



## Technical note: A view from space on global flux towers by MODIS and Landsat: The FluxnetEO dataset

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**Abstract.** The eddy-covariance technique measures carbon, water, and energy fluxes between the land surface and the atmosphere at several hundreds of sites globally. Collections of standardised and homogenised flux estimates such as the LaThuile, Fluxnet2015, National Ecological Observatory Network (NEON), Integrated Carbon Observation System (ICOS), AsiaFlux, and Terrestrial Ecosystem Research Network (TERN) / OzFlux data sets are invaluable to study land surface processes and vegetation functioning at the ecosystem scale. Space-borne measurements give complementary information on the state of the land surface in the surroundings of the towers. They aid the interpretation of the fluxes and support the training and validation of ecosystem models. However, insufficient quality, frequent and/or long gaps are recurrent problems in applying the remotely sensed data and may considerably affect the scientific conclusions drawn from them. Here, we describe a standardised procedure to extract, quality filter, and gap-fill Earth observation data from the MODIS instruments and the Landsat satellites. The methods consistently process surface reflectance in individual spectral bands, derived vegetation indices and land surface temperature. A geometrical correction estimates the magnitude of land surface temperature as if seen from nadir or 40° off-nadir. We offer to the community pre-processed Earth observation data in a radius of 2 km around 338 flux sites based on the MCD43A4/A2, MxD11A1 MODIS products and Landsat collection 1 Tier1 and Tier2 products. The data sets we provide can widely facilitate the integration of activities in the fields of eddy-covariance, remote sensing and modelling.

### 15 1 Introduction

The installation and maintenance of instrumental infrastructure at eddy-covariance (EC) sites worldwide require considerable financial and logistical efforts and labour force. The precious data sets of land-atmosphere fluxes and environmental conditions allow fundamental insights on ecosystem functioning (Baldocchi, 2008; Baldocchi et al., 2018; Baldocchi, 2020; Migliavacca



et al., 2021; Nelson et al., 2020). A significant achievement is the central processing, quality control, and open standardised  
20 distribution of a large number of the available observational records in data collections such as the LaThuile, Fluxnet2015,  
ABCflux (amongst others, Papale et al., 2006; Baldocchi, 2008; Pastorello et al., 2020; Virkkala et al., 2021b; Papale, 2020) to  
which many site teams contribute.

Complementary information from satellites or cameras (phenocams, Wingate et al., 2015) aid and refine studies of local land-  
atmosphere interactions as they relate to ecosystem structure, phenology, and functioning and the state of the land surface (e.g.,  
25 Migliavacca et al., 2015). Earth observation (EO) data for varying regional sizes around the sites can represent the actual area  
that contributes to the flux measurements - partly even more accurately than similar ground-based measurements can (Gamon,  
2015) - provided sufficiently high spatial resolution and temporal overlap with the site-level records. Next to local studies, the  
combination of flux and satellite observations is also a basic ingredient for upscaling exercises of the in-situ fluxes to larger  
areas or even the globe (Ueyama et al., 2013; Tramontana et al., 2016; Jung et al., 2019, 2020; Joiner et al., 2018; Reitz et al.,  
30 2021; Virkkala et al., 2021a; Zeng et al., 2020).

Independent of the nature of the scientific application, the quality control and gap structure of both the EC and the EO data are  
the groundwork of each analysis. Different criteria help to identify problematic data points with differing levels of strictness  
depending on the given application. Moffat et al. (2007) and Falge et al. (2001) describe techniques to fill gaps due to missing  
data points in the EC data. The literature also offers a diverse set of methods to gap-fill EO data that include spatial, temporal or  
35 cross-sensor approaches (to name a few, Wang et al., 2012; v. Buttler et al., 2014; Weiss et al., 2014; Verger et al., 2011, 2013;  
Kandasamy et al., 2013; Moreno et al., 2014; Moreno-Martínez et al., 2020; Yan and Roy, 2018; Ghafarian Malamiri et al.,  
2018; Li et al., 2018; Dumitrescu et al., 2020). The pre-processing steps are laborious and they are key to the results of the  
analyses. This contribution proposes a set of steps for the systematic quality assurance and gap-filling of key land surface  
indicators from EO data at varying resolutions. We apply them to official data products from the Moderate Resolution Imaging  
40 Spectroradiometer (MODIS) instruments and the sensors on board the Landsat satellites. Both MODIS and Landsat have long  
observational coverage with a high temporal overlap with most freely available EC records. Landsat measurements are of partic-  
ular interest because they resolve small spatial details in pixels of 30 m size, but at the cost of missing out on short temporal  
features. The opposite is true for MODIS data products, which partly average over heterogeneous areas in spatially compara-  
tively coarse pixels of several hundred meters. However, MODIS offers daily, partly even sub-daily temporal resolution. We  
45 process EO data sets of both surface reflectance and land surface temperature (LST) for a limited area around a given flux  
site. For both the quality control and the gap-filling, the approaches aim to be generalisable across all sites without accounting  
for specific local conditions, yet flexible enough to accurately reproduce phenological behaviour and characteristic features  
such as disturbances or fast transitions in managed ecosystems. The procedure shall be as simple as possible, computationally  
efficient and not resort to additional data sources to facilitate a potential application to EO data at global scale.

50 Observation geometries need special attention as the MODIS instruments measure in a wide swath to obtain high temporal  
coverage. They scan across their track from right to left with view zenith angles up to 65 degree from nadir. The wide range of  
viewing geometries leads to different fractions of surface types seen from one overpass to the next for a given site. In addition,  
vegetation structure and topography, together with the position of the sun relative to the sensors, cause variable shadowing



55 effects. The reflectance product (MODIS MCD43A4, Schaaf and Wang (2015b)) partly accounts for these anisotropy effects  
and simulates a nadir view. In order to partly account for variability in the observed LST that is related to changing observation  
geometry (Rasmussen et al., 2011; Guillevic et al., 2013; Ermida et al., 2014), a correction approach developed by Ermida  
et al. (2018) estimates an LST offset as if the instrument would measure from directly above a site. For some applications,  
an oblique view might be favourable over a nadir constellation, for example to enhance the contribution of vegetation canopy  
60 to the LST estimate and minimise fractions of soil or understorey. In addition, we provide LST corrected to a viewing zenith  
angle of 40 degrees. In contrast to MODIS, the Landsat sensors acquire images at much smaller view angles around 7.5-degree  
from nadir. Ground control points and a digital elevation model help to correct for small directional effects related to terrain  
structure and viewing angles (Wulder et al., 2019).

The FluxnetEO products of surface reflectance, vegetation indices, and LST, that result from the proposed processing, are freely  
available by the services of the ICOS Carbon Portal (see data availability statement, (Walther et al., 2021a, b)). Each data set  
65 has a complementary data layer with additional flags to inform the user whether data points correspond to actual good quality  
observations according to the proposed criteria or whether they have been estimated in different gap-filling steps. For all sites,  
the FluxnetEO products cover the period 1984-2017 and 2000-2020 for Landsat and MODIS, respectively. We describe details  
about data inputs in section 2.2, explain the quality control and gap-filling approaches in section 3, and provide information on  
the resulting products in table 2 and the data availability section.

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## 2 Data

### 2.1 Eddy-covariance sites

Here, we select the 338 sites from the LaThuile, Fluxnet2015 (Pastorello et al., 2020) and ICOS Drought 2018 Initiative  
(Drought 2018 Team and ICOS Ecosystem Thematic Centre, 2020) flux data releases. Site coordinates given in different  
75 sources (Ameriflux, Asiaflux, Europe-Fluxdata, Fluxdata.org, and a previously compiled in-house Fluxnet-site location list)  
may differ. In that case, the coordinates with the highest precision were selected. In case the coordinates differed by more than  
0.001° for a given site, a manual check in Google Earth identified the correct or most probable location of the site. The final  
set of 338 sites for which we process the MODIS and Landsat EO data is listed in table A1. Forests and grasslands are best  
represented among the 338 sites. The collection includes fewer sites from savannas and shrublands, and only one site from a  
80 deciduous needleleaf forest (table 1).

### 2.2 MODIS and Landsat

The MCD43A4 product combines AQUA and TERRA observations and provides estimates of surface reflectance in the  
MODIS bands 1-7 (Schaaf and Wang, 2015b). Time series represent observations modelled at nadir view at a resolution of  
16 days and 500 m spatial pixels. For the quality control of MCD43A4, a complementary product, MCD43A2, contains band



**Table 1.** Representation of different plant functional types and Koeppen climate classes across the 338 sites in the FluxnetEO collection.

plant functional type	number of sites	Koeppen main climate	number of sites
evergreen needleleaf forest (ENF)	86	arid	26
evergreen broadleaf forest (EBF)	25	equatorial	23
deciduous needleleaf forest (DNF)	1	warm temperate	171
deciduous broadleaf forest (DBF)	40	snow	103
mixed forest (MF)	13	polar	12
woody savanna (WSA)	10	undefined	3
savanna (SAV)	11		
closed shrubland (CSH)	6		
open shrubland (OSH)	19		
grassland (GRA)	58		
crops (CRO)	36		
wetlands (WET)	32		
snow (SNO)	1		

85 specific information on the quality of the inversion of the bidirectional reflectance distribution function as well as snow cover, platform information and land/water coverage in the scene (Schaaf and Wang, 2015a).

