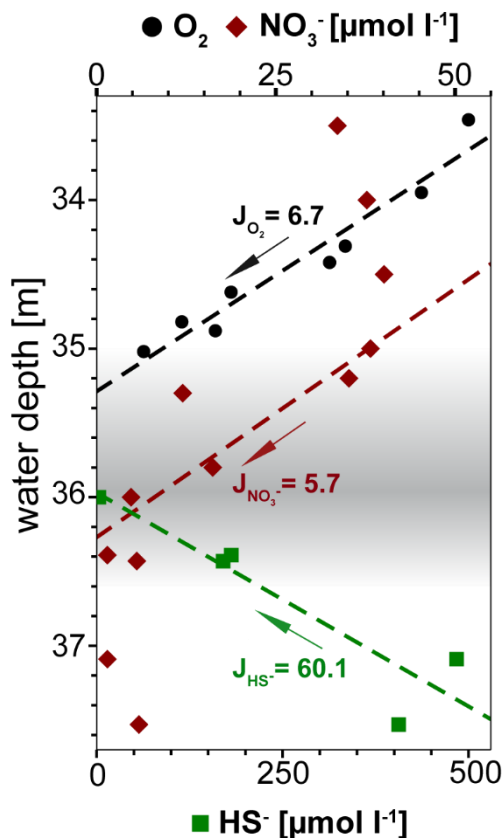
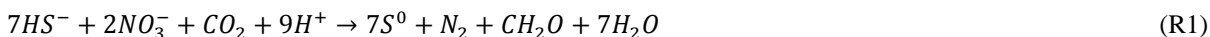


Fig. 8: Concentration profiles of dissolved O_2 , NO_3^- and HS^- (calculated with PhreeqC) in water depths around the redoxcline. The fluxes J are given in $10^{-5} \mu\text{mol m}^{-2} \text{s}^{-1}$. The linear fit of O_2 and HS^- is calculated for the range of plotted values, while for NO_3^- it is calculated only for the range from 34.4–36.6 m water depth. Only O_2 values above detection limit ($6.3 \mu\text{mol l}^{-1}$) from winch profile 2 were considered for the calculation (Table S1).

The oxidation of sulfide in the redoxcline is likely anaerobic, as sulfide vanishes at around 36 m while dissolved oxygen is already at undetectable levels at 35 m water depth and both concentration profiles are not overlapping (Fig. 8). Furthermore, the oxygen flux towards the redoxcline is around one magnitude lower than the flux of the reduced sulfur species HS^- indicating that sulfide oxidation via aerobic pathways is minor (Fig. 8). Thus, sulfide oxidation within the redoxcline must be predominantly via anaerobic pathways. As the downward flux of NO_3^- towards the redoxcline intersects with the upward flux of HS^- (Fig. 8), assimilatory anaerobic sulfide oxidation could be obtained with nitrate as terminal electron acceptor producing elemental sulfur and nitrogen under the consumption of protons (e.g. Bailey et al., 2009). The overall mass-balanced energy generating reaction for chemoautotrophic nitrate-driven anaerobic sulfide oxidation (ND-SO) is given in reaction (R1):



According to reaction (R1) ND-SO could account to one third of the HS^- oxidation, despite the flux of NO_3^- towards the redoxcline is around on order of magnitude lower than the HS^- flux (Fig. 8). Furthermore ND-SO is acid consuming and

sulfide oxidation to elemental sulfur is more acid consuming than the full sulfide oxidation to sulfate (see also Visscher and Stolz, 2005). The abundance of elemental sulfur particles found in the turbid layer filtrate (Fig. 7) indicates that sulfide oxidation to elemental sulfur is predominant. Full oxidation of sulfide to sulfate is less likely as no increase of sulfate is observed in the redoxcline (Fig. 3c). Maxima in pH are known to occur when sulfide is oxidized to elemental sulfur with nitrate as electron acceptor (Kamp et al. 2006). In consequence, the minimum of nitrate in the redoxcline and slight alkaline pH shift, indicate that ND-SO is a relevant process in the redoxcline (Fig. 3b and c, Fig. 8). Therefore, the proton consuming ND-SO could be the biogeochemical process in the redoxcline creating a disequilibrium in the carbonate dissolution-precipitation reaction, favoring calcite precipitation. This mechanism was recently reported for the formation of stromatolites below the photic zone of the Arabian Sea. There, a collective effect of proton-consuming ND-SO and alkalinity-producing sulfate driven-oxidation of CH₄ (SD-OM) leads to authigenic carbonate precipitation in microbial mats in the vicinity of CH₄-seeps (Himmeler et al., 2018).

