Response to reviewers for REddyProc BGD paper

immediate

Abstract. Dear editor,

We submit a revised version of our manuscript "Basic and extensible post-processing of eddy covariance flux data with REddyProc".

First, we provide a point-by-point response to the reviews. We repeat each reviewer's comment (RC) before each of our author's comment (AC) replies in blue.

The most important changes are

- restructuring of package dependencies (see RC1-4) and
- shortening of the manuscript by moving parts of the original manuscript to a new online supplement (see RC1-2).

Next, we provide a marked-up version of the manuscript, where deleted parts show red strike-through and added parts show to blue wiggly-underlined. Note that in addition to the marked changes, also figures changed axis labels.

1 Response

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1.1 Response RC1

AC: We thank referee 1 for his constructive comments.

RC1-1: The paper adds to a growing segment of extensive method descriptions for reproducible computational research. While the technical focus is laudable, it is also the reason that the manuscript in its current form somewhat misses the scope of BG ("interactions between the biological, chemical, and physical processes"; https://www.biogeosciences.net/index.html). I imagine this might be part of the reason that the other referees declined. Alternative Copernicus journals like AMT ("techniques of data processing"; https://www.atmospheric-measurement-techniques.net/index.html) or GMD ("statistical models", "technical papers"; https://www.geoscientific-model-development.net/) should provide a much better fit.

AC: While the suggested alternative journals maybe fit the content slightly better, they target an audience primarily interested in data processing or model development. Contrary, we want to target an audience of researchers who *use* EC flux data in their studies. We argue that Biogeosciences is the best open access journal to reach this audience. For instance Biogeosciences Journal recently hosted a special issue on eddy covariance data collected in the Australia and New Zealand, which include also more methodological papers "OzFlux: a network for the study of ecosystem carbon and water dynamics across Australia and New Zealand, 2016". Also a search of the keyword "eddy covariance" in the main text on April 4th, 2018 give back 1620 records. We also looked at the number of views in the open review phase and they were more than 350 (April 4th), supporting the argument that Biogeosciences can be the right audience for the article.

RC1-2: Here a few points for consideration in such re-submission: - I suggest shortening the Discussion paper manuscript. It should be straightforward to consolidate 22 pages of heavy methodological detail by about 1/3, and focus on novel aspects.

AC: We shortened the manuscript mainly by moving parts to an online appendix. We strive to report the details in some way for reproducibility.

Specifically we summarized section 2.2.3 (MDS) in the general gap-filling section and moved the section including former Fig. 3 to the online supplement. Next, we moved details of the day-time flux-partitioning to an online appendix. Further we moved former Appendices A1 and C to the online supplement.

RC1-3: - 1. 5: "standard tools available in open source environment for processing high frequency (10 or 20 Hz) data into half-hourly quality checked fluxes". At this time open-source environments for eddy-covariance data processing that actually facilitate open development are only emergent. REddyProc provides a substantial and much appreciated contribution to this movement. I suggest to either substantiate the claim of an abundance of open-source high-frequency data processing environments through providing examples, or to provide a more differentiated overview.

AC: We agree with the reviewer that open source tools for eddy covariance processing are only now emerging. There are now available some open source tools for processing of high frequency data, for instance the Eddy Pro software, which is open source and widely used. In the revised version we included a footnote with a link to a web-page that lists such packages both in the abstract and the introduction.

RC1-4: - 1. 8: While it is true that R is a cross-platform language, this does not mean that research is reproducible by using an R-package across platforms. Known as "dependency hell", installing e.g. REddyProc on a standard Debian Linux distribution requires the co-installation of several operating-system-side libraries (libudunits2-0, libudunits2-dev, udunits-bin, libnetcdf-dev) and even more R-side dependencies (backports, praise, evaluate, highr, mlegp, logitnorm, ncdf4, RNetCDF, minpack.lm, segmented, rprojroot, testthat, knitr). In some operating systems such as Windows, there is hardly any automation available for resolving operating-system-side dependencies, making R-packages with heavy dependencies inaccessible to less experienced users. Most importantly, dependency resolution itself is not reproducible among operating systems, thus rendering reproducible research impossible. A balanced discussion of how REddyProc can be used for reproducible research alongside examples for dependency resolution would add much substance and usability to the manuscript.

AC: The library dependency issue were caused by requirements of system libraries with the two NetCDF packages. In an updated package version we 1) moved NetCDF related functionality to an own package REddyProcNCDF and 2) migrated function from package logitnorm directly to REddyProc. Hence, in the updated package, besides the suggested packages mainly used in generating documentation, only two dependencies are there, specifically the tidiverse package suite and the mlegp Gaussian process package. The first one is almost as standard as base R, and at least one package dependency will be required if we want of use Gaussian processes.

Furthermore, we provide automatically-built docker images, a kind of lightweight virtual machine, on docker hub, so that with docker installed, REddyProc can be used without installation of R software and other package dependencies with a single command. In the revised version we only slightly extended the section on installation (Appendix A, p21) but rather refer to the REddyProc homepage on github for those technical issues.

RC1-5: - 1 15: It could be pointed out that REddyProc has already been adopted for computational research by the flux community, such as in Metzger et al. (2017). These authors also point to a community solution for "dependency hell", an pre-compile REddyProc alongside its dependencies into compute images that contain a turn-key, reproducible and shareable processing environment.

5 **AC**: Thanks for this suggestion. In the revised version we cited the paper when referring to the provided Docker image (Appendix A, p21L11)

1.2 Response RC2

RC2-1: The paper describes the R package REddyProc containing tools for post-processing of eddy covariance data with a focus on CO_2 . The REddyProc package contains tools for reading halfhourly data from different formats, estimating u* thresholds, gapfilling, flux-partioning, data visualisation, and estimation of uncertainties. The results obtained with the routines of the package are compared with other state-of-the art tools. This resulted in no significant differences in the results from the different tools on a montly or annual scale. The REddyProc package contain a more sophisticated way of the treatment of seasons and the possibility to run and take advantage of multi-year data sets. Discussion paperThe paper provides a good description of the rationale of the package and the diffence in calculation method between this package and earlier available tools. The paper also includes an appendix with example of how to run the routines on a full annual dataset. In general, I find that the paper is well written and documented, and that it can provide a very useful reference for scientist using the package. I have some suggestions for improvements in the specific comment below. Providing this package is very good for the flux reasearch community and by using standardized methods the results from different research groups become much more comparable.

AC: We thank reviewer 2 for his encouraging comments.

1.2.1 Specific comments

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RC2-2: Table 1. I find it a bit lazy to write "uStar" instead of the correct symbol " u_* ".

AC: In the revised version we use now symbol " u_* ", aside from referring to source code, where the multiplication character cannot be used.

RC2-3: Figure 2. Caption: I suppose it should be "night-time NEE". The "season–temperature subset" is only understandable after careful reading of the main text.

AC: In a revised version we modified the figure 2 caption to "Concept of the u_* -filter for a subset of data from DE-Tha. Night-time NEE at low u_* friction velocities, i.e with low turbulence, is biased towards lower NEE values compared to cases with higher u_* , although other environmental conditions are similar. The u_* value below which this effect is considered significant, i.e. u_* threshold (dashed line), is estimated by a moving point method on u_* bins (crosses) across half-hourly records (circles)."

RC2-4: p.6, l.12-13: It is not completely clear from the text which time-frame is used for the flagging. What is the chosen aggregation period for a "robust u* estimate" (l. 10).

AC: u_* value of each half hour is compared to the threshold determined for the corresponding season. In the revised version we slightly extend the paragraph (P6L16)

Both, the aggregation periods for which different thresholds are established, i.e. seasons, and the aggregation across different seasons is specified by the user. The first defaults to three months, starting in months, 3,6,9, and 12, the second defaults to maximum across one year as described at (P11L24ff). The choice of both, however, can be adapted to the knowledge of the site. We added extended discussion at P14L15 and Appendix B7.

RC2-5: p.7, l.8.: Probably this is the first time in the paper the "BGC online tool" is mentioned, and should thus be explained and referenced here.

AC: Thanks for noting. In the revised version, we extended the description of the BGC online tool in the benchmarking section (3), include a reference in this earlier method description section (P8L10), consistently refer to the BGC online tool as BGC16 (also in variable subscripts in figures), and added this acronym to Table 1.

RC2-6: p.9, 1.22: The abbreviation "LRC" should be explained.

AC: Thanks for noting. In a revised version, we will keep the description of the Light response curve (LRC) in the specific day-time partitioning section (2.3.2 page 9), but write in this earlier general section (P8L24) "fits a model to observations of daytime NEE and global radiation"

RC2-7: p.12, l.13: The choice of the software used for comparison (here called "benchmarking") should be explained better. The u* filtering is compared to Papale et al. (2006) and the gap-filling and flux partitioning to the BGC online tool. Do these represent (previous) state-of-the-art or have exactly those routines been used for calculating the FLUXNET data-sets?

AC: The required more detailed description of the role of the benchmark tools was described in the more specific sections, 3.2 and 3.5. In the revised version we keep some details in the specific sections but also extended the general description at beginning of the benchmark section (p10L20).

RC2-8: p.14, l.6: Should read "To explore these consequences ..."

AC: Thanks for noting the typo that is now corrected (P11L16)

RC2-9: Discussion paper Fig.4+5: I find it quite interesting that there seems to be a fairly large scatter in the estimated u* thresholds, but this does not seem to translate into a similar scatter in the annual NEE estimates. It would be valuable to have a further discussion on this. Does it e.g. mean that the value of the u* threshold is not very important? A lower u* threshold means that more data are kept which could potentially lower the uncertainty on the annual estimate.

AC: We also did not expect such low sensitivity of the aggregated flux results to the differences in the u_* threshold estimates, and discussed this at sections 3.2.3: "indicates that the sensitivity of NEE to the u_* threshold estimate in the inferred ranges is low, which also explains the large uncertainty of the u_* threshold estimate".

In the revised version we elaborated a bit more on this point (section 3.2. on page 14)

RC2-10: Fig. 6: The symbols of the quality flags are very difficult to read and distinguish. The legend refers to the sub-figures as "top" and "bottom". Here it should be "left" and "right".

AC: In the final (as opposed to review) two column mode formatting, the labels "top" and "bottom" are the correct ones.

The symbol transparency is a compromise between reading a single symbol and conveying the message that most of the values overplot at the 1:1 line. In a revised version we adjust the figure towards less transparency.