The MODIS MOD11A1 (TERRA, starting in 2000) and MYD11A1 (AQUA, starting in 2002) products (hereafter jointly referred to as MxD11A1, Wan et al. (2015a, b)) provide daily LST and emissivity estimates aligned with quality and view angle information at 1 km spatial pixel sizes. The LST values represent instantaneous values and are selected based on viewing  
90 zenith angle and LST values (MOD11A1 user guide, [https://lpdaac.usgs.gov/documents/118/MOD11\\_User\\_Guide\\_V6.pdf](https://lpdaac.usgs.gov/documents/118/MOD11_User_Guide_V6.pdf)). Four LST data streams are available: TERRA<sub>day</sub> with observations around 10.30 am local time, AQUA<sub>day</sub> with observations around 1.30 pm, TERRA<sub>night</sub> around 10.30 pm and AQUA<sub>night</sub> around 1.30 am. For each of them, observation times vary between overpasses by about  $\pm 1.5$  hours.

95 Reflectance-based Landsat time series comprise the entire multi-temporal collection 1 of the Landsat 4, 5, 7 and 8 archives (<https://landsat.gsfc.nasa.gov/data>) covering the period 1984-2017 at 30 m spatial pixel size. The seven spectral bands of the Landsat product were collected: BLUE, GREEN, RED, near infrared (NIR), shortwave infrared 1 and 2 (SWIR1, SWIR2), and thermal infrared (TIR) (<https://landsat.usgs.gov/what-are-band-designations-landsat-satellites>). Landsat data have been pre-processed using the Landsat Ecosystem Disturbance Adaptive Processing System (LEDAPS, Schmidt et al., 2013) and  
100 the Landsat Surface Reflectance Code (LaSRC, <https://landsat.usgs.gov/landsat-surface-reflectance-data-products>) for atmospheric correction. The pixelQA layer contains information related to clouds, cloud shadows, snow, and ice and is useful for the quality control of the Landsat data (Zhu and Woodcock, 2012; Zhu et al., 2015).



105 The services by Google Earth Engine (Gorelick et al., 2017) provided cutouts of the above mentioned products at the EC sites. Independently of the product and its spatial resolution, the cutout area was limited to a maximum distance of 2 km between a given tower and the centre of a given satellite pixel. Downloading the EO data in tiff-format avoided intransparent reprojection of the data from sinusoidal to regular grid by Google Earth Engine, which would have been problematic for the quality flags in the MCD43A2 and MxD11A1 products. The Landsat data were already provided in regular grid by Google Earth Engine.

### 110 3 Methods

Data processing works separately for each pixel in a cutout (henceforth subpixel). We describe here the overall concept and rationale of the quality filter and the gap-filling, but report all technical details in the Appendix B.

#### 3.1 Processing steps of reflectance-based indicators

##### 3.1.1 Quality control and computation of spectral indices

115 Quality control of the MODIS reflectance-based vegetation indices focused on three aspects: good inversion quality of the bidirectional reflectance distribution function as indicated by the BRDF\_Albedo\_Band\_Quality\_Bandx flags in the MCD43A2 product, snow-free conditions according to the Snow\_BRDF\_Albedo flag, and the omission of reflectance values that are affected by the presence of water in the field of view using the BRDF\_Albedo\_LandWaterType flag. For the selected data samples which passed those criteria we computed a large set of spectral vegetation indices (table 2). An additional check  
120 removed possible values of the vegetation indices outside their defined ranges. Some of the time series contained obvious outlier values. We employed an empirical filter which largely removed those samples which had a particularly large difference to the median of their surrounding values in a temporal window (Papale et al., 2006, technical details on all filters in Appendix B).

In the Landsat data, the flag pixel\_qa provided quality attributes (CFMask, Foga et al., 2017) and removed pixels that  
125 contained snow/ice, cloud, and/or cloud shadow using a binary flag of presence. Similar to the MODIS product, we computed a series of spectral vegetation indices (table 2) using the good quality observations and removed possible values of the indices outside their defined ranges. A slightly modified filter removed possible outlier values also for the Landsat data (see details in Appendix B.)

##### 3.1.2 Gap-filling

130 In the literature several gap-filling and smoothing approaches are available which work in one or more dimensions (e.g., Wang et al., 2012; Kandasamy et al., 2013; v. Buttler et al., 2014; Weiss et al., 2014; Yan and Roy, 2018; Zhang et al., 2021) or use fusion methods between sensors (Verger et al., 2011; Moreno-Martínez et al., 2020). They differ in their levels of sophistication and computational efforts. One of our requirements for the gap-filling approach was that it employs exclusively



temporal operations and does not use additional data sources. This allows the gap-filling to be generally applicable to a single  
135 time series per site, to several subpixels in a cutout around a site and also to global EO data. The idea was to retain the good  
quality data and make as realistic estimates as possible for the gaps between them instead of representing a gap-free time series  
from fitting functions to the valid data. The following recipe describes the steps to estimate missing data points conceptually,  
all technical details we report in Appendix B:

1. Fill short non-snow related gaps ( $\leq 5$  days or  $\leq 1$  month for MODIS and Landsat, respectively) with a median across  
140 valid values in moving windows of 16 days (3 months for Landsat). The moving median only fills gaps, it does not  
change/ smooth valid data points.
2. Fill snow related gaps with a constant baseline value which is identified as the average of valid data points adjacent to  
snow covered periods, i.e. immediately before snow fall or after snow melt (after Beck et al., 2007, but see details in  
Appendix B). Consider all times with a snow flag larger than 0.1 or missing snow information as snow covered. The latter  
145 periods are included as the snow flag appears to systematically miss snow periods in higher latitudes in the beginning of  
the winter. Still, frequent gaps with missing snow information also occur during the growing season. In order to avoid  
wrong filling with a constant value during the growing season this gap-fill step is not applied when the probability of  
snow cover is low, i.e. when the average seasonal cycle indicates typically snow-free conditions at a given time of the  
year, or when typically no snow occurs at all at a given site.
- 150 3. Subsequently, another moving median in windows of 40 days (4 months for Landsat) fills gaps shorter than 65 days (2  
months for Landsat).
4. Compute the median seasonal cycle and use it to fill longer gaps by linearly scaling it to the time series in temporal  
windows. This windowed operation guarantees more flexibility to correctly represent inter-annual variations in the time  
series and might even partly account for changes in the shape of the seasonal cycle due to disturbances. It is, however,  
155 not suited to fill regularly recurring gaps at a certain time of the year, e.g. during rain seasons (Verger et al., 2013).
5. Fill the remaining gaps by piecewise cubic polynomial interpolation. Time series with less than 300 valid data points in  
the whole record after application of all the previous gap-filling steps will not be meaningful for analysis but are still  
filled by nearest neighbour interpolation.
6. Temporal operations cannot meaningfully fill gaps at the beginning and at the end of the record. Therefore the first/last  
160 valid data points are repeatedly appended at the beginning/end of the record.

The described processing steps are generalisable across a range of spectral vegetation indices and can reliably fill missing  
data points across sites globally (see examples in section 4). However, a number of sites have extremely low data availability  
after quality checks, and the gaps in their time series are challenging to temporally interpolate in a meaningful way. This  
can lead to problematic gap-filled data points whose reliability and realism are questionable. Examples are tropical sites and/  
165 or sites with a pronounced wet season with permanent cloud cover. The same generally applies for MODIS in the years



2000-2002 when observations stem mainly from the TERRA satellite and therefore data availability is comparatively low. For Landsat, the number of available scenes is relatively heterogeneous across the globe (<https://www.usgs.gov/media/images/cumulative-number-scenes-landsat-archive>) with some regions having a very good coverage (e.g., North America) while other regions are observed less frequently (e.g., Russia and Africa). Such differences in the availability of good quality data between sites strongly affect the quality of the gap-filling at site level. FluxnetEO therefore provides for each data layer a gap-fill flag which describes whether and if so, how a certain data sample has been imputed which allows users to explore individual sites and use (parts of) the gap-filled data or resort to only using the high quality original data points.