Anaerobic SD-OM (e.g. Bailey et al. 2009) is likely to occur at the redoxcline, as dissolved CH₄ concentrations vanish at around the same depth of sulfide (~36.5 m) and $\delta^{13}\text{C-CH}_4$ values show a strong peak towards higher values at the same water depth (Fig. 4). Autotrophy also supports calcite precipitation by taking up CO₂ for the synthesis of biomass (Castanier et al., 1999; Kosamu and Obst, 2009). Although a decrease of DIC is not observed at the redoxcline, chemolithoautotrophy is indicated by the $\delta^{13}\text{C-HCO}_3^-$ in the water body (Fig. 4). The peak of higher values in the redoxcline indicates inorganic carbon assimilation by microorganisms (dark-CO₂ uptake). Organisms usually prefer to metabolize ¹²C (it takes less energy to break the ¹²C bond instead of ¹³C), which results in higher $\delta^{13}\text{C-HCO}_3^-$ values in the remaining dissolved inorganic carbon. Hence, the peak towards more positive $\delta^{13}\text{C-HCO}_3^-$ values identified in the redoxcline of El Zapote at ~36 m water depth, may be attributed to microbial CO₂ assimilation or dark CO₂ fixation.

4.3 Hypothesis on Hells Bells formation

It was suspected before that Hells Bells form within the freshwater body of El Zapote cenote (Stinnesbeck et al., 2017b). The depth zone of Hells Bells formation within the fresh water layer can now be narrowed down by the application of [Eq. \(2\)](#) with a given distribution coefficients D(Mg) of the temperature-dependent partitioning of Mg into calcite. ~~in equation (Eq. 2):~~

$$(Mg/Ca)_{\text{solution}} = \frac{(Mg/Ca)_{\text{solid}}}{D(Mg)} \quad (\text{Eq. 2})$$

Applying the mean value of Mg/Ca_{solid} determined for Hells Bells calcites (Table 1) and D(Mg) at 25°C given by Huang and Fairchild (2001) and Rimstidt et al. (1998), the calculation of Mg/Ca of the solution from which the Hells Bells calcite precipitated yields a Mg/Ca_{solution} of 0.73 and 1.06, respectively. Mg/Ca_{solution} ratios in this range are found in the water of the redoxcline and the uppermost top of the halocline in 36–37 m water depth, thus supporting the interpretation that Hells Bells formation takes place in the redoxcline (Fig. 3c and Table S4).

The calcite crystals found in the turbid layer filtrate give further hints on calcite precipitation in the redoxcline (Fig. 7c). It is not yet known whether these particles represent autochthonous matter of the turbid layer. Nevertheless, formation of calcite crystals at the density boundary is likely, as fine particulate matter is accumulated there and may act as crystallization seeds. This process is indicated by calcite crystal formation around silica shells (Fig. 7c1). The high sulfur contents found in Hells Bells calcite also supports this assumption as small sulfur particles are abundant in this water layer and are easily enclosed in calcite crystals growing there (Table 1).

Based on the indications of Hells Bells formation in the redoxcline and taking the biogeochemical processes discussed before into account we propose the following scenario illustrated in Figure 9. It summarizes the biogeochemical processes inducing calcite oversaturation and calcite precipitation in the turbid layer and the redoxcline of El Zapote cenote.

Heterotrophic bacterial decomposition of organic matter in sediment of the debris mound releases CO_2 (HCO_3^-), nutrients (P_{ortho}) and reduced species of sulfur ($\text{S}(-\text{II})$) and nitrogen (NH_4^+). Due to the stagnant conditions in the cenote, these species are transported via diffusion, thereby allowing for the formation of a defined and stable redoxcline. Here, anaerobic chemolithoautotrophy, and especially proton consuming nitrate-driven sulfide oxidation (ND-SO), increase alkalinity, thus favoring calcite precipitation (Fig. 9). The required Ca^{2+} ions for calcite precipitation are constantly supplied to the redoxcline by upward diffusion from the Ca-enriched saline water body (Stinnesbeck et al., 2017b).