RC2-11: Appendix B: I tried out the REddyProc package on my MacBook, First I tried to update all packages and after a little tweaking I managed to load REddyProc. Following the example went very smooth until running the function sEstUStarThresholdDistribution() where I got the message "Error: could not find function "sEstUStarThresholdDistribution". This is probably a minor trivial issue in my implementation that can be solved.

AC: Thanks for noting. sEstUStarThresholdDistribution is a method of the EddyProc class, not a function, and is called as in example B2. In the revised version we updated the line in B5 to explicitly include the class: EddyProc.C\$sEstUStarThresholdDistribution(...) and explicitly refer to section B2.

Basic and extensible post-processing of eddy covariance flux data with REddyProc.

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Abstract.

With the eddy-covariance eddy covariance (EC) technique, net fluxes of carbon dioxide (CO₂) and other greenhouse trace gases as well as water and energy fluxes can be measured at the ecosystem level. These flux measurements are a main source for understanding biosphere-atmosphere interactions and feedbacks by cross-site analysis, model-data integration, and up-scaling. The raw fluxes measured with the EC technique require an extensive and laborious data processing. While there are standard tools¹ available in open source environment for processing high-frequency (10 or 20 Hz) data into half-hourly quality checked fluxes, there is a need for more usable and extensible tools for the subsequent post-processing steps. We tackled this need by developing the REddyProc package in the cross-platform language R that provides standard CO₂-focused post-processing routines for reading (half-)hourly data from different formats, estimating the uStar u* threshold, gap-filling, flux-partitioning, and visualizing the results. In addition to basic processing, the functions are extensible and allow easier integration in extended analysis than current tools. New features include eross year cross-year processing and a better treatment of uncertainties. A comparison of REddyProc routines with other state-of the art tools resulted in no significant differences in monthly and annual fluxes across sites. Lower uncertainty estimates of both uStar u* and resulting gap-filled fluxes with the presented tool was achieved by an improved treatment of seasons during the bootstrap analysis. Higher estimates of uncertainty in day-time partitioning resulted from a better accounting of for the uncertainty in estimates of temperature sensitivity of respiration. The provided routines can be easily installed, configured, used, and integrated with further analysis in further inquiry. Hence the eddy covariance community will benefit from using the provided REddyProc package, allowing easier integration of standard post-processing with extended analysis.

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¹fluxnet.fluxdata.org/2017/10/10/toolbox-a-rolling-list-of-softwarepackages-for-flux-related-data-processing

1 Introduction

The availability of ecosystem level observations of net ecosystem exchange (NEE) of carbon dioxide (CO₂) and other gases, latent heat (LE) and sensible heat (H) fluxes measured by the eddy covariance (EC) method (Aubinet et al., 2000) allowed a boost in ecosystem understanding at site to global scales (Baldocchi et al., 2017). The EC method provides half-hourly or hourly records of turbulent fluxes between an entire ecosystem and the atmosphere. These data are derived from high frequency measurements (10 or 20 Hz) of wind speed and direction together with measurements of air scalar characteristics such as CO₂ and water vapor concentration, or temperature. Methods to compute fluxes from high frequency measurements, for the quality check and quality assessment (QA/QC), and for the storage corrections have been consolidated in the last decades (Rebmann et al., 2012; Foken et al., 2012; Aubinet et al., 2012). Depending on the site characteristics, the NEE fluxes also need to be storage correctedand are available as open source software². Although measured continuously, the (half)hourly EC data contain gaps due to instrument malfunction or meteorological malfunctioning or records which are not representative of the ecosystem because of micrometeorological conditions under which the assumptions of the EC technique are not met. Such conditions include insufficiently developed turbulence or low signal stability (details in e.g. Foken and Wichura, 1996; Foken et al., 2004; Göckede et al., 2004; Foken et al., 2012). Data obtained under these unfavorable conditions are subject to several constraints, technically correct but not representative of the ecosystem. Hence, (half-)hourly records are flagged for different quality levels marked with different quality flags and need further extensive post-processing as described by Papale et al. (2006).

NEE from periods with low friction velocity ($uStaru_s$) (Aubinet et al., 2012) need to be detected and filtered out to avoid systematic biases in nighttime NEE (Papale et al., 2006). The screened flux time series with gaps need to be filled (Reichstein et al., 2005a) using the available flux data and auxiliary micrometeorological meteorological measurements. An additional information can be obtained from NEE thanks to flux partitioning methods, that provide model estimates of gross primary production (GPP) and ecosystem respiration (R_{eco}) (Reichstein et al., 2005a). These gross fluxes are important to understand land-atmosphere interactions.

All these post-processing steps need to be performed routinely for EC data. Hence, it is desirable to have automated and reproducible post-processing tools available that can be easily used, extended, and integrated into researcher's own workflow. For this purpose we compiled all routines for the important CO_2 -focused post-processing steps in the REddyProc package for the cross-platform, free R language. The REddyProc package loads time series of quality checked and storage corrected fluxes and the basic set of meteorological variables and provides a software environment to perform uStar u_x threshold detection and filtering, gap filling and partitioning. Furthermore, a series of other functionalities like data import routines and data visualization are provided.

The objectives of the paper are to 1) provide a reference that describes the methodology of the processing used in the REddyProc package, and 2) show that the obtained results do not differ systematically from results obtained with standard

² fluxnet.fluxdata.org/2017/10/10/toolbox-a-rolling-list-of-softwarepackages-for-flux-related-data-processing

post-processing implemented in the FLUXNET community (based on Papale et al., 2006; Reichstein et al., 2005a; Lasslop et al., 2010; Pastorello et al., 2017).

Table 1 explains used abbreviations. The first part of the paper (section 2) describes the post-processing methods. The second part (Section 3) presents the benchmarks the REddyProc implementation with standard post-processing tools. It details differences in the implementations and possible consequences in obtained results and aggregated fluxes. An appendix gives an overview of the package with general design, an example of the post-processing, and links to help and resources so that readers can get started resources that get readers started with post-processing their own data.

2 Methods of Post-processing

The post-processing relies on half-hourly or hourly measurements of NEE and ancillary meteorological data of uStar, Rg, Tair u_{*}, global radiation (Rg), air or soil temperature , and VPD(Tair, Tsoil), and vapor pressure deficit (VPD). The fluxes should be quality checked and, if applicable, storage corrected before their use in the package.

The post-processing follows a specific workflow:

- 1. determination and filtering of periods with low turbulent mixing (uStar-filtering y_* -filtering),
- 2. replacing missing data in the half-hourly/hourly records (gap-filling), and
- 3. partitioning NEE into the gross fluxes GPP and R_{eco} (flux-partitioning).

Usage of the REddyProc package follows this data post-processing workflow (Fig.1). The following sections explain the steps in more detail.

2.1 uStar-filtering u_* -filtering

Determining periods with low turbulent mixing is a critical step in the EC data post-processing. Standard steady state and integral turbulence characteristics tests in the initial processing exclude the most problematic records of sensible (H), latent (LE) heat and CO_2 fluxes (Foken and Wichura, 1996). However, it is well known (summarized in Aubinet et al., 2012, chapter 5), that such a quality checking strategy is not sufficient, especially in the case of CO_2 . Stable stratification that is present often during the night-time dampens turbulence and leads to an underestimation of the night-time NEE, i.e. the ecosystem respiration (Gorsel et al., 2007). Massman and Lee (2002) proposed that unfavorable conditions could be detected by inspecting the relationship of night-time NEE versus $\frac{1}{1000}$ within similar time period and similar environmental conditions respiration should not be dependent on the $\frac{1}{1000}$ At low $\frac{1}{1000}$ will will a threshold of $\frac{1}{1000}$ and $\frac{1}{1000}$ be established above that nighttime fluxes are considered valid. Hence, the $\frac{1}{1000}$ threshold is the minimum $\frac{1}{1000}$ above which respiration reaches a plateau (Fig. 2). This threshold

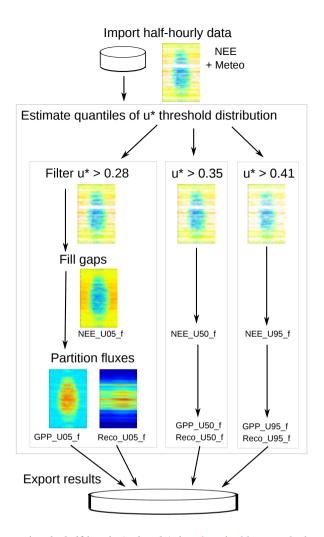


Figure 1. The workflow starts with importing the half-hourly (or hourly) data, here in this example the year 1998 of site DE-Thais plotted as an example. Next, a probability distribution of uStar ux threshold is estimated for each season. Gap-filling and flux-partitioning is performed for several quantiles of this distribution for an estimate of uncertainty. Finally the results are exported.

is specific for each season of a site year. Uncertainties in the $\frac{\text{uStar}}{u_x}$ threshold estimate represent one of the largest uncertainty components in the post-processing of NEE.

There are at least two methods of estimating the uStar u, threshold: the moving point method (Reichstein et al., 2005c; Papale et al., 2006), which is currently more routinely used, and the break-point detection method (Barr et al., 2013).

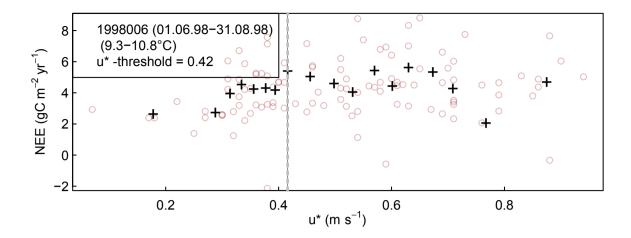


Figure 2. Concept of the uStar-filter: u_* -filter for a subset of data from DE-Tha. Night-time NEE at low uStar- u_* friction velocities below a threshold, i.e with low turbulence, is biased low-towards lower NEE values compared higher uStar-to cases with otherwise similar-higher u_* , although other environmental conditions are similar. The uStar- u_* value below which this effect is considered significant, i.e. u_* threshold (dashed line), is estimated by a moving point method on uStar- u_* bins (crosses) across half-hourly records (circles)here for a season-temperature subset of data from DE-Tha.

2.1.1 Moving point method for uStar u_*

The method of (Papale et al., 2006) Papale et al. (2006) detects a plateau in the relationship of night-time NEE versus uStar u_* among all records within temperature subset by a moving point test of records binned into different uStar u_* bins.