## 3.2 Preprocessing of MODIS land surface temperature

### 3.2.1 Quality checks

The quality control of the MODIS LST did not use the flags provided in the MxD11A1 products, but focussed on the removal of outlier values. Negative outlier values in LST might represent residual cloud contamination, whereas unusually high values might originate from undetected saturation in the level 1 data. Empirical quality checks followed the procedure for the MODIS reflectances, i.e. they discarded data points that deviated strongly from the median of their surrounding values in temporal windows of 30 days (Papale et al., 2006). An additional sanity check eliminated any daytime LST that was lower than the minimum of AQUA and TERRA nighttime LST for a given day.

### 3.2.2 Geometrical correction

For several applications, variable viewing geometries as inherent in the MODIS LST observations are not desirable. A geometrical correction approach developed by Ermida et al. (2018) accounted for directionality in LST retrievals due to vegetation structure and topographical effects. A parametric model estimates the magnitude of LST as if constantly observed from nadir or from an angle of 40 degrees between the sensor and the zenith above a given site. Ermida et al. (2018) derived the coefficients for this geometrical model at a resolution of 0.05 degree. We followed the pragmatic approach of selecting the model coefficients for the correction from the pixel containing a given site, and acknowledge that we did not investigate to what extent the given site conditions represent the overall characteristics of the land surface in the allocated pixel. Further input to the geometrical model were the viewing azimuth angles, solar angles at the overpass time and estimates of daily potential radiation at the top of the atmosphere. The geometrical correction was applied to each subpixel in a cutout separately.

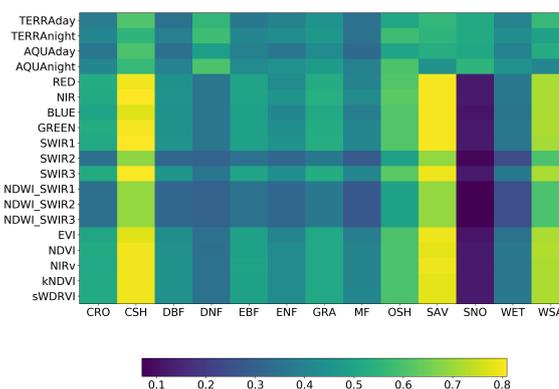
### 3.2.3 Gap-filling

Also for the gap-filling of LST several approaches are present in the literature (e.g., Gerber et al., 2018; Ghafarian Malamiri et al., 2018; Li et al., 2018; Dumitrescu et al., 2020). When using exclusively operations in time and no ancillary data to estimate invalid LST observations, one needs to take care to respect the shorter autocorrelation of LST compared to the reflectance-based indicators. According to Vinnikov et al. (2008) the weather-related component of clear-sky LST has an autocorrelation of about 3 days. The following sequence of steps filled the four MODIS LST data streams (for technical details refer to Appendix C):



1. Similar to the reflectances, a first step consisted in a temporal moving median to fill gaps, but in shorter windows of eight days.
- 200 2. A second step was inspired by Li et al. (2018) and Crosson et al. (2012) and foresaw to use one of the four MODIS LST time series as a 'reference' to fill gaps in a second 'imputed' one. We computed a median seasonal cycle of the difference of the 'reference' and the 'imputed' MODIS LST. This average shift was linearly scaled to the actual shift in temporal windows. The sum of the scaled average shift and the 'reference' LST filled gaps in the 'imputed' LST time series. This procedure iteratively used three of the MODIS LST data streams to fill the fourth, i.e. each one is imputed once by all three others (see details in Appendix C). This gap-fill step was only possible in cases where not all four MODIS LST  
205 observations were invalid during a given day, but extremely advantageous to preserve short synoptic variability in the gap-fill estimates.
3. In fully cloudy days without any valid LST observation, or in case a period has too few valid observations for a meaningful calibration of the linear model in the previous step, the gap-filling followed the same steps like for the reflectance-based spectral indices:  
210 Linearly scale the valid LST observations of each of the four data streams to their own median annual cycle in temporal windows.
4. Interpolate the remaining gaps with cubic polynomials, or nearest neighbour in case of very low data availability (less than 300 valid data points in an entire time series).
- 215 5. Missing values at the beginning and the end of the record cannot be meaningfully filled by temporal methods and are therefore simply repeated.

Steps 3-5 produced very smooth and therefore less realistic LST estimates than steps 1-2. Also, one needs to be aware that any LST estimate in data gaps from this procedure necessarily represents an LST estimate under clear sky conditions, which can be very different from the real LST under overcast skies (Ermida et al., 2019). This needs to be considered for a given application to prevent effects of clear-sky bias in the LST data sets on the results. Like for the vegetation indices, also the LST data layers  
220 have a gap-fill flag in FluxnetEO describing which data points are original and which gap-filling step filled the missing values.



**Figure 1. Fraction of good quality data in the MODIS time series.** Colours represent the median data availability in tower pixels across sites grouped by plant functional type for 2003–2020 (the time period when both TERRA and AQUA satellites are in space).

## 4 Results and Discussion

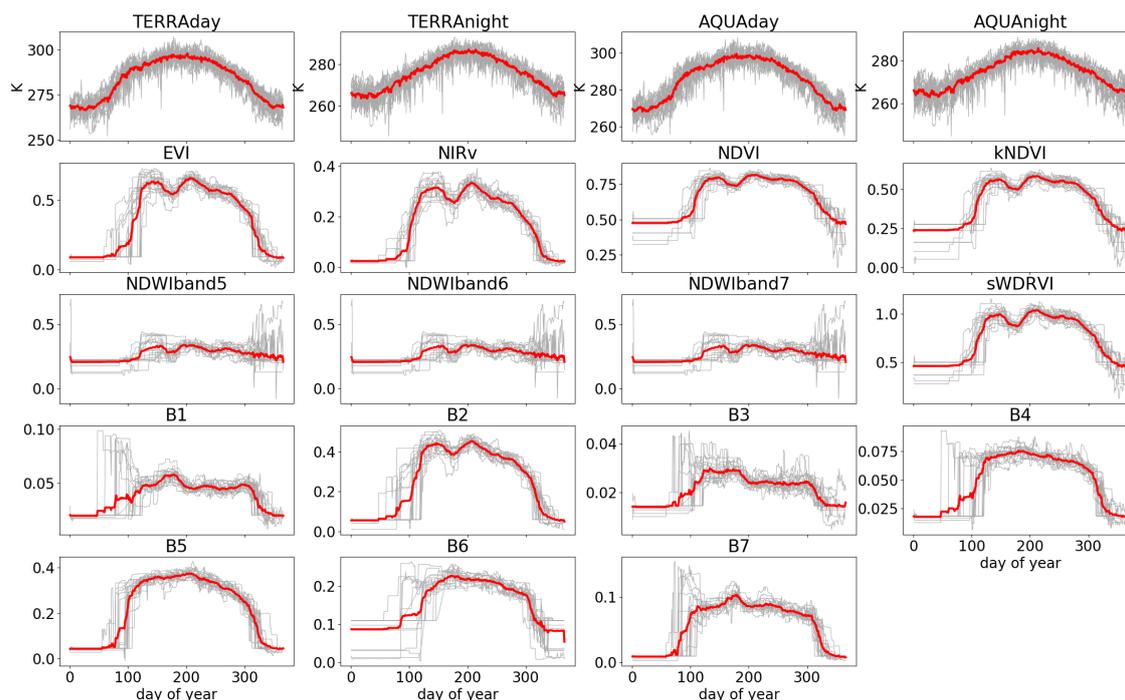
### 4.1 Gap-statistics across indices

Data availability after quality screening is highly variable between sites and depends on the data stream (Fig. 1). MODIS LST generally has less valid data points among the data sets than the reflectance-based indicators, and often less during daytime than nighttime. While the LST are instantaneous values, the reflectances represent averages over 16-day periods. A lower number of good quality observations in indices that rely on band 6 relate to degraded detectors in AQUA MODIS band 6. Large differences in the amount of good quality data between groups of plant functional types, especially for the reflectances, mirror general atmospheric conditions in different regions.