Biogeochemical processes involved in Hells Bells formation

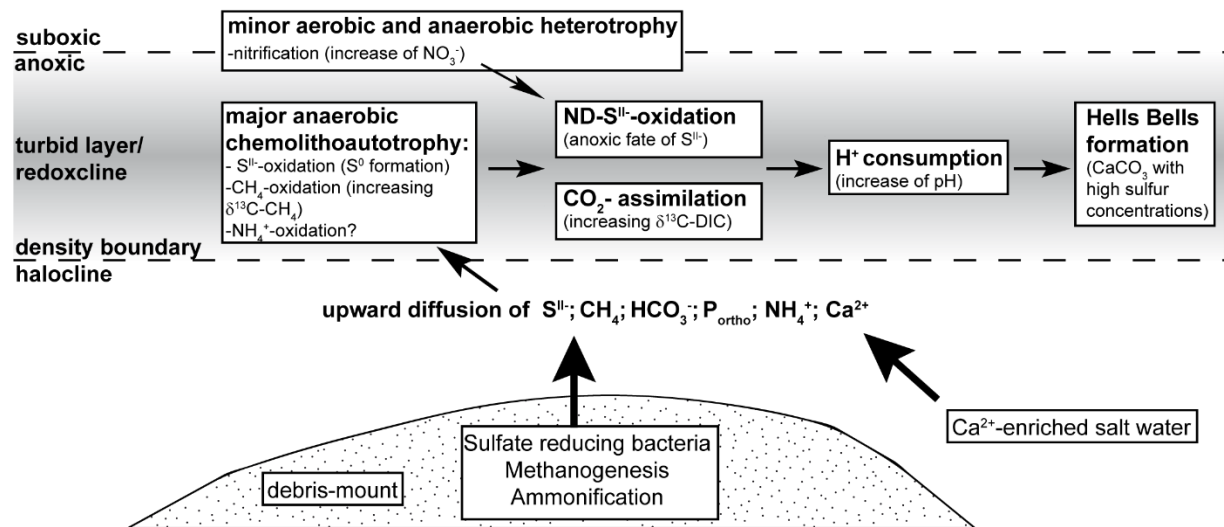


Fig. 9: Scheme of the biogeochemical processes involved in the sediment and redoxcline of the water column of the El Zapote cenote that lead to Hells Bells formation.

4.3.1 Calcite precipitation rates

In order to test the plausibility of the hypothesis on Hells Bells formation within the redoxcline, calcite precipitation rates of both biogeochemical processes, ND-SO and CO₂-assimilation are assessed. All used parameters and results are given in Table S5.

- 5 The overall chemical reaction of the carbonate balance is given in R2 and the partial reactions of R2 are given in R2a–R2d. Under equilibrium conditions, carbonate precipitation after R2 is acid producing because for each mole of precipitated calcite in reaction (R2) one proton is released to compensate for the abstracted carbonate ion (R2b) due to proton shift in the partial reactions (R2b–R2d).



- 15 Assuming that up to one third of the hydrogen sulfide flux towards the redoxcline (Fig. 8) is oxidized by the proton consuming ND-SO after reaction (R1) and that the proton consumption of this process is buffered by both, the observed increase in pH values of 0.04 within the redoxcline (Fig. 3b) and calcite precipitation after reaction (R2), allows for an estimation of the calcite precipitation rate R_{ND-SO} derived from ND-SO within the redoxcline after ~~equation~~ (Eq. (3)).

$$R_{ND-SO} = \frac{1}{3} J_{HS^-} * \left(\frac{\alpha}{m_{HS^-}} - \frac{\beta}{m_{HS^-}} \right) \left[\frac{mol}{m^2 s} \right] \quad (Eq. 3)$$

With $m_{HS^-} = 7$ (moles of HS⁻ in R1); $\alpha = 9$ (moles of H⁺ in R1); $\beta = 0.91$ (moles of H⁺ consumed in pH increase of 0.04).

- 20 Equation 3 yields that 7.3 mmol calcite m⁻² a⁻¹ or 0.73 g calcite m⁻² a⁻¹ could be precipitated within the redoxcline due to ND-SO. This adds up to a total calcite precipitation rate R_{ND-SO} of 2.2–6.2 kg calcite a⁻¹ in the whole redoxcline of the circular 60–100 m wide El Zapote cenote.

- 25 Additionally, the second biogeochemical process in the redoxcline that might lead to calcite precipitation, the CO₂-assimilation or dark CO₂-fixation (Fig. 9), has to be taken into account. Under equilibrium conditions, for each mole of abstracted or assimilated CO₂ one proton is consumed due to the carbonate balance (R2a–R2d) and this proton consumption is compensated by calcite precipitation (R2). The CO₂-assimilation rate in the redoxcline can be approached by the upward flux of DIC ($J_{DIC(HCO_3^-)}$) towards the redoxcline (calculated as HCO₃⁻ and given in Fig. S4). The calcite precipitation rate $R_{CO_2-assim.}$ derived from the CO₂-assimilation is then limited to the half of the DIC flux towards the redoxcline as given in ~~equation~~ (Eq. (4)).