The nighttime data (default: Rg < 10 Wm⁻²) is split into different times of year, here called seasons, to account for differing surface roughness. Then the data of each season is split into default six temperature subsets of equal size (according to quantiles). Within each temperature subset data is split into 20 about equally sized uStar us bins. The default moving point method, called Forward2, determines the threshold based on these uStar us bins. It checks for each bin if the mean NEE is higher than 0.95 times the mean of the following 10 bins. If this holds true also for the next bin, the mean uStar us of the bin is reported as threshold. There are often subsets of data, where no clear threshold can be detected. Hence, there are quality criteria on whether the estimate of a given subset is used in subsequent aggregation. One quality criterion specifies that temperature and uStar us should not be correlated within the temperature subset, another requires a minimum number of valid records within a subset. Next, the uStar us estimates for different temperature classes and periods seasons (details in section 3.2.1) are aggregated to derive a robust uStar us estimate. Within one season, the median is taken across the estimates of different temperature subsets. Within one year, the maximum is taken across the associated seasons.

Records during night-time with uStar-u* smaller than the estimated threshold are flagged as invalid and are replaced in the subsequent gap-filling processing step.

In addition, each half-hour after records with u_* smaller than the threshold is flagged to be invalid.

2.1.2 Breakpoint detection method for uStaru*

Alternatively, a breakpoint-detection can be applied to the unbinned data, that avoids the sensitivity of the Moving point method to the specifics of the binning schemes (Barr et al., 2013). REddyProc provides this method by estimating the breakpoint based on unbinned records within the seasons/temperature subsets using the segmented R-package (Muggeo, 2003, 2008). However, REddyProc differs from (Barr et al., 2013) Barr et al. (2013) by keeping the same aggregating scheme of seasonal/temperature estimates to annual thresholds as with the Moving point method.

10 2.1.3 Bootstrapping uncertainty of the uStar u_* threshold

Estimates of the uStar- u_* threshold are often sensitive to the specifics of the combination of methods and the data, e.g. the binning, minimum number of records within a season or temperature subset, criteria in aggregation, etc. Therefore, bootstrap (re-sampling with replacement) is applied to generate 200 artificial replicates of the dataset and for each replicate the threshold is estimated (Efron and Tibshirani, 1986; Davison and Hinkley, 1997). The 5%, 50% and 95%-5th, 50th, and 95th percentile of the estimates are reported as a range of threshold estimates. The subsequent post-processing steps of gap-filling and partitioning are then repeated using those different thresholds to propagate the uncertainty of uStar- u_* threshold estimation to derived quantities such as annual NEE, GPP and R_{eco}

2.2 Gap-filling methods

Filling of gaps in half-hourly NEE data is necessary to obtain complete time series for the calculation of daily averages or balances such as monthly or seasonal sums.

The After quality checks and u_* -filtering, the dataset of half-hourly fluxes after quality checks and uStar-filtering NEE fluxes may contain up to 50 percent gaps, sometimes (sometimes this fraction is even higher, depending on the site conditions). For the site-year-benchmark datasets used in this manuscriptas a benchmark, the percentage of gaps ranged from before u_* -filtering was on average of 32% before uStar-filtering to and after u_* -filtering 60% or and 48% after gap-filling for upper or lower uStar threshold estimaterespectively.

The gap-filling method implemented in for upper and lower u_* threshold estimate, respectively. Filling of gaps in half-hourly NEE data is necessary to obtain complete time series for the calculation of daily averages or balances such as monthly or seasonal sums. The following three gap-filling methods are implemented in REddyProc is the so-called marginal

distribution sampling (MDS) by (Reichstein et al., 2005b). In a comparison with gap-filling methods for net carbon fluxes by (Moffat et al., 2007), this method performed well for the different artificial gap scenarios ranging from single half-hours to several days. Due to its flexibility in dealing with missing meteorological input data and its fast and highly automated routines available as BGC online tool, the MDS gap-filling method has been widely used. The algorithm exploits the covariation of the fluxes with the meteorological variables and the temporal autocorrelation of the fluxes based on two basic methods, the (moving) look-up table approach (LUT) and the mean diurnal course (MDC), which are described next.

2.2.1 Look-up-table for gap-filling Look-up tables

In the look-up table (LUT) approach, the fluxes are binned by the meteorological conditions within a certain time window. Within the chosen time window and respective bin, each meteorological variable deviates less than a fixed margin to ensure similar meteorological conditions. The missing value of the flux is then calculated as the average value of the binned records and its uncertainty estimated from their standard deviation.

The original LUT of Falge et al. (2001) consisted of fixed periods over a year(Falge et al., 2001), while in the MDS algorithm, REddyProc the meteorological conditions are sampled with a moving window around the gap to be filled. Within the chosen time window and respective bin, each meteorological variable should not deviate more than a fixed margin to ensure similar meteorological conditions. The default meteorological variables are Rg, Tair, and VPD with default margins of 50, 2.5 °C, and 5.0, respectively.

2.2.2 Mean diurnal course for gap-filling

The NEE fluxes have a mean diurnal course (MDC) that follows the course of the sun with only respiration (release of) during nighttime and a combination of respiration and photosynthesis during daytime. This autocorrelation of the fluxes is exploited by taking the average value at the same time of day within a moving time window of adjacent days (Falge et al., 2001). In REddyProc, the same time of day includes also the fluxes of the adjacent hour (±1 hour).

Though the MDC method only showed a medium performance in the gap filling comparison for net carbon fluxes by (Moffat et al., 2007), it has the advantage that this approach can be used even if no meteorological information is available.

In the MDC algorithm, the same time of day includes also the fluxes of the adjacent hour (±1 hour). The number of adjacent days need to be specified.

Flow diagram of the MDS gap-filling algorithm as implemented in REddyProc. See table 1 for abbreviations.

2.2.3 Marginal distribution sampling for gap-filling

The so-called marginal distribution sampling combines the (MDS) by (Reichstein et al., 2005b) exploits the covariation of the fluxes with the meteorological variables and their temporal autocorrelation based on the two methods LUT and MDC methods depending described above.

- The filling of each half-hourly NEE with the MDS algorithm depends on the availability of the meteorological data . Three different conditions are identified for each half-hourly NEE flux:
 - The data of all three meteorological variables (of Rg, Tair, and VPD) are available. Tair or VPD are missing, but Rg is available. Also Rg is missing.
- Case 1): The missing value is replaced by the average value under similar meteorological conditions in a LUT approach. Similar meteorological conditions based on Rg, Tair and VPD. If no similar meteorological conditions (minimum of two half-hourly fluxes) are present within the starting time window of 7 days, the windows size is increased to 14 days.
 - Case: 1) If all three meteorological variables are available, LUT will be used with default margins of 50 Wm⁻², 2.5 °C, and 5.0 hPa, respectively. 2): The same LUT approach is taken, but similar meteorological conditions can only be defined via Rg within a time window of 7 days.
- Case If Tair or VPD are missing, only the variable Rg will be used. 3): The missing value is replaced with the mean diurnal course (MDC). The If no meteorology is available, the gaps are filled with MDC. Following a specific sampling procedure, the MDS algorithm increases the number of days start with one day, thus a linear interpolation of available data at adjacent hours (±1 hour) at the same day. The number is then increased to ±1 and ±2 days.
- If after these steps the NEE values could not be filled, the procedure is repeated with increased window sizes until the value can

 be filled, see flow diagram in Figure ?? in the vicinity of the gap until there are enough data points (at least two) for gap-filling.

 A more detailed description with a flow diagram is provided in Supplement 1.
 - The MDS algorithm is optimized for carbon dioxide and water fluxes and can also be used to estimate the uncertainty of the half-hourly fluxes. In the comparison of gap-filling methods by (Moffat et al., 2007), the MDS algorithm performed well for different artificial gap scenarios ranging from single half-hours to several days. Due to its flexibility in dealing with missing meteorological input data and its fast and highly automated routines available as online tool (BGC16, section 3), the MDS gap-filling method has been widely used.

2.3 Flux-partitioning methods

The gross fluxes of GPP into the land system and R_{eco} out of the land system are the two opposing parts of NEE: NEE = R_{eco} - GPP. Availability of GPP and R_{eco} is pivotal as they are the two biggest terms of the carbon cycle (e.g. Chapin et al., 2006; Jung et al., 2011). Moreover, understanding their sensitivity to environmental drivers (e.g., radiation, temperature, and soil moisture)

is important to interpret land-atmosphere interactions and to improve earth systems models (Reichstein et al., 2012). Therefore several methods were developed to partition NEE into these two components (Reichstein et al., 2005c; Lasslop et al., 2010; Moffat, 2012; Wehr and Saleska, 2015; Desai et al., 2008; Stoy et al., 2006).

The two most widely used methods are the so-called night-time partitioning and day-time partitioning (Reichstein et al., 2012). The night-time partitioning (Reichstein et al., 2005c) relies on the temperature response function of nighttime NEE fluxes that are representative of R_{eco}. It assumes that this relationship is applicable also to daytime data. The relationship is then used to predict R_{eco} from measured temperature and GPP is computed as a difference between R_{eco} and NEE. This method is currently the most widely used approach. Alternatively, the day-time partitioning (Lasslop et al., 2010) fits a LRC to daytime NEE observations model to observations of daytime NEE and global radiation, accounting for the effects of radiation and VPD on GPP as well as the effects of temperature on R_{eco}.

2.3.1 Nighttime flux-partitioning

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The method of (Reichstein et al., 2005c) Reichstein et al. (2005c) estimates a temporally varying respiration-temperature relationship from night-time data. First night-time data is selected by a threshold of $Rg < 10 \text{ Wm}^{-2}$, which is congruent with the BGC online tool (BGC16, section 3), but differs from the 20 Wm⁻² reported in (Reichstein et al., 2005c) Reichstein et al. (2005c).

Next, temperature sensitivity, E_0 of the Lloyd and Taylor (1994) relationship (1) is estimated by fitting the model to successive 15 day periods of night-time data, and the resulting E_0 series is aggregated to an annual estimate.

$$R_{eco}(T) = R_{Ref} \underbrace{exp}_{\sim} \left[E_0 \left(\frac{1}{T_{Ref} - T_0} - \frac{1}{T - T_0} \right) \right], \tag{1}$$

where T_0 is kept constant at -46.02°C Lloyd and Taylor (1994) (Lloyd and Taylor, 1994) and where reference temperature T_{Ref} is 15°C, which is congruent with the BGC online tool (BGC16, section 3) but differs from the 10°C reported in (Reichstein et al., 2005c). For robustness each fit is repeated on a trimmed data set excluding records with residuals outside the 5%-95% residual distribution. The annual aggregate is the mean across the 3 valid estimates having the lowest uncertainty in the fit. Single estimates of E_0 are considered valid, if there were minimum of 6 records, temperature ranged across at least 5°C, and estimates were inside the range of 30 to 450 K.