### 4.2 Temporal patterns of the gap-filled time series

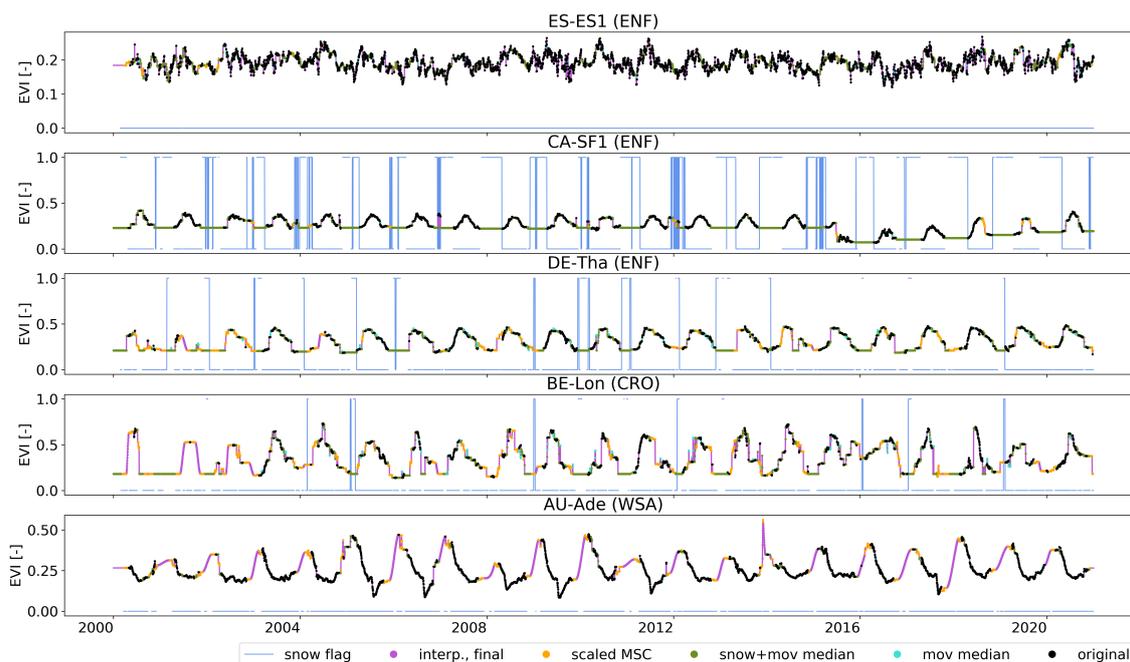
We illustrate some characteristics of the MODIS time series in FluxnetEO using example sites. The Austrian site Neustift (AT-Neu) was situated in a valley in the Alps and surrounded by grasslands which were typically mown three times a year (Wohlfahrt et al., 2008). According to their nature, the LST time series exhibit faster variability than the vegetation indices (Fig. 2). Midday observations (AQUAday) partly show an LST increase after the first harvest event in a year around the day of the year 150. The MSC of most vegetation indices clearly marks the mowing timing, although the relative magnitude varies between indices. Constant values in winter represent snow-covered times.

Focusing on the example of the MODIS EVI, other sites illustrate a few characteristics of the gap-filling procedure in more detail (Fig. 3): At the evergreen needleleaf forest site El Saler in Spain (ES-ES1) much data passes the quality control



**Figure 2.** Median seasonal cycle (red) and individual yearly trajectories (gray) of the LST (top row) and vegetation indices and surface reflectance (second to last rows) in the pixel containing the Austrian site Neustift (AT-Neu).

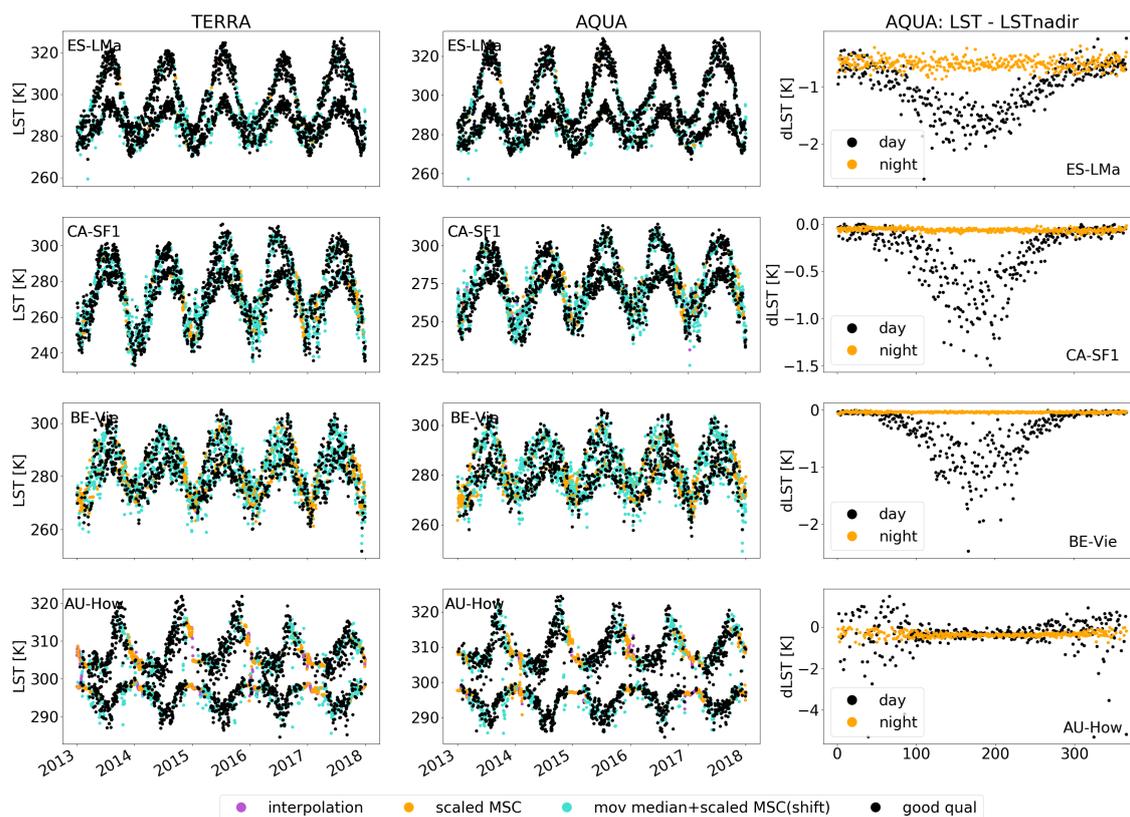
and mostly short gaps are reliably filled also in the absence of a very regular seasonal cycle in EVI. The boreal forest site  
240 Saskatchewan (CA-SF1) illustrates the effect of a disturbance that happened in 2015 (though the site was operated only until  
2006). The gap-filling procedure adapts to the modified conditions both abruptly when the disturbance happens and gradually  
during recovery in the following years. There is a problematic group of high EVI values during winter 2006/07. The moving  
window outlier filter applied to the MODIS reflectances is by design unable to detect those outliers as they occur consecutively  
in a short period of time. In Tharandt (DE-Tha, evergreen needleleaf forest) and Lonze (BE-Lon, crops), data are scarce in  
245 the years 2000–2002 where only TERRA was in operation and the estimated values are less reliable. Also, false filling by  
the snow baseline value during the growing season could not entirely be prevented, causing an unrealistic dip in one year in  
each of the sites. Note that the snow flag contains partly long data gaps in CA-SF1, DE-Tha and BE-Lon. Finally, the woody  
savanna site Adelaide River (AU-Ade) is a typical example of EC sites in climates with a dry and a wet season. While in the  
dry season basically no data gaps occur, cloud coverage in the rainy season is long enough such that mainly the last gap-filling  
250 steps of a linearly scaled MSC and interpolation take effect. Although the scaling of the MSC does not fully succeed in all  
years to produce smooth transitions between the good quality data and the gap-filled ones, the interpolation is able to preserve



**Figure 3.** Illustration of gap-filling steps in the pixel containing selected eddy-covariance sites for the MODIS EVI.

inter-annual variations in the EVI.

Missing LST values were estimated most reliably in the gap-filling steps 1-2 (moving median and scaled average shift to  
255 observations at other overpass times) because the typical short-term variability in the time series could be preserved. In the Spanish site Majadas de Tietar (ES-LMa, Fig. 4 top panel), savanna-type vegetation is prevalent with a dry summer and wet winter. Visually the gap-filling procedure succeeds in preserving the typical higher LST variability in the dry season and seasonally changing diurnal amplitudes. Also, in Saskatchewan (CA-SF1) gap-filling step 2 successfully estimates the largest fraction of missing values for each data stream from the complementary observation times. The EVI indicated a disturbance  
260 event in the beginning of 2015 (Fig. 3) that continued to strongly affect the EVI also in the following year. The event also marks the LST time series in that daytime LST, and therefore, the diurnal amplitude clearly increase in summer after 2015. The gap-filling procedure follows this behaviour. Relative to Majadas de Tietar or Saskatchewan, in the mixed forest in Vielsalm (BE-Vie), data gaps are much more persistent throughout a day and the gap-filling works more often with the third gap-filling step using an average seasonal cycle of LST to estimate missing observations. Finally, at the woody savanna site Howard  
265 Springs in northern Australia (AU-How, Fig. 4 bottom panel) there is a strong seasonal phasing between daytime and night-time LST. Data availability also changes with the seasons. In the monsoon season, synoptic variability in the filled data points



**Figure 4.** LST gap-filling steps in the pixel containing selected eddy-covariance sites for daytime and nighttime LST. The rightmost column shows the average annual cycle of the correction factor between LST from variable viewing angles and LST corrected to nadir view.

is unrealistically low because the gap-filling needs to resort to filling by a median seasonal cycle of LST (obtained from those years in which the monsoon starts late) or by interpolation.