30 $R_{CO_2-assim.} = \frac{1}{2} J_{DIC(HCO_3^-)} \quad (Eq. 4)$

Equation 4 yields a calcite precipitation rate $R_{\text{CO}_2\text{-assim.}}$ of 40 mmol calcite $\text{m}^{-2} \text{a}^{-1}$, which is 12–34 kg a^{-1} in the whole redoxcline. Such rates seem reasonable when compared to reported dark CO_2 -fixation rates from 0.2–2.7 $\mu\text{mol CO}_2 \text{l}^{-1} \text{d}^{-1}$ of deep marine pelagic redoxclines in the Black Sea (Jørgensen et al., 1991), the Baltic sea (Glaubitx et al., 2009; Jost et al., 2008) as well as for a deep karstic lake pelagic redoxcline (Noguerola et al., 2015) (Table S5). Applied to the redoxcline of El Zapote cenote, these rates would yield calcite precipitation rates from 11–420 kg calcite a^{-1} in the 0.5 m thick and 60–100 m wide redoxcline of the circular El Zapote cenote.

The summed up calcite precipitation rates of $R_{\text{ND-SO}}$ and $R_{\text{CO}_2\text{-assim.}}$ in the redoxcline can be converted to calcite growth rates from 0.27–1.46 $\mu\text{m a}^{-1} \text{m}^{-2}$, taking a calcite density of 2.7 g cm^{-3} into account. Calcite growth, however, is likely to be concentrated on a much smaller area, i.e. the crystal or substrate surfaces of Hells Bells hanging in the redoxcline, the cenote walls and the tree stem within the redoxcline (Figs. 2 and 3). This would result in a higher actual calcite growth rate, for example if the calcite growth is concentrated to 1 % of the area, this would result in actual growth rates of 27–146 $\mu\text{m a}^{-1}$.

These growth rates are close to the reported net growth rates of 12–90 $\mu\text{m a}^{-1}$ for a U-series dated Hells Bells specimen from the tree (Stinnesbeck et al., 2017b) demonstrating the plausibility of Hells Bells formation by the biogeochemical mechanisms proposed in this study. However, the actual calcite growth rates must be significantly higher than the reported net growth rates as these rates also comprise repeated phases of calcite dissolution (see Sects. 3.1.1 and 3.2; Stinnesbeck et al., 2017b).

Eventually, the comparison of the estimated actual calcite precipitation rates in the redoxcline and the reported net growth rates is hindered by a lack of data of both, the actual area of calcite precipitation and the time and intensity of calcite dissolution.

4.3.2 The role of halocline elevation in Hells Bells formation

Hells Bells formed in modern to at least historic times and occur in a relatively ~~wide-thick~~ vertical zone of about 10 m from 28–38 m water depth (Stinnesbeck et al., 2017b). This indicates that their underwater growth occurred under environmental conditions similar to the ones detected by us, as modern sea-~~water~~-levels were already reached at about 4.5 ka BP (Hengstum et al., 2010) and thus significantly earlier. ~~Modern~~ Hells Bells therefore precipitate either permanently in the entire depth zone reaching from 28–38 m, or in the narrow 1–2 m ~~wide-thick~~ redoxcline or turbid layer above the halocline (Fig. 9). According to the data presented here the latter hypothesis appears much more likely to us. Therefore, we propose that growth of Hells Bells is a non-permanent episodic process which majorly depends on the halocline elevation in the cenote (Fig. 10). The depth of the halocline generally increases with increasing distance to the coast (Fig. 10a) (e.g. Bauer-Gottwein et al., 2011). The halocline depth position is a function of the hydrostatic pressure of the overlying fresh water layer, i.e. its thickness and the sea level, ~~with a general trend of increasing depth with increasing distance to the coast (Fig. 10a) (e.g. Bauer-Gottwein et al., 2011).~~ Therefore, the halocline elevation can vary on multiple timescales in response to droughts, recharge events and annual tidal fluctuations that are superimposed upon on a longer term sea level change. Extensive droughts occurred repeatedly in the Holocene (Hodell et al., 2001) and could have led to a prolonged elevation of

the halocline as the fresh water layer was thinner in consequence of decreased precipitation (Evans et al., 2018). Although on much shorter timescales, extraordinary recharge events (e.g. hurricanes) must have an effect on the depth position of the halocline (Fig. 10b). Generally, during these events of enormous precipitation, the halocline is temporarily pushed downwards by the amount of the fresh water infiltrating into the Yucatán karst aquifer. However, it has been reported that