Subsequently, the respiration at reference temperature, R_{Ref} , is re-estimated from night-time data using the annual E_0 temperature sensitivity estimate for 7-day windows shifted consecutively for 4 days. The estimated value is then assigned to the central time-point of the 4 day period and linearly interpolated between periods. Hence, the obtained respiration-temperature relationship varies across time.

Finally, R_{eco} is estimated for both day- and night-time from the temporarily varying R_{eco} temperature relationship and daytime GPP is computed as R_{eco} - NEE.

2.3.2 Daytime flux-partitioning

The method of (Lasslop et al., 2010) models NEE using the common rectangular hyperbolic light-response curve (LRC) (Falge et al., 2001):

$$NEE = \frac{\alpha \beta R_g}{\alpha R_g + \beta} + \gamma, \tag{2}$$

where α (µmol CO_2 J $^{-1}$) is the canopy light utilization efficiency and represents the initial slope of the light-response curve, β (µmol CO_2 m $^{-2}$ s $^{-1}$) is the maximum CO_2 uptake rate of the canopy at infinite Rg, associated with light saturation, and γ (µmol CO_2 m $^{-2}$ s $^{-1}$) is a term accounting for ecosystem respiration. The hyperbolic light-response curve is modified to account for the temperature dependency of respiration after Gilmanov et al. (2003) by setting respiration γ to the Lloyd and Taylor respiration model (Lloyd and Taylor, 1994) (1). Further, the LRC setting-constant parameter β in (2) to is replaced by an exponential decreasing function (Körner, 1995) at higher VPD values (3).

$$\beta = \begin{cases} \beta_0 \exp\left[-k(\text{VPD} - \text{VPD}_0)\right] & \text{if VPD} > 10 \text{ hPa} \\ \beta_0 & \text{otherwise} \end{cases} , \tag{3}$$

where the VPD_0 threshold is 10 hPa in accordance with earlier findings at the leaf level (Körner, 1995) ignoring potential vegetation specific differences.

Parameter T_0 in (1) was fixed as in the nighttime approach night-time partitioning (section 2.3.1). Parameter T_{Ref} was fixed in each window to the median temperature within the window. The other parameters $(E_0, R_{Ref}, \alpha, \beta_0, k)$ of the model are estimated by the following steps. 1) A time varying temperature sensitivity E_0 is estimated from night-time data for a window shifted by two days. 2) The E_0 estimates are smoothed across successive windows by fitting a Gaussian Process (Rasmussen and Williams, 2006; Menzer et al., 2013) using the mlegp R-package that also estimates uncertainty of the smoothed E_0 . Next, a prior respiration, R_{Ref} , for reference temperature $T_{Ref} = 15$ °C is re-estimated from night-time data for each window with smoothed E_0 . 3) Parameters of the rectangular hyperbolic light-response curve $(R_{Ref}, \alpha, \beta_0, k)$ are fitted using only day-time data and the previously determined temperature sensitivity (E_0) for each window. 4) Finally, for each NEE record, GPP and R_{eco} are estimated with the parameter set of the previous valid window and the parameters of the next valid window, and the two results are interpolated linearly by the time difference to the window centers. The following paragraphs add-Supplement 1 reports necessary technical details to these steps.

In step 1, parameter E_0 is estimated for 12 day windows. Only records with temperature above -1 °C are valid for estimation. Reference temperature T_{Ref} in (1) is set to the median temperature of the window in order to decrease correlation between estimates of R_{Ref} and E_0 . A missing estimate is reported for non-valid windows with too few valid records (minNRecInDayWindow = 10), non-convergence of the fitting procedure, or an E_0 estimate outside the bounds 50,400. Missing estimates are filled during the smoothing in step 2.

In step 2, the Gaussian Process takes into account the uncertainty of the E_0 fit from night-time in each window. However, if the correlation of E_0 across subsequent windows is high, the uncertainty is reduced similar as with repeated measurements. The respiration R_{Ref} for windows where no fit could be obtained is set to the value from the previous valid window. Again, only records with temperature above -1 °C are used.

In step 3, fitting of other parameters is done for each window centered at the same record as the windows of step 1. By default the fit uses the same weak prior on parameters as the BGC online tool (Lasslop et al., 2010). The prior locations are k = 0.05, $\alpha = 0.1$, R_{Ref} = night-time estimate, and β = range of NEE values, i.e. the difference between 97% and 3% quantile. The prior uncertainty is $sd_k = 50$, $sd_\beta = 600$, $sd_\alpha = 10$, and $sd_{R_{Ref}} = 80$.

In each daytime fit, the NEE records are weighted according to their variance (Lasslop et al., 2010, eq. 5). In order to avoid unreasonable leverage of records with a very low estimate of NEE uncertainty, the records of the above 0.7 quantile of weights are associated the weight of the 0.7 quantile. This assigns low influence to records with high uncertainty, but avoids the problem of the high leverage with very low estimates of NEE-uncertainty.

There are certain quality criteria and fall-backs during the daytime fitting in order to obtain reasonable fits. If there are too few valid records (minNReeInDayWindow < 10), or the fitting did not converge, or VPD parameter k < 0, then the fit is repeated without the VPD effect, because often there are records where VPD is missing but other variables are available. If fitting did not converge or parameter estimate of α is larger than 0.22, the fit is repeated with α fixed to the last valid value of α from fits in previous windows. If there are still too few records or the fitting was not valid, a missing result is reported for the window. In addition a missing result is reported if estimated $\alpha < 0$ or $R_{Ref} < 0$ or $\beta_0 < 0$ or $\beta_0 > 250$, or if $\beta_0 > 100$ and at the same time estimated standard deviation $sd_{\beta_0} > = \beta_0$ (Table A1 in Lasslop et al., 2010). The Variance-Covariance matrix of the parameter uncertainty is estimated by bootstrapping the day-time fit. In each sample, the prescribed temperature sensitivity E_0 is drawn from a normal distribution with standard deviation estimated in step 2. In this way also the uncertainty of the night-time fit propagates to the uncertainty of the day-time parameters and subsequently to the inferred gross fluxes.

In step 4, variance of the flux estimates are computed based on the Variance-Contrivance matrices obtained in step 3 with each parameter estimate (Lasslop et al., 2010, eq. 6). The two standard deviations based either on the previous and subsequent valid estimates of the Variance-Contrivance matrices are linearly interpolated with respect to the time difference to the estimates.

Note, that contrary to the night-time based flux-partitioning, both GPP and R_{eco} are model predictions and do not add up exactly to observed NEE.

3 Benchmarking REddyProc post-processing steps

The post-processing steps implementations of REddyProc were benchmarked with eurrent widely used the post-processing tools widely used in the FLUXNET processing. Specifically, REddyProc (version 0.8.1) uStar-filtering u_* -filtering results were compared with results by a C-implementation from Dario Papale (Papale et al., 2006), here referred to as DP06. Results of REddyProc (version 1.1.2) gap-filling and flux-partitioning were compared with results obtained by the year-2015-version of the 2016-version web based tool provided by the institute for biogeochemistryMax Planck Institute for Biogeochemistry, Jena, best described in Reichstein et al. (2005a). The tool was accessed in 2016 (29.07.2016) and is hereafter referred as the old BGC online toolBGC16. Annually and monthly aggregated values, here, refer to the mean across all valid values in a month or a year-In the presence of large gaps, these values, which can differ from real annual or monthly budgets. The manuscript proceeds describing in the presence of large gaps. The first section describes the dataset used for benchmarking and thereafter sections for the benchmarking of each processing steps for each processing step implemented in the package. Within each of those sections the following sections for the processing steps, subsections describe differences in the code, report the results of benchmarking, and discuss the main consequences them. Supplement 1, additionally, provides more detailed results and statistics.

3.1 Dataset used for benchmarking

Data of twenty-five sites with open data policy of the LaThuile FLUXNET dataset³ were used for benchmarking. The sites are located in different climate zones and belong to a variety of plant functional types (Table 2) to guarantee testing different conditions (i.e. presence of snow, management such as cuts and crop rotation, sites disturbed) and ecosystem types (e.g. deciduous vs evergreen forests, grasslands and forests). Site data contained croplands). For each site the following variables were used: NEE already filtered for quality flags (Foken and Wichura, 1996), despiked and uStar u_x filtered (Papale et al., 2006), random error of NEE computed as described by Reichstein et al. (2005a), Tairand soil temperature (Tsoil), Tsoil, Rg, and VPD. Moreover, NEE time series before the uStar-filtering and the uStar u_x-filtering and the u_x data were downloaded from AMERIFLUX and the European Flux Database to test the uStar u_x threshold estimation. Finally, time series of gap-filled NEE (NEE_f), GPP partitioned with the night-time based method (GPP_{NT}) (Reichstein et al., 2005a) were downloaded from the LaThuile dataset, while GPP partitioned with the daytime method (GPP_{DT}) were computed with BGC online toolBGC16.

25 3.2 UStar-filtering u*-filtering: Benchmark-Benchmarking with DP06

Estimation of uStar us threshold by REddyProc using the default moving point method (section 2.1.1) was benchmarked to estimation based on Papale's DP06 C-implementation (Papale et al., 2006). The benchmark applied a boostrap bootstrap sample of size 60 and recorded lower, median, and upper quantiles of 10%, 50% and 90% instead of the default 5% and 95% based on a larger sample size to save computing time.

³www.fluxdata.org

The different estimates of the $uStar_{u_*}$ threshold have potential consequences for the inferred fluxes. To explore these consequences, we used the different thresholds to mark gaps, gap-fill the data, and compute the annual NEE based on the gap-filled time series. NEE uncertainty was estimated by the difference between NEE based on lower quantile $uStar_{u_*}$ and NEE based on upper quantile $uStar_{u_*}$ estimate.

5 3.2.1 Differences in code

The biggest difference of REddyProc compared to DP06 is that REddyProc by default employs seasons that can span across years. With the 'Within one year' classification option, which is employed also by DP06, records of December are associated to the same season as January and February of the same year. With the default 'continuous' classification, seasons start the same as in DP06 by default in March, June, September, and December. However, December is treated in the same season as January and February of the next year to avoid discontinuities at year boundaries. The annual uStar ux threshold is then applied according to those continuous seasons spanning year boundaries. For example, the processing of 2014 data would by default use data from Winter 2014 (starting in December 2013) to Autumn 2014 (ending in November 2014). REddyProc also allows more flexibility with the 'User specified' classification into seasons as explained below.