Geometrical corrections to nadir viewing angle are much larger and have a stronger seasonality for daytime LST than for nighttime observations (rightmost panel in Fig. 4, Ermida et al. (2018)). The daytime LST value from a nadir view is consistently estimated to be several Kelvin higher than from an oblique view. The Australian Howard Springs is an exception in that the correction offset to nadir has no consistent sign during the wet season .



**Table 2. Data sets presented in FluxnetEO.** b1, b2, b3, b4, b5, b6, b7 refer to the spectral bands. For each of the variables, we provide both the individual subpixels in a radius of 2 km around a given site as well as a single time series which represents an average over all pixels within 1 km of the site weighted with the inverse distance. Each data set spans the time period 2000–2020 (MODIS) and 1984–2017 (Landsat) and contains a flag describing whether a data point is good quality or whether and how it has been estimated in the gap-filling procedures.

index/ variable	MODIS	Landsat 4,5 and 7	Landsat 8	notes
reflectance-based indicators				
EVI	$2.5 \cdot \frac{b2-b1}{b2+6*b1-7.5*b3+1}$	$2.5 \cdot \frac{b4-b3}{b4+6*b3-7.5*b1+1}$	$2.5 \cdot \frac{b5-b4}{b5+6*b4-7.5*b2+1}$	Huete et al. (2002)
NDVI	$\frac{b2-b1}{b2+b1}$	$\frac{b4-b3}{b4+b3}$	$\frac{b5-b4}{b5+b4}$	Tucker (1979)
kNDVI	$\tanh\left(\frac{b2-b1}{b2+b1} \cdot \frac{b2-b1}{b2+b1}\right)$	$\tanh\left(\frac{b4-b3}{b4+b3} \cdot \frac{b4-b3}{b4+b3}\right)$	$\tanh\left(\frac{b5-b4}{b5+b4} \cdot \frac{b5-b4}{b5+b4}\right)$	Camps-Valls et al. (2021)
NDWI_SWIR1	$\frac{b2-b5}{b2+b5}$	$\frac{b4-b5}{b4+b5}$	$\frac{b5-b6}{b5+b6}$	Gao (1996)
NDWI_SWIR2	$\frac{b2-b6}{b2+b6}$	$\frac{b4-b7}{b4+b7}$	$\frac{b5-b7}{b5+b7}$	Gao (1996)
NDWI_SWIR3	$\frac{b2-b7}{b2+b7}$	NA	NA	Gao (1996)
NIRv	$(NDVI - 0.08) \cdot b2$	$(NDVI - 0.08) \cdot b4$	$(NDVI - 0.08) \cdot b4$	Badgley et al. (2017)
sWDRVI	$\frac{(0.3-1)+(0.3+1) \cdot \frac{b2-b1}{b2+b1}}{(0.3+1)+(0.3-1) \cdot \frac{b2-b1}{b2+b1}} + \frac{1-0.3}{1+0.3}$			Gitelson (2004)
RED	b1	b3 (RED)	b4 (RED)	
NIR	b2	b4 (NIR)	b5 (NIR)	
BLUE	b3	b1 (BLUE)	b2 (BLUE)	
GREEN	b4	b2 (GREEN)	b3 (GREEN)	
SWIR1	b5	b5 (SWIR1)	-	
SWIR2	b6	-	b6 (SWIR1)	
SWIR3	b7	b7 (SWIR2)	b7 (SWIR2)	
MODIS land surface temperature				
LST	TERRA <sub>day</sub> , TERRA <sub>night</sub> , AQUA <sub>day</sub> , AQUA <sub>night</sub>			each with variable viewing zenith angle
LST_nadir	TERRA <sub>day</sub> , TERRA <sub>night</sub> , AQUA <sub>day</sub> , AQUA <sub>night</sub>			corrected to viewing zenith angle = 0 degrees
LST_oblique	TERRA <sub>day</sub> , TERRA <sub>night</sub> , AQUA <sub>day</sub> , AQUA <sub>night</sub>			corrected to viewing zenith angle = 40 degrees



### 4.3 On the importance of spatial context

The type and distribution of the vegetation around a given EC measurement station are not necessarily homogeneous. Instead, clusters of different vegetation or land use types might prevail in different sections of the immediate surroundings of a site. The area that a given flux measurement is representative of (the flux footprint, Schmid, 1997) changes rapidly with wind direction, turbulence conditions, atmospheric stability, and surface resistance (Schmid, 1997; Vesala et al., 2008; Chu et al., 2021). An exact match between the flux footprint and EO data (or a model grid cell) is challenging due to the often unknown or uncertain flux footprints and coarse spatial grid sizes. The scale mismatch is equally important in validation exercises for site-level measurements of surface reflectance (Román et al., 2009; Cescatti et al., 2012), site-level energy-balance closure (Stoy et al., 2013) and model-data integration (Williams et al., 2009). The role that the scale-mismatch between site-level and EO data plays for ecosystem analyses clearly depends on the site and the application. Some applications try to account for the mismatch (Pacheco-Labrador et al., 2017; Wagle et al., 2020), others ignore it and use a custom area around each EC site. Approaches to quantify and account for heterogeneity within a satellite pixel or a certain area around a given site do exist in the literature (Román et al., 2009; Chu et al., 2021; Duveiller et al., 2021), but seem less exploited. In this section, we present different examples for the relevance of spatial context. We computed the average flux footprints for every day (MODIS) and month (Landat) around three example EC stations (Majadas de Tietar, ES-LM1, Gebesee, DE-Geb, and Zotino, RU-Zo2). We illustrate how the relationship between EC-derived gross primary productivity (GPP) and EVI as an EO-derived proxy of the same changes according to whether the footprint area is taken into account or custom cutout sizes are chosen. In RU-Zo2, we compare surface temperature inverted from long-wave outgoing radiation to LST and illustrate how the pixel sizes relate to the flux footprint area (see details on the data processing in Appendix D).

The site ES-LM1 (El-Madany et al., 2018) is a tree-grass ecosystem which is very homogeneous at the remote sensing scale (pixels  $\geq 20$  m). While the trees are evergreen, the herbaceous layer senesces in summer and re-greens in winter (Luo et al., 2018). The EO cutout of  $2 \times 2 \text{ km}^2$  includes irrigated agricultural areas north of the flux footprint. These fields are barren in winter and are covered with crops in summer. MODIS and Landsat EVI are strongly negatively correlated to GPP derived from EC in the pixels over agricultural areas, as are the anomalies of EVI and GPP (Fig. D1 a-d). Conversely, high positive correlations prevail across the remaining larger parts of the EO cutouts. Landsat EVI overlaid by the average flux footprint for two example months illustrates that the EC GPP is only representative of the tree-grass ecosystem (Fig. 5e, g). Hence, the spatial representativeness of EO data for EC fluxes might differ strongly depending on which satellite pixels are chosen for the analysis. We computed the average EVI that is representative of the flux footprint (henceforth fpa for footprint area). We compared it with an average EVI weighted with the probability density function of the flux footprint in order to take into account the decreasing influence of subpixels further away from the tower (henceforth fpw for weighted footprint area), as well as with two pragmatic approaches in case a flux footprint is unknown: an EVI average over all subpixels in the cutout with a radius of 2 km (henceforth fex for full extent) or only the single subpixel that contains the tower (cpx for center pixel). The most noticeable difference between the time series for the different intersection methods is that the full extend (fex) in both