5 the halocline can also be elevated as a response to precipitation events (Escolero et al., 2007). ~~Escolero et al. (2007) observed a piston-like effect regarding the halocline elevation in response to recharge in the Northern Yucatán aquifer. During recharge events, the water flow is predominantly vertical and salt water is temporarily pushed down. In the period after these events, the halocline bounces back and oscillates until a static equilibrium is again reached. Additionally, Escolero et al. (2007) detected a halocline elevation of 17.5 m in between measurements in the years 2000 and 2003 in an observation well indicating a highly dynamic halocline elevation within the Yucatán karst aquifer, that might be recharge-driven.~~

10 Regional and local vertical and lateral hydraulic transmissivities of both the epi- and the phreatic karst can result in spatially variable hydraulic pressure of the fresh water lens (Williams, 1983), thus leading to a lowered halocline beneath areas of higher, and an elevated halocline beneath areas of less vertical hydraulic transmissivity. El Zapote cenote is located in the

15 Holbox fracture zone that in the area is characterized by N-S trending lineaments of increased permeability (Bauer-Gottwein et al., 2011, and references therein). Hurricanes that pass the area frequently (Farfán et al., 2014) could therefore lead to episodic elevation of the halocline (Fig. 10).

Halocline dynamic at El Zapote Cenote

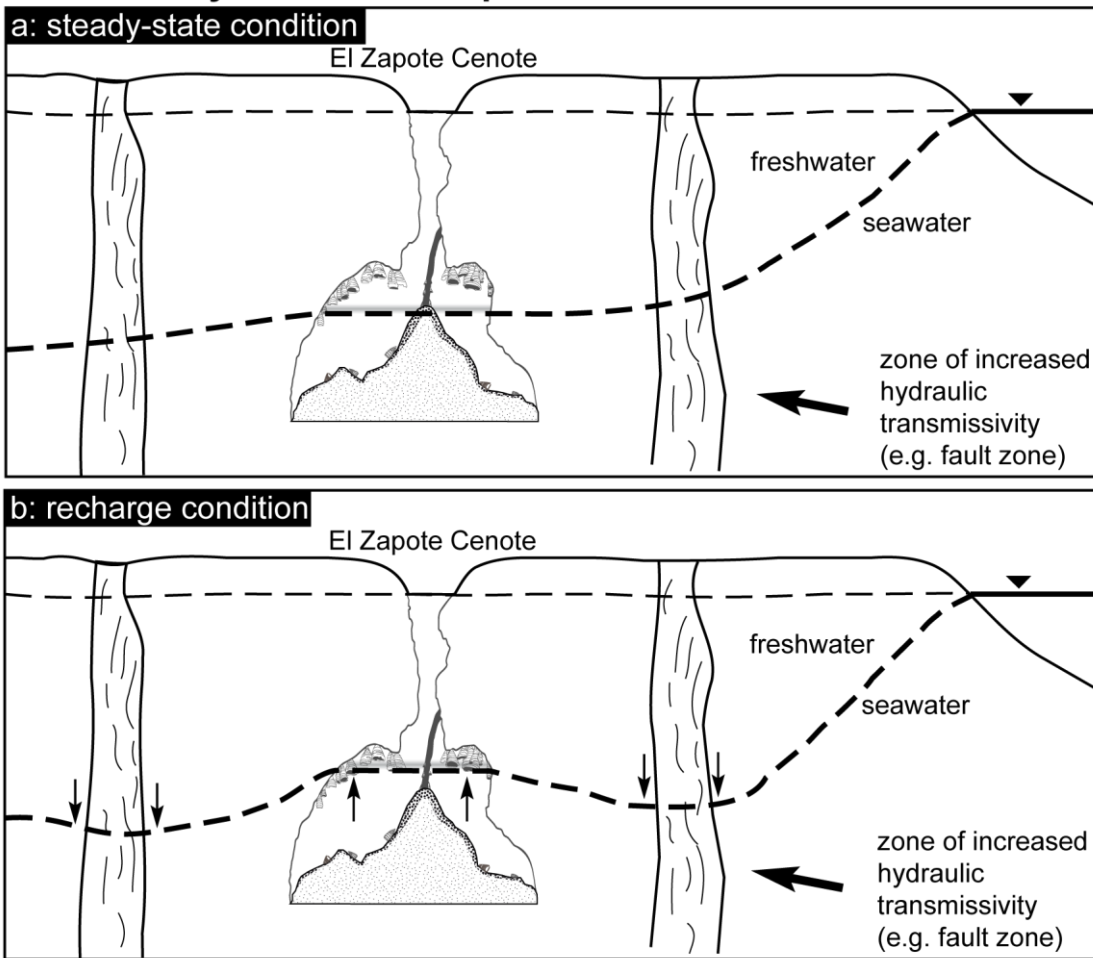


Fig. 10: Sketch of dynamic halocline elevation within the Yucatán Karst aquifer. Halocline depth increases with increasing distance from the coast in a steady-state condition (a). Recharge events (e.g. hurricanes) result in a lower halocline beneath areas of high vertical transmissivities and an elevated halocline in areas of low hydraulic transmissivity, e.g. El Zapote cenote (b).