There are further slight differences between REddyProc and DP06. Both methods bin in a way such that the number of records in each bin is similar. If there are numerically equal uStar us values, they are sorted into the same bin, resulting in bins with unequal record numbers. In DP06, less and sometimes no records are sorted into the subsequent bins hampering the moving point detection. Contrary, the binning with REddyProc ensures that there are a minimum number of records in all bins. This often results in fewer bins than without numerically equal uStar us values. Moreover, differing from the Papale C-implementation, REddyProc employs several more quality criteria. First, when comparing the threshold bin to NEE in the following bins, it makes sure that there are least 3 bins to infer a plateau in NEE. Next, when aggregating the thresholds of different temperature classes to season, it ensures that a threshold was found in at least 20% of the temperature classes. For those seasons where no threshold could be determined, the annual estimate is used. When there are too few records within one year, a single season comprising all records is used for threshold estimation.

In difference to DP06, REddyProc re-samples data only within seasons instead across the entire year during the bootstrap, in order to protect periods of similar uStar-NEE us-NEE relationship and to avoid seasonal biases in re-sampling.

3.2.2 Benchmark results

The general relationship in the estimation of the <u>uStar u_* </u> threshold was retained (slope of 0.94) between the two methods (Fig. 3), although there was a big scatter. The exceptionally high threshold value of > 0.6 m s⁻¹ for site FR-Pue was very probably an overestimate by DP06. However, one has to remember that each estimate has a high uncertainty, and the differences between the two methods were in the range of this uncertainty (Appendix Fig. ??) . suppl 1) The estimate of the uncertainty of the <u>uStar u_* </u>

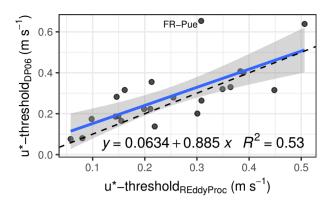


Figure 3. Large scatter but retained relationship between uStar ux threshold derived with DP06 versus uStar ux threshold derived with REddyProc across site-years as shown by a regression (solid line with shaded uncertainty bound) close to the 1:1 line (dashed).

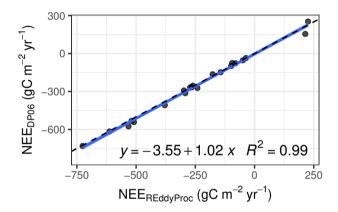


Figure 4. Strong correspondence in NEE based on uStar u* threshold estimated by REddyProc and NEE based on uStar u* threshold estimated by DP06 across site-years.

thresholds with REddyProc was, however, only half of the uncertainty range estimated by DP06 (Appendix Fig. ??). suppl 1). This increased precision was mainly due to the modified bootstrapping scheme, that respects the uStar us seasons.

When propagating the differences in uStar u_* to differences in annual NEE, there was no bias and very low scatter across sites between all the methods (Fig. 4), despite the differences in uStar u_* threshold. The absolute differences in annual NEE between the methods were small (mostly < 20 gC m⁻² yr⁻¹), and mostly lower than half of the uncertainty range estimated from the bootstrap (Appendix Fig. ??). suppl 1) REddyProc estimates uStar u_* thresholds with roughly double the precision compared to DP06, due to protecting seasons during bootstrap (Appendix Fig. ??). suppl 1).

3.2.3 Discussion of uStar- u_* threshold estimation

The agreement between NEE based on uStar u_* estimates of REddyProc moving point implementation and current FLUXNET standard post-processing (DP06) (Fig. 4) indicates that the sensitivity of NEE to the uStar u_* threshold estimate in the inferred ranges is low, which also explains the large uncertainty of the uStar u_* threshold estimate. One reason for the missing effect could be site selection of this study without many sites affected by advection, which show limited saturation of the NEE - u_* dependence. Since in such cases the filtering does not work properly anyway, it should not change the conclusions for NEE. Hence, we infer that u_* estimates of both DP06 and REddyProc are appropriate due to the negligible effect on NEE sums. The agreement implies that both methods can be interchanged in studies that are based on aggregated values, such as annual carbon budgets or for upscaling, without the need to reprocess data.

However, the increase in estimated precision, i.e. lower standard deviation, of the <u>uStar u</u> threshold estimate also yields an increase in estimated precision of the annual NEE (Fig. ??). suppl 1). This will lead to improved accuracy and usability of EC measurements and any downstream, post-processed data products in model-data integration studies.

While the default seasons and their aggregation are in line with previous approaches, REddyProc allows to use site specific knowledge to derive better threshold estimates. For example, if there is disturbance such as harvest, the u_* threshold is expected to change and a different threshold should be applied for filtering before and after the disturbance. In this case the user can define a season change at harvest date and use season-specific threshold estimates instead of the annually aggregated estimate (section B7).

3.3 Gap-filling: Benchmark with the BGC online toolBGC16

The gap-filling implementation of REddyProc was benchmarked with the BGC online tool (BGC16, section 3) that used pvWave-code from Reichstein et al. (2005c).

3.3.1 Differences in code

Compared to the BGC online toolBGC16, the new implementation of the MDS algorithm in REddyProc was not limited to single years but filled the gaps with a window moving continuously over all years in the input data. This had the advantage of smoother filling-gap-filling over the end of year time stamp timestamp and will especially be of interest for sites which are in which vegetation is not dormant during this time. This new feature led to differences in and different, probably more realistic gap-filled NEE values in the beginning and end of the year.

There were also slight differences in the window size between the old and new code. For MDC, the old window size had a few more intermediate day steps than the new implementation which affected longer gaps with missing meteorology. The default meteorological variables and margins for LUT (see Chapter 4.2.2 above) were the same in both implementations.

While REddyProc restricted restricts gap-filling to interpolation of gaps, the BGC online tool BGC16 also extrapolated missing records in periods without measurements.

3.3.2 Benchmark results and discussion for gap-filling

In the benchmark, REddyProc gap-filling was run using the same measured NEE as input that passed the QA/QC routines and $\frac{\text{uStar-filtering}}{u_*}$ -filtering. The annually aggregated values comprised both, filled gaps and originally valid records.

REddyProc gap-filling results agreed with the results from the BGC online toolof BGC16. Few discrepancies at half-hourly time scale were mostly at longer gaps due to usage of fewer window sizes, as shown for the DE-Tha case (Fig. 5 top). At annually aggregated time scale, the agreement between methods was strong ($R^2 = 0.99$) (Fig. 5 bottom). The outlier of site RU-Cok is due to the availability of only a few months of data for the whole year. While REddyProc filled gaps in the time period with data available, the old BGC online tool BGC16 extrapolated also into the time before and after. The seasonal cycle was well reproduced at each site (Appendix ??). suppl 1).

3.3.3 Discussion of gap-filling

The good agreement between NEE based on gap-filling by REddyProc and the BGC online tool gap-filling by BGC16 (Fig. 5) imply that these both gap-filling tools can be used interchangeably without the need to reprocess data.

15 3.4 Nighttime flux-partitioning: Benchmark with BGC online toolBGC16

The night-time based flux-partitioning was benchmarked to the BGC online tool (BGC16, section 3), that used pvWave code developed by Reichstein et al. (2005c).

3.4.1 Differences in code

The main features of the REddyProc implementation of the night-time based partitioning algorithm were very similar to the BGC online toolBGC16, with using a reference temperature of 15°C and trimming the estimates of temperature sensitivity E_0 before aggregating them (section 2.3.1). REddyProc differed from the BGC online tool BGC16 in computing the potential radiation that is used in subsetting the night-time data to derive E_0 and R_{Ref} (Reichstein et al., 2005). While REddyProc used the exact solar time for the calculation of the potential radiation, where the sun culminates exactly at noon, the BGC online tool BGC16 used the local winter time which differs from the solar time depending on the location within the time zone.

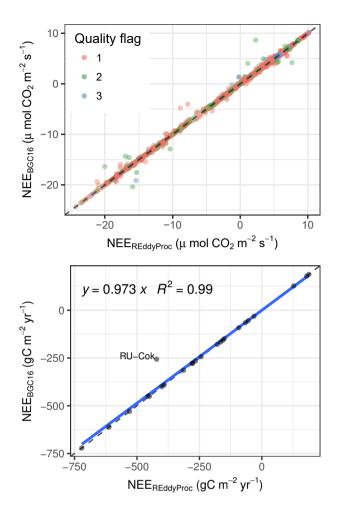


Figure 5. Predictions of NEE by REddyProc after gap-filling agree with the BGC online tool_BGC16 both at half-hourly values (top), shown for the DE-Tha 1998 use case, and annual means across sites (bottom). Larger quality flags are associated with larger window sizes.

3.4.2 Benchmark results and discussion for night-time flux-partitioning

Annual aggregated values of R_{eco} predicted by REddyProc were in very good agreement ($R^2 = 0.99$; slope ≈ 1) with the BGC online tool-BGC16 as shown in Fig. 6 and in the Appendix ??. supplement 1.

In order to evaluate the effects of the differences introduced in the code described above, we also computed R_{eco} by prescribing in REddyProc either E_0 , selection of night-time data, or both from the BGC online toolBGC16 output. Results are reported in the Appendix ?? supplement 1 and showed that the most important factor affecting the R_{eco} computed with REddyProc was the different selection of night-time data, though the differences were almost negligible at annual time scale(Appendix ??).

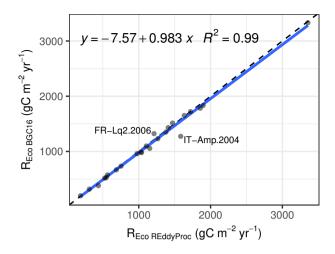


Figure 6. Predictions of annually aggregated ecosystem respiration, R_{eco} from REddyProc night-time partitioning agree with the predictions by the BGC online toolBGC16.

3.4.3 Discussion of nighttime flux-partitioning

The two implementations agree very well for most sites at annual time-scale. Because of no systematic deviations across sites, the spatial upscaling of fluxes should not be affected by REddyProc implementation. However, for some sites, such as IT-Amp, the quite large relative errors indicates problems related to selection of night-time data and problems due to a large gaps in a the dataset.

3.5 Daytime flux-partitioning: Benchmark with **BGC** online toolBGC16

The day-time flux-partitioning was benchmarked with results of the BGC online tool . The online tool used (BGC16, section 3), which is based on pvWave code developed by Lasslop et al. (2010) that was used with slight variations also and used in the processing of the 2015 Fluxnet release (Pastorello et al., 2017).