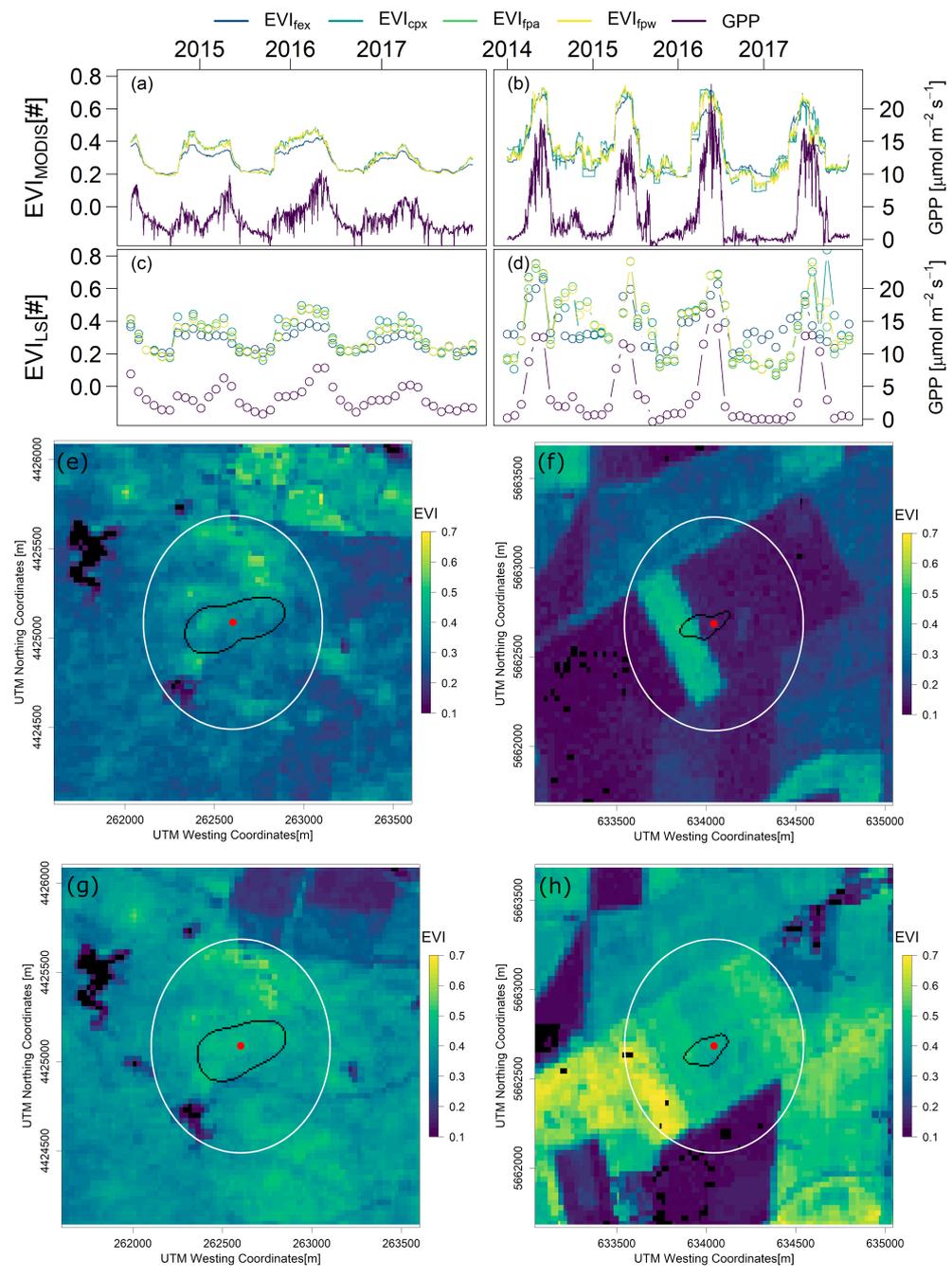


Landsat and MODIS EVI is comparatively lower during the winter period (Fig. 5a,c). The agricultural areas contribute to fex, while the footprint intersection methods (fpa and fpw) and the center pixel (cpx) EVI consistently indicate high greenness in the tree-grass ecosystem.

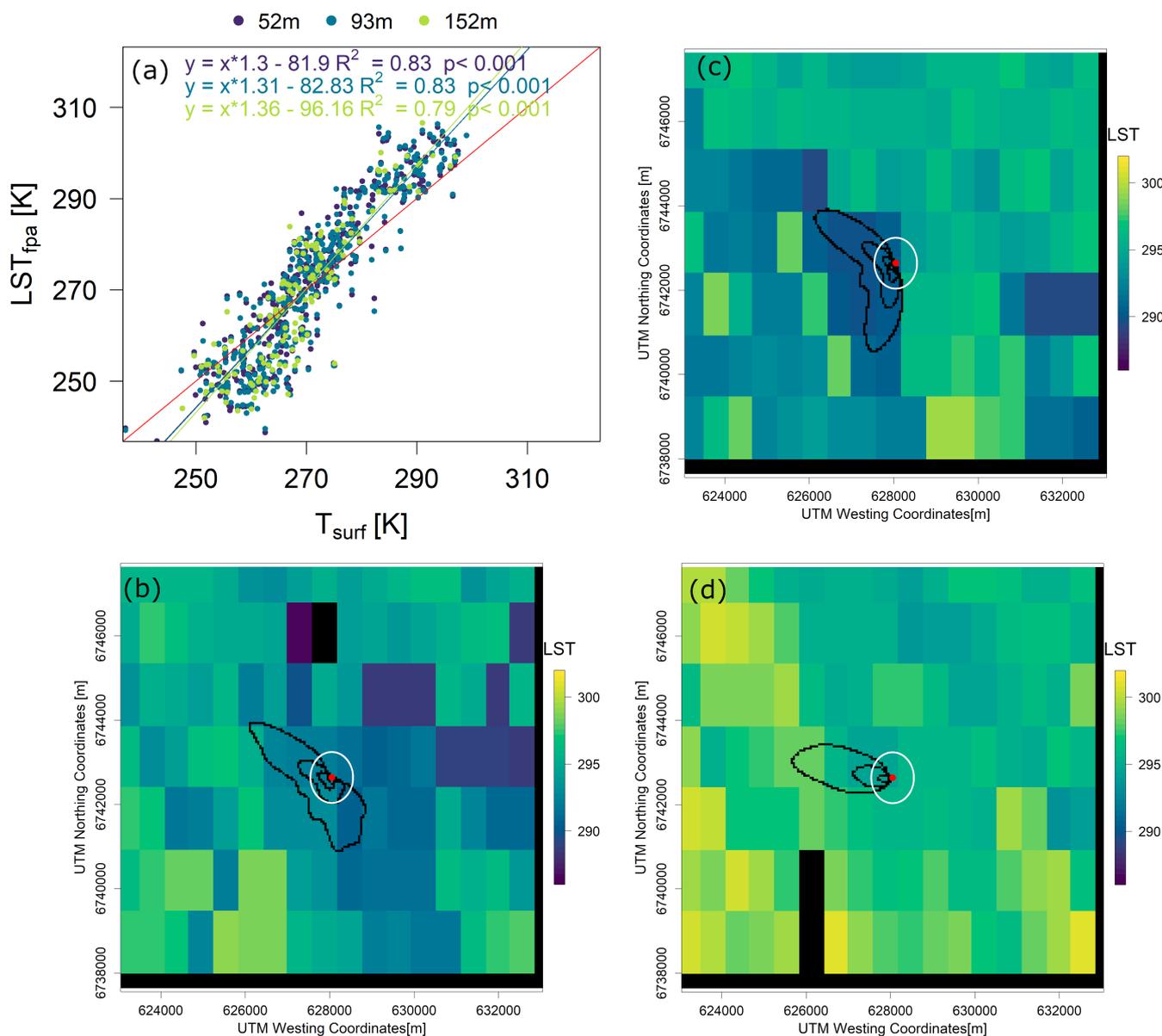
310 Gebesee, DE-Geb, is an agricultural site. The common approach in conducting EC measurements is to put the tower in a location where the land use is as homogeneous as possible, to be able attribute fluxes to a targeted ecosystem, e.g. a known crop type. In Gebesee, this was assured for most of the years in the long site history (e.g. Fig. 5h), but not from 2011-2013. In these years, the field was split into two different adjacent crop types that contributed to the measured fluxes (Fig. 5f), raising the risk for pitfalls in the analyses of the fluxes. Also, in situations/ years when the flux footprint represents a single field, additional  
315 potential difficulties originate from phenological differences between fields within the EO cutouts (Fig. 5f,h) if not properly matched. For example, the anomalies of both GPP and EVI are only highly correlated with each other in the immediate surroundings of the tower (Fig. D1g-h). Phenological heterogeneity between fields might explain why the EVI averaged over the full cutout (fex) is clearly different from the EVI in the footprint area (fpa, fpw) or the tower pixel (cpx) during the growing season maxima in 2015/16 (Fig. 5b,d). Also, consistently with the GPP, the EVI in the tower pixel indicates slightly later  
320 senescence in 2017 than averaged over the footprint area or the full cutout, highlighting considerable effects of a mismatch between the flux footprint and the EO area.

Irrespective of the match between flux footprint and the area that the EVI is representative of, Fig. 5 illustrates the complementarity between MODIS and Landsat in terms of resolution. Although Landsat offers high spatial detail, the temporal patterns that can be resolved with monthly averages are much coarser than the shorter variations that daily MODIS data can describe.  
325 Depending on the application the user of FluxnetEO might choose one or the other.