- 5 Both, a prolonged halocline elevation during droughts and short-term but frequent recharge-driven halocline elevations could result in the presence of Hells Bells in a zone of 28–38 m water depth. Furthermore, repeated phases of precipitation and dissolution indicated by the alternating layers of dog-tooth calcite and microcrystalline calcite of Hells Bells thin sections (Stinnesbeck et al., 2017b) suggest an episodic halocline elevation. During episodes of an elevated halocline, precipitation or dissolution of Hells Bells may occur in lower depths due to the concurrent elevation of the calcite precipitating redoxcline
- 10 and the underlying sulfide-rich and carbonate undersaturated water.

~~The time scale of the frequent recharge-driven episodic halocline elevations is unclear to date, but could reach from a few days up to several weeks. Nevertheless, the estimations of calcite precipitation rates in the redoxcline suggest that Hells Bells growth on such short time scales is possible (Sec. 4.3.1).~~

A variable halocline depth position at El Zapote cenote is also supported by the positive correlation with water depth of Sr/Ca and a negative correlation of $\delta^{13}\text{C}$ of the Hells Bells calcite (Fig. 6). Hells Bells formed in lower water depths show slightly higher contents of the trace elements Sr, Mg and Ba and slightly lower $\delta^{13}\text{C}$ values than Hells Bells formed in greater water depths (Table 1 and Fig. 6). The higher incorporation of the trace elements Sr, Mg and Ba is either obtained by faster growth rates (Tesoriero and Pankow, 1996), or by elevated concentrations in the solution from which the calcite precipitated. The latter process is more likely, as the amount of ~~seawater-saltwater~~ increases in the turbid layer when the halocline is located at lower water depths. Lower $\delta^{13}\text{C}$ values support this assumption, as lowest $\delta^{13}\text{C-HCO}_3^-$ values are detected in the halocline of the modern El Zapote cenote (Fig. 4).

The increase in ~~seawater-saltwater~~ in the lowermost fresh water and the turbid layer could result from increased salinity of the freshwater body during droughts. Furthermore, turbulences induced by halocline elevation as a reaction to recharge events (Kovacs et al., 2017b). Minor mixing of the water bodies would be sufficient to increase the concentrations of Sr, Mg, Ba and decrease $\delta^{13}\text{C-HCO}_3^-$ in the turbid layer, as seen at El Zapote.

Following ~~this~~ model of ~~an episodic -hydraulically driven-~~ halocline elevation, the question remains why Hells Bells are restricted to a zone of 28–38 m water depth. This range could solely depend on the hydraulic conditions, e.g. Hells Bells formation reflecting maximum and minimum elevations of the halocline ~~during-as a result of droughts,~~ recharge events ~~and long-term sea level changes~~. The lower (38 m) level of Hells Bells formation may represent the stable environmental conditions in the modern El Zapote cave, influenced only by the thickness of the freshwater body and the mean sea level. The upper range boundary, on the other hand, could well be given by the shape of the sinkhole and limnological conditions in the narrow cenote shaft reaching from 0–30 m water depth. In this latter unit the water body is mixing- rather than diffusion-dominated (Fig. 3. Sect. ~~ion~~ 4.1). A rise of the halocline to about 28 m water depth would therefore lead to an exposure to fast and convective oxygen supply from the mixed-in fresh water body above, and consequently to aerobic microbial sulfide oxidation, which is an acid-producing reaction (e.g. Jones et al., 2015). Hells Bells formation would then stop as it is tied to ~~the~~-anaerobic ND-SO. The occurrence of a zone of brown-colored manganese oxide coatings on Hells Bells and ~~the~~ cave wall at and above 30 m water depth, indicates that the redoxcline must temporarily have reached up to this level (Fig. 2c, d and Fig. S6). Manganese dissolved in the halocline and turbid layer was then oxidized to manganese oxide precipitates (Fig. S6).