10 3.5.1 Differences in code

The BGC online tool BGC16 differed from REddyProc (section 2.3.2) mainly in aspects of separation of nighttime data, estimation of temperature sensitivity from night-time data, uncertainty estimation, treatment of missing values, and optimization library code.

While for separating nighttime data REddyProc used exact solar time where sun culminates exactly at noon, the BGC online tool BGC16 used the local winter time zone time.

For the estimation of temperature sensitivity E_0 from night-time data the BGC online tool BGC16 used a reference temperature of 15 °C, instead of the median temperature inside the window. Hence, it estimated stronger correlations between parameters for windows with a different temperature range. Moreover, it omitted smoothing of the estimated E_0 across time, often leading to large fluctuations of the E_0 estimates across few days (Fig. ??) suppl 1), larger estimates of its uncertainty, and differences in subsequent estimation of LRC parameters.

For uncertainty estimation, the BGC online tool BGC16 relied on the curvature of the LRC fit's optimum instead of a bootstrap procedure. Hence, it could not take into account the uncertainty of E_0 estimated from night-time data before the daytime day-time LRC fit. Moreover, during interpolation of fluxes based on previous and next valid estimates, the distance weights differed. While REddyProc assigned assignes the estimates to the time of the mean of valid record in a window, the BGC online tool BGC16 assigned it to the start of the third day, also if there were only valid data for the first day in the window.

For weighting the records in the LRC fit, the BGC online tool BGC16 used the raw estimated NEE uncertainty of each record. It did not check for high leverage of spurious low NEE uncertainty estimates. Its estimates, therefore, were in some windows very strongly influenced by a few records, and failed if a NEE uncertainty estimate of zero was provided. Moreover, when there were missing values or values below zero in provided NEE uncertainty, it set all uncertainty to 1, while REddyProc filled fills the gaps by setting the missing uncertainty to the maximum of 20% of respective NEE but at least 0.7 µmolCO₂m⁻²s⁻¹.

Treatment of missing values was not considered by the BGC online tool_BGC16 and assumed to be handled prior to the processing. Hence, it did not handle missing VPD values and did not re-try the LRC fit without the VPD-effect in order to use also records with missing VPD. Moreover, as described above, when there were missing values of NEE uncertainty, weighting records in the LRC fit was omitted.

For compatibility with the BGC online toolBGC16, the above code-differences can be disabled in REddyProc. But differences in optimization library code and specifically the conditions of non-convergence on scattered data could not be eliminated, which led to differences in results as shown in the following section.

3.5.2 Benchmark results for day-time partitioning

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Annually GPP predictions of both implementations showed no significant bias across the test sites (Fig. 7), although, there was some scatter for the individual predictions. A similar scatter was observed when comparing the predictions of the default REddyProc options to the predictions with compatibility options. Most of the scatter was caused by skipping the test on differences were caused by decreasing the unreasonable high influence of NEE records with small NEE uncertainty (Appendix Fig. ??). suppl 1).

The largest differences in aggregated fluxes between implementations were due to the extrapolation of fitted parameters to periods where no parameter fits were obtained. In many of those cases, there were fits at the boundaries of the difficult these periods, whose validity was questionable. Whether those fits passed the quality check or not had a large influence on the

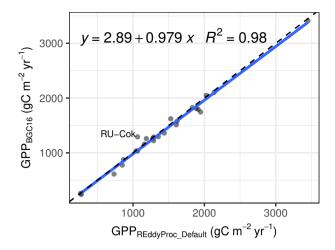


Figure 7. Prediction of annually aggregated GPP from REddyProc day-time partitioning agree with the BGC online tool_BGC16 across sites.

extrapolation and hence on the aggregated values. For example, at RU-Cok parameter estimates for valid periods agreed between implementations. However, no valid parameters could be obtained for winter months. While REddyProc reported missing values, the BGC online tool BGC16 reported GPP values based on summer parameterizations also for periods further away from summer, which in turn lead to higher annual GPP estimates.

Uncertainty estimates of gross fluxes were larger with REddyProc due to accounting for uncertainty in temperature sensitivity estimate stimates from night-time data (Fig. ??). suppl 1).

3.5.3 Discussion of daytime day-time flux-partitioning

Agreement between aggregated fluxes predicted by the day-time method and absence of bias for the test sites (Fig. 7) suggest that the methods can be used interchangeably for upscaling, although differences in results of influential sites can potentially propagate to differences in upscaled estimates. REddyProc provides a quality flag for the results of the day-time partitioning, that allows excluding less reliable data in upscaling studies. For the results associated with good quality flags, we set stronger trust in the REddyProc-based estimates.

The day-time flux-partitioning is quite sensitive to the details of LRC fit. Small changes in treatment of extreme or missing NEE uncertainty estimates or changes in pre-processing and treatment of missing values cause different estimates of LRC parameters and propagate to predicted fluxes of GPP and R_{eco}. Although, we put much effort in trying to reproduce the BGC online tool results results of BGC16, we were not able to eliminate all differences, especially in the subtle details in the parameter optimization library codes. The differences in predicted half-hourly fluxes, however, average out across sites and across time (Fig. ?? bottom) suppl. 1) making this issue less severe at larger scales.

The estimated uncertainties are even more sensitive. Both implementations occasionally produce unreasonably high outliers that affect the aggregated values. REddyProc, in general, estimates higher uncertainties of predicted fluxes, because it accounts for uncertainty in temperature sensitivity. Note, that the uncertainty introduced to annually aggregated fluxes due to flux-partitioning is smaller than uncertainty due to uncertain uStar us threshold estimate. Hence, differences or difficulties in uncertainty estimation caused by flux-partitioning do affect conclusions of the overall uncertainty estimates to a lesser extent.

4 Conclusions

The REddyProc software provides a set of tools for the CO_2 -focussed post-processing of eddy covariance flux data including uStar-filtering, gap-filling and flux-partitioning. The freely available R-package allows flexible integration into extended workflows and automated routines for the and propagation of the uncertainty from the uStar-filtering u_* -filtering to the gap-filled NEE and partitioned GPP and R_{eco} .

The freely available R-package enables researchers to integrate the flux data processing into their own, offline environment or work stream without the need of uploading data. This seamless integration allows to improve overall workflow, to speed up processing routines and to ultimately generate cleaner, reproducible scientific results.

The compatibility of the implemented methods to the available standard tools provides continuity of the data analysis, when adopting REddyProc for processing EC-data. REddyProc can closely reproduce results of the widely used BGC online tool (BGC16, section 3).

A number of enhancements provide more flexibility to the user in the processing of their data. For instance, the new processing allows to treat multi-year data without breaks at annual boundaries that can significantly affect sites in the southern hemisphere or sites characterized by vegetation activity in winter. Another new feature of REddyProc is the flexibility to define different seasons for the application of the uStar-filtering ux-filtering and gap-filling routines, which is important for sites with discontinuous surface cover associated with snow melt, dry seasons, or harvest.

Sensitivity of the results to subtle details of the implementation, however, call for caution when interpreting results. This is especially true for $uStar u_x$ threshold estimation and the daytime-based day-time flux-partitioning, and especially for data with long gaps.

Continued integration of new methodological developments into the package will support research using EC data. We strive to provide new developments in a basic and extensible manner, while paying attention to compatibility with results of reference implementations.

In summary, research using (half-)hourly eddy covariance data can benefit from building blocks for standardized and extensible post-processing provided by REddyProc.

Appendix A: The REddyProc package

The REddyProc processing tool is freely available in two options: a) online as a web-service⁴ with a smaller range of user option, and b) as a package of the open-source R environment with a larger set of user options and with each of the steps and methods available independently.

The REddyProc package can be installed by typing at the R-terminal:

```
install.packages("REddyProc")
library(REddyProc)
?REddyProc
```

The package relies on quasiquotation and requires a quite recent versions of the tidyverse packages, specifically dplyr version 0.7. Installations with older versions should rund the following code before installing.

install.packages("tidyverse") update.packages(oldPkgs="dplyr")

10 A1 General design

There are some general principles and choices in the design of that explain some trade-offs.

Preventing accidental errors is one goal of the package. Therefore Alternatively, there is extensive checking on formatting and data availability when importing the data and also throughout post-processing steps. Especially, the high-level routines will issue warnings or stop post-processing if they detect inconsistencies or lack of sufficient data, or changes in specification of critical standard parameters. These checks reinforce a sound standard post-processing also for non-expert users. On the other hand, these checks will render the standard routines not usable for datasets where not enough data are available. It is still possible to use—with sparse data for some purposes by using lower-level routines. But this requires experience in both, post-processing and R programming.

Smooth inter-annual processing is the next goal. The data is not partitioned into annual chunks for post-processing to avoid artificial discontinuities between years. The user can still specify different periods or seasons, e.g. when meteorological conditions change after harvest, but the boundaries do not need to align with years. therefore works with the entire dataset in memory. Potentially this can lead to longer post-processing times when working with many years of EC data but still can be handled on a usual notebooka automatically built docker image (Metzger et al., 2017) at docker hub allowing to run RStudio with REddyProc from a browser without any other installation besides docker. Regarding installation issues and docker images we refer the reader to the github project homepage⁵.

⁴ www.bgc-jena.mpg.de/bgi/index.php/Services/REddyProcWeb

⁵github.com/bgctw/REddyProc

Memory efficient processing required some extended R programming. Specifically we used R5 classes⁶ to avoid frequently copying the entire dataset in memory. Users should be aware that calling functions on the not only provides a return value but also changes the data of Some general principles and choices in the R5 class (Chapter 'OO field guide' in Wickham, 2014). Users who want to integrate the post-processing in their own codes are encouraged to learn about R5 classes, however, it is not a prerequisite.

Continuity with other tools ensures that switching to design of REddyProc does not introduce discontinuities with results obtained from other tools. Specifically, we reproduced the uStar threshold estimation of the C implementation by Dario Papale, and the gap-filling, night-time flux partitioning, and day-time flux-partitioning from the BGC online tool (Reichstein et al., 2005). In some details, the standard parameterization differs from the old tool, e.g. using seasons that span across calendar years in uStar threshold estimation, but it is usually possible to set compatible parameters with that lead to trade-offs between robustness and flexibility are explained in supplement 1.

⁶www.rdocumentation.org/packages/methods/versions/3.4.3/topics/ReferenceClasses

Table 1. Abbreviations used repeatedly in the paper.