RU-Zo2, the Zotino tall tower observatory ZOTTO, is located in the taiga-tundra transition zone. The landscape in the proximity of the EC station is a heterogeneous mix of forest, bogs and wetlands. At the tall tower, fluxes are measured at different heights above the canopy. The size of the flux footprint strongly increases with height and the fluxes at the highest level partly  
330 represent areas more than 2 km away from the site (Fig. 6b-d). Flux footprints of measurements closer to the canopy are usually much smaller than the MODIS pixel size of 1 km for the LST, but the flux footprints of the higher measurement levels at RU-Zo2 partly integrate over multiple of such pixels. Size and direction of the footprint extents strongly vary over time (note that Fig. 6b-d represent three consecutive days), such that the vegetation types and surface conditions sampled do not only differ between measurement heights but also between days. We compare spaceborne LST  $AQUA_{day}$  integrated over the flux footprint  
335 area ( $LST_{fpa}$ ) with surface temperature inverted from long-wave outgoing radiation measured at the tower for clear-sky days (Fig. 6a, see details about the methods in Appendix D).  $LST_{fpa}$  of all three heights is consistently about 30% higher than the inverted surface temperature for most of the year, with a notably higher scatter under freezing conditions. The slope between  $LST_{fpa}$  and surface temperature markedly decreases for the highest temperatures, which might indicate significant changes in surface emissivity during the brief peak growing season when vegetation extent is highest and the surface has drained from  
340 snow melt.



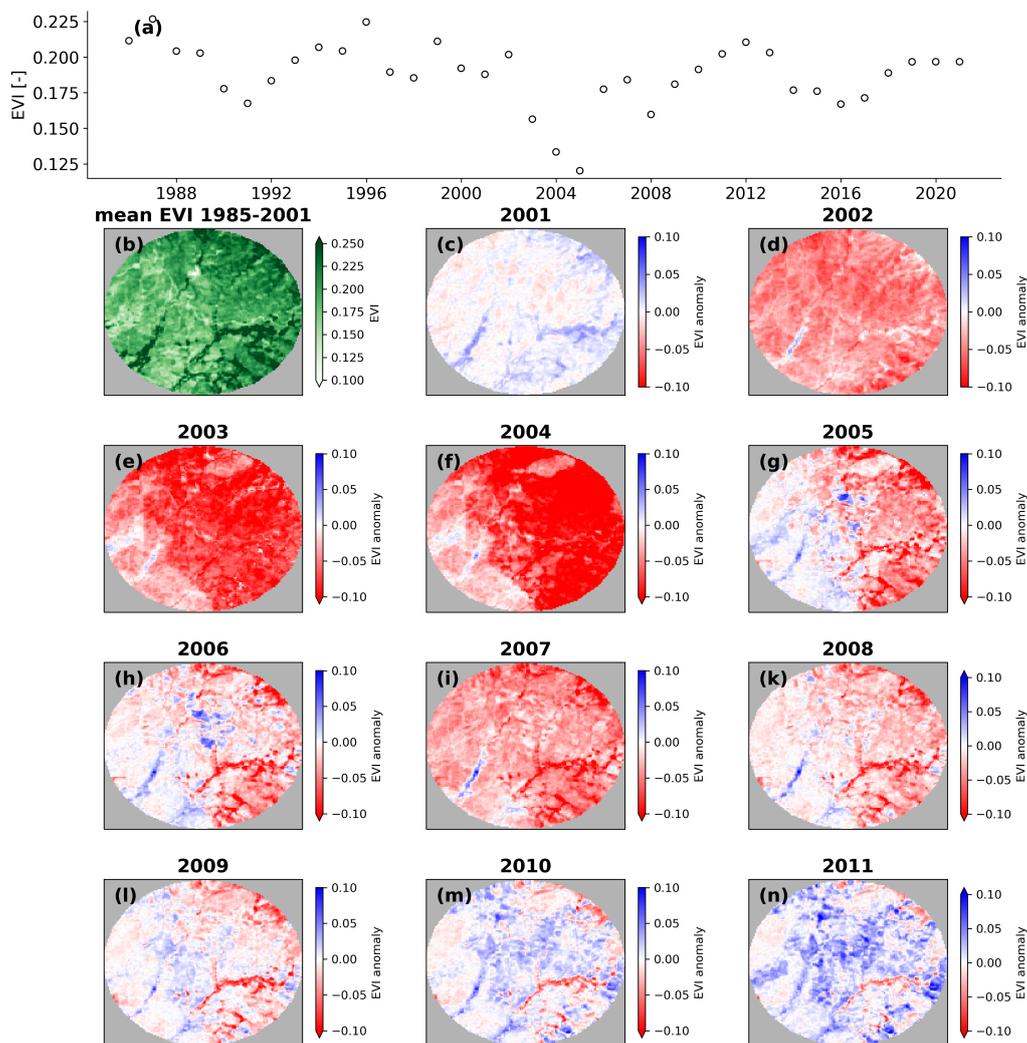
**Figure 5.** Time series of EVI and GPP for ES-LM1 (a,c) and DE-Geb (b,d). MODIS EVI (top row) and Landsat EVI (second row) represent areas with different extents: full extent of the cutout ( $EVI_{fex}$ ), the center pixel that contains a tower ( $EVI_{cpx}$ ), the EVI averaged over the flux footprint area ( $EVI_{fpa}$ ), and the  $EVI_{fpa}$  weighted with the flux probability density function ( $EVI_{fpw}$ ). Subplots e-h: Landsat EVI overlaid with the monthly flux footprint (black line) for ES-LM1 in November 2014 (e) and April 2016 (g), and for DE-Geb in February 2012 (f) and February 2016 (h). Non-original low quality EVI values are blacked out. Red circles indicate the location of the EC station, white circle denotes 1 km diameter from the station.



**Figure 6.** Relationship between MODIS AQUA<sub>day</sub> LST<sub>fpa</sub> and surface temperature calculated from the long-wave outgoing radiation at 302 m above ground (details about the methods in Appendix D). The red line represents the 1:1 line. Subplots b to d show example footprints at the three levels (black lines) overlaid on the LST map from May 31st to June 2nd, 2017, respectively. Non-original low quality LST values are blacked out. The white circle indicates the 1 km diameter around the tower.



Next to matching the flux footprints with the EO data pixels, spatial context is equally important in studies of vegetation recovery after a disturbance event. The Sky Oaks-Young Stand (US-SO3) is a closed shrubland with woody vegetation less than 2 m tall. The US-SO3 site experienced a fire during the period 2002-2003, followed by regrowth. Landsat allows to observe  
345 the impact structure and the spatially very heterogeneous recovery dynamics with remarkable detail (Fig. 7): The fire caused lower than average EVI in large parts of the cutout during the period 2002-2004 (Fig. 7d-f). From 2005 onwards, some patches, particularly the western part of the cutout, appear to have recovered faster from the disturbance than other patches (Fig. 7g). By 2011, EVI has reached pre-fire values in most parts of the area around the site with only small patches as exceptions indicating that regrowth was complete (Fig. 7n). This example illustrates how high spatial resolution EO combined with EC at the site-  
350 level can provide complementary insights for better understanding disturbance regimes and the associated recovery dynamics.



**Figure 7.** Annual EVI dynamics at the site US-SO3 as observed by Landsat. Time series of spatial average annual EVI for the full 2x2km<sup>2</sup> cutout (a) and the long-term temporal average spatial patterns of EVI (b). Annual anomalies of EVI for the period 2003-2011 in panels c-n (anomaly  $EVI_{year\ n} = EVI_{year\ n} - \text{mean}(EVI_{1985-2001})$ ).



## 5 Conclusions

The proposed methods aim at assuring good quality and producing as reliable as possible gap-free estimates of EO-derived surface reflectance, vegetation indices, and LST for pixels around EC sites. Depending on the question/ application at hand, MODIS or Landsat EO data with their inherently very diverse spatial and temporal resolutions might be more suitable. The requirements for the strictness of the quality checks and the sophistication of the gap-filling methods differ by use case. No approach can fit all requirements, but we expect FluxnetEO to offer many opportunities to advance our understanding of land-atmosphere fluxes for individual sites, across regional networks and globally. It helps bridging the Fluxnet, remote sensing, and modelling communities, and facilitates consistent benchmarking of EO-based flux models of any kind. We anticipate that this will accelerate our ability to monitor and understand land-atmosphere fluxes across spatial and temporal scales. For the future we plan to maintain, update and improve FluxnetEO. This will include extending the time series to most recent years, adding EC sites as measurements become available in one of the networks, improving the processing based on newly identified drawbacks and/ or user needs (e.g., Landsat sensors harmonisation), and updating to new EO data collections (e.g. Landsat collection 2). Importantly, forthcoming FluxnetEO versions shall more strongly facilitate complementary usage of multiple missions to exploit their synergy potential, so that future additions will include further EO products, for example the Sentinel missions. Although temporal overlap with most of the EC records is low, it will grow with the lifetime of the different Sentinels and because strong efforts in the EC community target the timely, free and open distribution of site-level measurements.