The dynamic history of halocline elevation at El Zapote cenote cannot be resolved to date but raises an interesting issue of further research on the dynamic hydraulic response of the Yucatán aquifer to extraordinary recharge events, especially as this process could be a key factor for the formation of Hells Bells. We are currently addressing this issue by logging the hydraulic head of the fresh water body and the electrical conductivity at a fixed position in the halocline of El Zapote cenote.

4.3.3 Shape of Hells Bells

Previously, we attributed the growth of Hells Bells to microbial mediation (Stinnesbeck et al., 2017b). We hypothesized that autotrophy (ammonia oxidation) and denitrification are the main factors that trigger calcite precipitation at the surface of the Hells Bells and that calcite precipitation could further be supported by the presence of negatively charged extracellular polymeric substances (EPS), leading to the accumulation of Ca^{2+} ions and to supersaturation of calcite within biofilms (e.g. Dupraz et al., 2009). However, the large size and form of the dog-tooth calcite crystals of Hells Bells rather resemble slow growing inorganic calcite crystals rather than biologically-mediated precipitates (Fig. 2e and Fig. 5). This hypothesis is supported by Bosak & Newman (2005) who investigated microbial kinetic controls on calcite morphology and found that microbially mediated calcite precipitated at low calcite supersaturation shows more anhedral crystal morphologies, compared to the more euhedral abiotic ones. Although the microbial activity in the redoxcline induces calcite oversaturation, the hypothesis presented in this study is compatible with an inorganic calcite precipitation of the Hells Bells from a biologically mediated water layer.

Ultimately, the hypotheses of dynamic halocline elevation and biogeochemically induced calcite precipitation in the redoxcline can be integrated to explain some morphological features of the Hells Bells. They grow downward and are conically divergent, with a strict horizontally lower margin and a hollow interior. Specimens also tend to be oriented towards the cenote center (Stinnesbeck et al., 2017b). The horizontal downward growth is indicative for a precipitation from a defined layer within the water column (i.e. the redoxcline). Also, an abrupt elevation of the redoxcline as a response to recharge events and a subsequent decelerated drop towards its original position serves to explain the downward growth of Hells Bells. This is indicated by the tendency towards downward orientation of the calcite crystal growth axis (Fig. 2e). The fact that Hells Bells specimens growing on the inclined cave wall are always oriented towards the cenote center could result from a lateral gradient in the chemolithoautotrophic intensity. The “energy sources” (~~i.e. sulfide~~) used by the chemolithotrophic microbial community in the redoxcline ~~are~~ is the released ~~d~~-of reduced carbon, sulfur and nitrogen species from anaerobic organic matter decay in the organic-rich sediments on the debris mound. Both the morphology of the cenote and the diffusive mass transport likely result in radial concentration gradients of upwards diffusing reduced species from sediment of the debris mound. These conditions limit the availability of reduced species in locations of the redoxcline distal to the debris mound, and vice versa. Consequently, the intensity of chemolithoautotrophy and hence calcite oversaturation is preferentially higher in the center proximal to the debris mound and decreases towards the cenote walls. This accounts for both the inclined bells as well as for horse-shoe like horizontal openings of Hells Bells which always face towards the wall. Furthermore, the often observed hollow interior of Hells Bells could be due to preferential growth of the outer edges of Hells Bells, especially the parts facing towards the cenote center. Once such parts of Hells Bells grew slightly more, hence into deeper water depth, this will result in a higher net growth as these parts are likely to be reached more frequently by the narrow zone of calcite precipitation (redoxcline). The conical shape and downward divergence of Hells Bells could be a

continuation from the microscopic to the macroscopic level as the angles of the botryoidal calcite phases in thin section (Fig. S5) strongly resemble those of the large specimens of Hells Bells (Fig. 2d).

Nevertheless, we can also not exclude the potential influence of microorganisms forming a biofilm community on the surface of Hells Bells. Stinnesbeck et al. (2017b) showed that this community does not resemble the planktonically growing microbial biocenosis but forms a distinct community that seems to thrive catalyzing the reduction and oxidation of different nitrogen species. However, it is not known to date to what percentage the activity of these organisms contributes to the shape of the speleothems.