Symbol	Description	
EC		
	eddy-covariance	
	eddy	
	covariance	
CO_2	carbon	
	diox-	
	ide	
NEE	net	
	ecosys-	
	tem	
	ex-	
	change	
	to-	
	wards	
	the	
	at-	
	mo-	
	sphere	
	in	
	$\mu \mathrm{mol}\mathrm{CO}_2\mathrm{m}^{-2}\mathrm{s}^{-2}$	
	(ag-	
	gre-	
	gated	
	in	
	$gC m^{-2}yr^{-1})$	
GPP	gross	
	pri-	
	mary	
	pro-	
	duc-	
	tiv-	
	ity	
	(same	
	units	
	as	
	NEE)	
R _{eco} 24	ecosystem	
Neco 27	res-	

pira-

Table 2. Description of the site sites and site year times used for benchmarking REddyProc.

^b Abbreviations for climate from Köppen-Geiger classification: Af: equatorial, rainforest; BSh: hot arid steppe; Cfa: humid, warm temperate, hot summer; Cfb: humid, warm temperate, warm summer; Csa: summer dry, warm temperate, hot summer; Dfb: cold, humid, warm summer; Dfc: cold, humid, cold summer.

Site	Year	Lat	Lon	Land cover ^a	Climate ^b
CA-NS7	2004	56.64	-99.95	OSH	Dfc
CA-TP3	2005	42.71	-80.35	ENF	Dfb
CH-Oe2	2004	47.29	7.73	CRO	Cfb
DE-Hai	2002	51.08	10.45	DBF	Cfb
DE-Tha	1998	50.96	13.57	ENF	Cfb
DK-Sor	2006	55.49	11.65	DBF	Cfb
ES-ES1	2000	39.35	-0.32	ENF	Csa
ES-VDA	2005	42.15	1.45	GRA	Cfb
FI-Hyy	1998	61.85	24.29	ENF	Dfc
FI-Kaa	2001	69.14	27.30	WET	Dfc
FR-Gri	2006	48.84	1.95	CRO	Cfb
FR-Hes	1998	48.67	7.06	DBF	Cfb
FR-Lq1	2006	45.64	2.74	GRA	Cfb
FR-Lq2	2006	45.64	2.74	GRA	Cfb
FR-Pue	2003	43.74	3.60	EBF	Csa
IE-Dri	2004	51.99	-8.75	GRA	Cfb
IL-Yat	2005	31.34	35.05	ENF	BSh
IT-Amp	2004	41.90	13.61	GRA	Cfa
IT-MBo	2005	46.02	11.05	GRA	Cfb
IT-SRo	2001	43.73	10.28	ENF	Csa
PT-Esp	2004	38.64	-8.60	EBF	Csa
RU-Cok	2004	70.62	147.88	OSH	Dfc
SE-Nor	1997	60.09	17.48	ENF	Dfb
US-Ton	2004	38.43	-120.97	WSA	Csa
VU-Coc	2002	-15.44	167.19	EBF	Af

^a Abbreviations for land cover type from International Geosphere-Biosphere Programme (IGBP) classification: CRO: cropland, DBF: deciduous broadleaf forest, EBF: evergreen broadleaf forest, ENF: evergreen needleleaf forest, GRA: grassland, OSH: open shrubland, WET: permanent wetland, WSA: woody savanna.

Appendix B: Example application

This section reports an example R session using REddyProc. Code is shown in a shaded area and corresponding output with monospace font.

B1 Importing the half-hourly data

The workflow starts with importing the half-hourly data. The example reads a text file with data of the year 1998 from the Tharandt-DE-Tha site and converts the separate decimal columns year, day, and hour to a POSIX timestamp column. Next, it initializes the sEddyProc class.

10 B2 Estimating the uStar u_* threshold distribution

The second step is the estimation of the distribution of $uStar u_*$ thresholds to identify periods of low friction velocity ($uStar u_*$), where NEE is biased low. Discarding periods with low $uStar u_*$ is one of the largest sources of uncertainty in aggregated fluxes. Hence, several quantiles of the distribution of the uncertain $uStar u_*$ threshold are estimated by a bootstrap.

The friction velocity, uStaru, needs to be in a column named "Ustar" of the input dataset.

```
uStarTh <- EddyProc.C$sEstUstarThresholdDistribution(
  nSample = 100L, probs = c(0.05, 0.5, 0.95))</pre>
```

```
#filter(uStarTh, aggregationMode == "vear")
   select (uStarTh, -seasonYear)
                                                    5%
                                                             50%
                                                                        95%
   ##
        aggregationMode
                         season
                                      uSt.ar
   ## 1
                  single
                            <NA> 0.4162500 0.3658125 0.4414583 0.6197520
   ##
                            <NA> 0.4162500 0.3658125 0.4414583 0.6197520
                    vear
                  season 1998001 0.4162500 0.3658125 0.4414583 0.6197520
   ##
      3
                  season 1998003 0.4162500 0.3198846 0.4037981 0.5615500
   ##
     4
                  season 1998006 0.3520000 0.3192476 0.3858571 0.4393708
   ##
     5
                  season 1998009 0.3369231 0.2195485 0.3906458 0.5235175
   ## 6
10
   ## 7
                  season 1998012 0.1740000 0.1829667 0.4189286 0.6180982
```

The output reports uStar us estimates of 0.42 for the original data and 0.37, 0.44, 0.62 for lower, median, and upper quantile of the estimated distribution. The threshold can vary between periods of different surface roughness, e.g. before and after harvest. Therefore, there are estimates for different time periods of the year, called seasons, reported as different rows. These season-estimates can be aggregated to entire years or to a single value across years, reported by rows with corresponding aggregation mode.

The subsequent post processing steps will be repeated using the three quantiles of the uStar u_* distribution. They require to specify a uStar-threshold u_* -threshold for each season and a suffix to distinguish the outputs related to different thresholds.

For this example of an evergreen forest site, the same annually aggregated uStar_u* threshold estimate will be chosen for each of the seasons within a year. In order to distinguish the automatically generated columns, the column names of the estimation results are written to variable uStarSuffixes.

```
uStarThAnnual <- usGetAnnualSeasonUStarMap(uStarTh) [-2]
uStarSuffixes <- colnames(uStarThAnnual) [-1]
print(uStarThAnnual)</pre>
```

```
## season U05 U50 U95
## 1 1998001 0.3658125 0.4414583 0.619752
## 2 1998003 0.3658125 0.4414583 0.619752
## 3 1998006 0.3658125 0.4414583 0.619752
## 4 1998009 0.3658125 0.4414583 0.619752
## 5 1998012 0.3658125 0.4414583 0.619752
```

15

B3 Gap-filling

The second post-processing step is filling the gaps using information of the valid data. In this case, the same annual $uStar u_*$ threshold estimate is used for each season, as described above, and the uncertainty will be computed also for valid records (FillAll).

```
EddyProc.C$sMDSGapFillAfterUStarDistr('NEE',
   UstarThres.df = uStarThAnnual,
   UstarSuffix.V.s = uStarSuffixes,
   FillAll = TRUE
)
```

The screen output (not shown here) already shows that the uStar-filtering u_* -filtering and gap-filling was repeated for each given estimate of the uStar u_* threshold, i.e. column in uStarThAnnual, with marking 22% to 38% of the data as gap.

For each of the different $uStar_u$, threshold estimates, a separate set of output columns of filled NEE and its uncertainty is generated, distinguished by the suffixes given with uStarSuffixes. Suffix "_f" denotes the filled value and "_fsd" the estimated standard devation of its uncertainty.

```
grep("NEE_.*_f$",names(EddyProc.C$sExportResults()), value = TRUE)
grep("NEE_.*_fsd$",names(EddyProc.C$sExportResults()), value = TRUE)

## [1] "NEE_U05_f" "NEE_U50_f" "NEE_U95_f"

## [1] "NEE_U05_fsd" "NEE_U50_fsd" "NEE_U95_fsd"
```

B4 Partitioning net flux into GPP and R_{eco}

The third post-processing step is partitioning the net flux (NEE) into its gross components GPP and R_{eco}. The partitioning algorithm needs a precise criteria between night-time and day-time. Therefore, the specification of geographical coordinates and the time zone need to be provided to allow computing exact solar time of sunrise and sunset. Further, the missing values in the used meteorological data need to be filled.

```
EddyProc.C$sSetLocationInfo(Lat_deg.n = 51.0, Long_deg.n = 13.6, TimeZone_h.n = 1)
EddyProc.C$sMDSGapFill('Tair', FillAll.b = FALSE)
EddyProc.C$sMDSGapFill('VPD', FillAll.b = FALSE)
```

Now we are ready to invoke the partitioning, here by the night-time approach, for each of the several filled NEE columns.

```
#variable uStarSuffixes was defined above at the end of uStar threshold estimation
resPart <- lapply(uStarSuffixes, function(suffix){</pre>
```

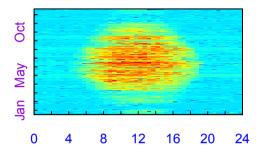
```
EddyProc.C$sMRFluxPartition(Suffix.s = suffix)
})
```

The results are stored in columns Reco and GPP_f modified by the respective uStar-u* threshold suffix.

```
grep("GPP.*_f$|Reco", names(EddyProc.C$sExportResults()), value = TRUE)
## [1] "Reco_U05" "GPP_U05_f" "Reco_U50" "GPP_U50_f" "Reco_U95" "GPP_U95_f"
```

5 The visualizations of the results by a fingerprint plot gives a compact overview.

```
EddyProc.C$sPlotFingerprintY('GPP_U50_f', Year.i = 1998)
```



B5 Estimating the uncertainty of aggregated results

First, the mean of the GPP across all the year is computed for each uStar-scenario u_* -scenario and converted from μ mol CO_2 m⁻²s⁻¹ to gC m⁻²yr⁻¹.

```
FilledEddyData.F <- EddyProc.C$sExportResults()
#suffix <- uStarSuffixes[2]

GPPAggCO2 <- sapply( uStarSuffixes, function(suffix) {
        GPPHalfHour <- FilledEddyData.F[[paste0("GPP_",suffix,"_f")]]
        mean(GPPHalfHour, na.rm = TRUE)
})

molarMass <- 12.011

GPPAgg <- GPPAggCO2 * 1e-6 * molarMass * 3600*24*365.25

print(GPPAgg)</pre>
```

```
10 ## U05 U50 U95
## 1894.097 1956.090 1985.817
```

The difference between those aggregated values is a first estimate of uncertainty range in GPP due to uncertainty of the uStar u_* threshold.

```
(max(GPPAgg) - min(GPPAgg)) / median(GPPAgg)
```

In this run of the example a relative error of about 4.7% is inferred.

For a better but more time consuming uncertainty estimate, specify a larger sample of uStar us threshold values, for each repeat the post-processing, and compute statistics from the larger sample of resulting GPP columns. This can be achieved by specifying a larger sequence of quantiles when calling sEstUstarThresholdDistribution in section B2.

```
%DIFDELCMD < sEstUstarThresholdDistribution%%%
EddyProc.C$sEstUstarThresholdDistribution (
nSample = 200, probs = seq(0.025,0.975,length.out = 39) )</pre>
```

B6 Storing the results in a csv-file

The results still reside inside the sEddyProc class. To export them to an R Data.frame, the newly generated columns need to be appended to the columns with the original input data. Then this data.frame is written to a text file in a temporary directory.

```
FilledEddyData.F <- EddyProc.C$sExportResults()
CombinedData.F <- cbind(EddyData.F, FilledEddyData.F)
fWriteDataframeToFile(CombinedData.F, 'DE-Tha-Results.txt', Dir.s = tempdir())</pre>
```

10 B7 Specifying seasons where $\frac{uStar}{u_*}$ thresholds differs

With changing surface roughness, e.g. on harvest or leaf-fall, also the uStar-NEE u_x -NEE relationship can change. Therefore the uStar u_x threshold needs to be re-estimated at different times of the year, called seasons. The default uses continuous seasons, for details see section 3.2.1. In order to yield results corresponding to DP06, the user can specify seasonFactor.v = usCreateSeasonFactorMonthWithinYear(EddyData.C\$sDATA\$sDateTime, startMonth= c(3,6,9,12)) as an argument to routine sEstUstarThreshold. By default the annual aggregate of the season thresholds, i.e. maximum across seasons, is used to identify unfavorable conditions. The, but the seasonal estimates can also be used instead.

Moreover, the users can specify also other user-defined seasons, e.g. when harvest dates are known (see package vignette DEGebExample). They can create a grouping by specifying exact starting days of the periods by function usCreateSeasonFactorYdayYear, or they can provide a column with the data that indicates e.g. the same group for two wet seasons. Each season is associated to the year corresponding to the center day between first and last day of the season.

With all methods, there is a required minimum number of 160 records within a season. If there are too few records, the data of the seasons within one year are combined and the uStar u_x threshold for this seasons is set to the estimate obtained for the data of the entire year.

Appendix C: Additional benchmark statistics

This section provides figures, in addition to the main figures presented in the benchmark section (3), that require the reader being comfortable with histograms and probability distribution functions.

The histograms display how often a certain value occurs across all the site-months. A centering of a difference away from zero or for a ratio away from one denotes a bias.

B1 uStar Threshold estimation

There was a bias in uStar estimation (Fig. ??). However, the bias and the differences were only a small fraction of the uncertainty of the estimate itself. Moreover, the bias did not propagate through estimated NEE based on different uStar values (Fig. ??).

Histogram of differences between uStar threshold estimates of different methods (DP06 -) normalized by the uncertainty range (90% quantile - 10% quantile estimated by DP06). Absolute values are smaller than one meaning that the difference between the methods is smaller than range of estimates by the DP06 method only. The clustering of values on the positive side suggests a bias towards larger values with DP06.

Histogram of differences between annual NEE based on uStar estimates of different methods (Papale -) normalized by the uncertainty range of NEE due to uStar (NEE based on 90% uStar quantile - NEE based on 10% uStar quantile estimated by DP06). Absolute values are mostly smaller than one, meaning that the difference in annual NEE between methods is smaller than the uncertainty due to uncertainty of uStar threshold from DP06 only.

There was a significant difference in the estimate of uncertainty of the uStar thresholds (Fig. ??). only estimated half of the uncertainty due to acknowledging the seasons of similar conditions also during bootstrap. This lower uncertainty propagated to the uncertainty estimates of NEE (Fig. ??)

Histogram of ratio (/ DP06) of uncertainty ranges of uStar (90% quantile - 10% quantile). The clustering around 0.5 shows that the estimated uncertainty with _ is only about half the uncertainty estimated by DP06.

Histogram of ratio (/ DP06) of uncertainty ranges of annual NEE (NEE based on 90% quantile of uStar threshold - NEE based on 10% quantile of uStar threshold). The clustering below a value of one indicates that the lower uncertainty in uStar threshold with also propagates to lower uncertainty in annual NEE.

B1 Gap-filling

20

The evaluation of annual aggregated NEE data obtained with and the BGC online tool showed good agreement across sites.

Note, that the aggregated NEE data contained both, measured and gap-filled data for the purpose of evaluating the impact of the processing on the aggregated NEE. The determination coefficient (R2) showed a very good agreement between the two

methods both at annual and monthly time scale (Table ??). The relative mean absolute error (RMAE) is low: about -3 % for annual aggregation and -0.47 % for monthly aggregation. Both Modeling Efficiency (EF) and Mean Bias Error (MBE) showed a very low bias between the two products for both monthly and annual aggregations (Table ??).

Gap-filling evaluation statistics of yearly (Fig. 5) and monthly cumulated and gap-filled NEE data obtained with and the BGC online tool. Statistic abbreviations are explained in section ??. Yearly Monthly N 25 281 pearson 0.99 1.00 MBE -0.02 0.00 RMBE 2.8% -0.08% MAE 0.02 0.00 RMAE -3.0% -0.5% RMSE 0.09 0.01 R2 0.98 1.00 EF 0.98 1.00

B1 Night-time partitioning

- The results show good agreement between estimated using and the BGC online tool for R2 and EF (Table ??). The relative RMSE is 6.56 and 13.45 % for yearly and monthly aggregation, respectively. The high RRMSE is due to few site years as reported in Table ??, and this is confirmed by the lower RMAE (3.87 % and 6.33 % for yearly and monthly, respectively), which is less sensitive to outliers.
- The difference in related to the selection of night-time data are not negligible: the differences in RRMSE of 1.03% and negligible differences in R2. Also, the use of E_0 prescribed from the BGC online tool lead to negligible difference in R2 of about 0.02. Therefore, though very small, the selection of night-time data is the most important difference introduced by .
 - Nighttime partitioning evaluation statistics across sites of annually (Fig. 6) and monthly aggregated ecosystem respiration () estimated with and the BGC online tool. Yearly Monthly N 25 297 pearson 0.99 0.99 MBE 25.4 2.15 RMBE 2.2% 2.2% MAE 44.8 6.17 RMAE 3.8% 6.3% RMSE 75.6 13.2 RRMSE 6.5% 13.5% R2 0.99 0.98 EF 0.99 0.97
- 20 RMBE RMAE RRMSE R2 EF CA-NS7 15.12 16.29 32.64 0.96 0.86 CA-TP3 -1.99 25.10 38.59 0.87 0.82 DE-Hai 6.00 6.48 10.40 0.99 0.96 DE-Tha 2.83 3.52 6.19 0.99 0.99 DK-Sor 3.50 4.45 7.34 1.00 0.99 ES-ES1 1.68 3.34 4.59 0.93 0.92 ES-VDA 3.96 9.98 15.93 0.95 0.93 FI-Hyy 2.72 3.05 4.47 1.00 1.00 FI-Kaa -4.52 5.51 9.07 0.99 0.99 FR-Gri 3.96 4.14 6.35 1.00 0.99 FR-Hes 4.11 5.06 9.06 0.99 0.99 FR-Lq1 -2.89 5.76 7.24 0.99 0.99 FR-Lq2 -8.27 9.49 11.49 0.99 0.97 FR-Pue 0.65 1.27 1.55 1.00 1.00 IE-Dri -0.03 2.41 3.24 1.00 1.00 IL-Yat 1.66 2.89 3.72 0.99 0.99 IT-Amp 24.34 29.38 45.33 0.91 0.48 IT-MBo 2.03 3.00 4.49 1.00 1.00 IT-SRo -1.46 1.86 2.10 1.00 0.99 PT-Esp 9.06 12.08 19.94 0.62 0.45 RU-Cok -0.76 28.54 38.98 0.27 -0.07 SE-Nor -1.05 2.09 3.62 1.00 1.00 US-Ton 2.93 7.45 11.82 0.93 0.92 VU-Coc 0.73 2.64 3.83 0.93 0.93 Nighttime partitioning evaluation statistics at sites level of the of ecosystem respiration () estimated with and the BGC online tool.

B1 Day-time partitioning

The time-variable estimate of temperature sensitivity of ecosystem respiration with day-time partitioning is a significant source of uncertainty for gross fluxes GPP and . accounts for this previously unaccounted uncertainty for estimating uncertainty of these gross fluxes by a bootstrap (Fig. ??). The estimate of annual uncertainty in Fig. ?? is a low estimate, because it assumed no correlation between half-hourly errors. An improved quantification of correlations requires the full variance-covariance

matrix of the LRC parameter fits (Lasslop et al., 2010; Menzer et al., 2013), which were not available for the BGC online tool.

- Density plot of estimated standard deviation of uncertainty of the annually aggregated GPP across sites due to uncertainty in parameters estimation during day-time based flux partitioning. Higher estimates with are caused by taking into account the uncertainty in temperature sensitivity, E₀.
 - The introduced uncertainty is reduced by smoothing the E_0 estimates across several successive windows (Fig. ?? top) before estimating parameters of the LRC. This smoothing has also an effect on predicted half-hourly gross fluxes (Fig. ?? bottom).
- 10 Effects of smoothing (top) successive estimates of temperature sensitivity, E₀, on predicted GPP (bottom) for site PT-Esp.
 - Results of daytime partitioning are sensitive to subtle details of the procedure. Hence, there is quite much scatter introduced by the differences of processing—with default options or with options that maximize compatibility with the BGC online tool (Fig. ?? top). One such a subtle options is to decrease or to not account for the unreasonably high leverage of some observations during the fit to the light-response curve by some records having a very small estimate of its uncertainty (Fig. ?? bottom). Due to the sensitivities of the day-time partitioning, there are still differences between—with compatibility options and the BGC online tool (Fig. ??).
 - Sensitivity of estimated monthly GPP fluxes to specific processing details results in scatter between GPP predictions based on different—options. Most of the differences between default options and compatibility options (top) are caused by differences in weighting different records during the fit (bottom.)
- 20 Histogram of difference between monthly GPP predictions of with Lasslop10 compatibility options and the BGC online tool.

(Reichstein et al., 2005a)

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