4.4 Prerequisites for the formation of Hells Bells

Hells Bells have so far been identified in a few cenotes only of a restricted area of the North-Eastern ~~Yucatán Peninsula-YP~~ (Stinnesbeck et al., 2017b), although the peninsula hosts many thousands of sinkholes (Bauer-Gottwein et al., 2011). The question thus arises which factors are needed for the generation of these underwater speleothems. The following apparent prerequisites for Hells Bells formation appear likely to us:

- The cenote or sinkhole must be deep enough to reach the halocline in order to have a density stratified water column (meromixis).
- Sufficient input of organic material to the cenote bottom is required to create anoxia in the halocline with a release of reduced sulfur, carbon and nitrogen species.
- A meromictic stagnant water body indicated by a thick halocline is needed that allows for the formation of a redoxcline in which anaerobic chemolithoautotrophy prevails ~~in a lightless environment~~. This leads to a narrow zone of calcite oversaturation in the water body.
- Special hydraulic conditions are needed which allow the halocline to rise and fall in order to form subaqueous speleothems.

5. Conclusion

The unique underwater speleothems termed Hells Bells recently described from El Zapote west of Puerto Morelos on the northern ~~Yucatán~~ Peninsula, Mexico, are most likely formed in the redoxcline, a narrow layer in the lowermost fresh water body immediately overlying the halocline. We propose a biogeochemical mechanism for the formation of these structures, that induces calcite oversaturation and favoring calcite precipitation within the redoxcline. The upward diffusion of reduced sulfur, carbon and nitrogen stimulates a chemolithoautotrophic microbial community thriving above the halocline at El Zapote cenote. Chemolithoautotrophy and proton-consuming nitrate-driven anaerobic sulfide oxidation lead to calcite precipitation, and hence Hells Bells formation, in a narrow depth zone confined to the redoxcline, or turbid layer. We further postulate a dynamic elevation of the halocline as an episodic hydraulic response to both droughts and recharge events ~~,-e.g.-~~ ~~hurricanes,~~ that may accounts for Hells Bells occurrence over a vertical range of 10 m water depth.

Video supplement

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Author contribution

- WS and AGG initialized the project and funding was acquired by WS, AGG and MIS. Sampling was planned and conducted by SR and CS. Underwater sampling was conducted by JAO, DP, CL and EAN, while underwater videographer TV was documenting part of the sampling producing the footage for the video supplement. Instrumentation and methodology was
- 25 provided by MIS for hydrogeochemical analyses, FK for gas and stable carbon isotope analytics, AH and GA for light microscope-imaging and AV for SEM-imaging and GA and. SR, LK and NS collected the data and CS, MIS, FK and AV validated it. SR, CS, MIS, LK, NS interpreted the results. SR developed the hypotheses equally supported by CS and MIS. SR visualized the data and prepared the original draft, MIS supervised. JG, WS and FK critically reviewed the manuscript.

Competing interests

The authors declare that they have no conflict of interest.

Disclaimer

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Table 1: Geochemistry of samples from the lowermost tips of Hells Bells growing on a subfossil ceiba tree that fell into the El Zapote cenote about 3500 cal yr BP. Individual tree bells were sampled at water depths from 31.3 to 37.3 m. The lack of samples in water depths from 34.3 to 36.8 m water depth is due to poor visibility in the turbid layer above the halocline (compare Fig. 1b). Given uncertainties represent 2σ standard deviations.

Sample No.	water depth [m] ^a	Mg/Ca [$\cdot 10^{-3}$]	Sr/Ca [$\cdot 10^{-5}$]	Ba/Ca [$\cdot 10^{-5}$]	Fe/Ca [$\cdot 10^{-5}$]	Mn/Ca [$\cdot 10^{-6}$]	S/Ca [$\cdot 10^{-3}$]	$\delta^{13}\text{C}$ [‰ _{V-PDB}]
1	31.3							
2	31.3	21.4	41.6	1.27	4.3	21	2.78	-13.47 ± 0.01
3	32.8	25.6	41.0	1.19	5.2	26	2.69	-13.69 ± 0.01
4	32.8							

5	33.3	22.0	42.7	1.30	3.8	24	3.05	-13.82 ± 0.01
6	33.3	22.3	38.5	1.14	3.0	16	2.54	-13.43 ± 0.01
7	33.8							
8	33.8	20.8	37.9	1.12	3.5	24	2.72	-13.52 ± 0.01
9	33.9	23.2	40.0	1.18	3.9	22	2.76	-13.68 ± 0.01
10	33.9							
11	36.8	22.3	34.0	0.88	11.3	39	3.14	-12.87 ± 0.02
12	36.8	21.1	34.1	0.88	10.6	28	3.02	-12.85 ± 0.00
13	37.3	24.0	37.3	0.92	6.2	32	3.17	-12.99 ± 0.01
14	37.3							
	Mean	22.5	38.6	1.10	5.76	25.8	2.87	-13.37
	2σ	2.9	5.9	0.31	5.84	12.8	0.42	0.70