*Data availability.* Data sets are available for open and free usage under ICOS Carbon Portal in separate collections Landsat (Walther et al., 2021a, [https://meta.icos-cp.eu/collections/-x7\\_Z4PGRuav5QzwgEY\\_DErM](https://meta.icos-cp.eu/collections/-x7_Z4PGRuav5QzwgEY_DErM)) and for MODIS (Walther et al., 2021b, <https://meta.icos-cp.eu/collections/tEakpU6UduMMONrFyym5-tUW>). Zipped folders package the data by continents and groups of countries. In the zip-directories, the files are organised by site and in two processing versions: One version contains spatially explicit data fields for each subpixel in the cutout of 2x2km<sup>2</sup> and is denoted by 'subpixel' in the file name. A second version is an average time series per site that represents the area within 1km radius of the site with the inverse distance to the tower as weight ('average\_cutout'). In this version, at every time step all valid subpixels closer than 1km to the site are averaged after the quality checks, and the gap-filling procedure takes this average time series as input. The data fields contained in both processing versions are listed in table 2. Each data field has a complementary data layer with a flag ('gapfilltype') indicating which data point is of original good quality or how a given point has been imputed in the gap-filling procedure. The processing version 'average\_cutout' has additional fields that indicate how many valid pixels within 1km of the tower contributed to the spatial average per time step ('N') and the spatial standard deviation of the vegetation index or LST for the given time step ('NSTD').

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## Appendix A: Site selection



**Table A1.** Sites in FluxnetEO: site codes and coordinates (latitude in degree N, longitude in degree E, rounded to 4 decimals). Site codes including a \* indicate sites for which currently only MODIS data are provided.

site code	latitude, longitude	site code	latitude, longitude
AR-SLu	-33.4648, -66.4598	AR-Vir	-28.2395, -56.1886
AT-Neu	47.1167, 11.3175	AU-ASM	-22.283, 133.249
AU-Ade	-13.0769, 131.1178	AU-Cpr	-34.0021, 140.5891
AU-Cum	-33.6152, 150.7236	AU-DaP	-14.0633, 131.3181
AU-DaS	-14.1593, 131.3881	AU-Dry	-15.2588, 132.3706
AU-Emr	-23.8587, 148.4746	AU-Fog	-12.5452, 131.3072
AU-Gin	-31.3764, 115.7138	AU-How	-12.4943, 131.1523
AU-RDF	-14.5636, 132.4776	AU-Rob	-17.1175, 145.6301
AU-TTE	-22.287, 133.64	AU-Tum	-35.6566, 148.1517
AU-Wac	-37.4259, 145.1878	AU-Whr	-36.6732, 145.0294
AU-Wom	-37.4222, 144.0944	AU-Ync	-34.9883, 146.2916
BE-Bra	51.3076, 4.5198	BE-Lon	50.5516, 4.7462
BE-Vie	50.3049, 5.9981	BR-Ban	-9.8244, -50.1591
BR-Cax	-1.7197, -51.459	BR-Ji2	-10.0832, -61.9309
BR-Sa1	-2.8567, -54.9589	BR-Sa2	-3.0119, -54.5365
BR-Sa3	-3.018, -54.9714	BR-Sp1	-21.6195, -47.6499
BW-Ma1	-19.9165, 23.5603	CA-Ca1	49.8673, -125.3336
CA-Ca2	49.8705, -125.2909	CA-Ca3	49.5346, -124.9004
CA-Gro	48.2167, -82.1556	CA-Let	49.7093, -112.9402
CA-Man	55.8796, -98.4808	CA-Mer	45.4094, -75.5186
CA-NS1	55.8792, -98.4839	CA-NS2	55.9058, -98.5247
CA-NS3	55.9117, -98.3822	CA-NS4	55.9144, -98.3806
CA-NS5	55.8631, -98.485	CA-NS6	55.9167, -98.9644
CA-NS7	56.6358, -99.9483	CA-Oas	53.6289, -106.1978
CA-Obs	53.9872, -105.1178	CA-Ojp	53.9163, -104.692
CA-Qcu	49.2671, -74.0365	CA-Qfo	49.6925, -74.3421
CA-SF1	54.485, -105.8176	CA-SF2	54.2539, -105.8775
CA-SF3	54.0916, -106.0053	CA-SJ1	53.908, -104.656
CA-SJ2	53.945, -104.649	CA-SJ3	53.8758, -104.6453
CA-TP1	42.6609, -80.5595	CA-TP2	42.7744, -80.4588
CA-TP3	42.7068, -80.3483	CA-TP4	42.7102, -80.3574
CA-TPD	42.6353, -80.5577	CA-WP1	54.9538, -112.467
CA-WP3	54.47, -113.32	CG-Tch	-4.2892, 11.6564
CH-Aws	46.5832, 9.7904	CH-Cha	47.2102, 8.4104
CH-Dav	46.8153, 9.8559	CH-Fru	47.1158, 8.5378



site code	latitude, longitude	site code	latitude, longitude
CH-Lae	47.4781, 8.365	CH-Oe1	47.2858, 7.7319
CH-Oe2	47.2863, 7.7343	CN-Anh	33.0, 117.0
CN-Bed	39.5306, 116.252	CN-Cha	42.4025, 128.0958
CN-Cng	44.5934, 123.5092	CN-Dan	30.4978, 91.0664
CN-Din	23.1733, 112.5361	CN-Do1	31.5167, 121.961
CN-Do2	31.5847, 121.903	CN-Do3	31.5169, 121.972
CN-Du1	42.0456, 116.671	CN-Du2	42.0467, 116.2836
CN-Du3	42.0551, 116.2809	CN-HaM	37.37, 101.18
CN-Hny	29.31, 112.51	CN-Ku1	40.5383, 108.694
CN-Ku2	40.3808, 108.549	CN-Qia	26.734, 115.0663
CN-Sw2	41.7902, 111.8971	CN-Xi1	43.5458, 116.6778
CZ-BK1	49.5021, 18.5369	CZ-BK2*	49.4944, 18.5428
CZ-Lnz	48.6816, 16.9464	CZ-RAJ	49.4437, 16.6965
CZ-Stn	49.036, 17.9699	CZ-wet	49.0246, 14.7704
DE-Akm	53.8662, 13.6834	DE-Bay	50.1419, 11.8669
DE-Geb	51.0997, 10.9146	DE-Gri	50.95, 13.5126
DE-Hai	51.0792, 10.453	DE-Har	47.9344, 7.601
DE-HoH	52.0853, 11.2192	DE-Hte	54.2103, 12.1761
DE-Hzd	50.9638, 13.4898	DE-Kli	50.8931, 13.5224
DE-Lkb	49.0996, 13.3047	DE-Lnf	51.3282, 10.3678
DE-Meh	51.2753, 10.6555	DE-Obe	50.7867, 13.7213
DE-RuR	50.6219, 6.3041	DE-RuS	50.8659, 6.4471
DE-RuW	50.5049, 6.331	DE-Seh	50.8706, 6.4497
DE-SfN	47.8064, 11.3275	DE-Spw	51.8922, 14.0337
DE-Tha	50.9626, 13.5652	DE-Wet	50.4535, 11.4575
DE-Zrk	53.8759, 12.889	DK-Eng	55.6905, 12.1918
DK-Fou	56.4842, 9.5872	DK-Lva	55.6833, 12.0833
DK-Ris	55.5303, 12.0972	DK-Sor	55.4859, 11.6446
ES-Abr	38.7018, -6.7859	ES-Amo	36.8336, -2.2523
ES-ES1	39.346, -0.3188	ES-ES2	39.2756, -0.3153
ES-LJu	36.9266, -2.7521	ES-LM1	39.9427, -5.7787
ES-LM2	39.9346, -5.7759	ES-LMa	39.9415, -5.7734
ES-LgS	37.0979, -2.9658	ES-Ln2	36.9695, -3.4758
ES-VDA	42.1522, 1.4485	FI-Hyy	61.8474, 24.2948
FI-Jok	60.8986, 23.5134	FI-Kaa	69.1406, 27.2698
FI-Let	60.6418, 23.9595	FI-Lom	67.9972, 24.2092



site code	latitude, longitude	site code	latitude, longitude
FI-Sii	61.8326, 24.1928	FI-Sod	67.3624, 26.6386
FI-Var	67.7549, 29.61	FR-Aur	43.5497, 1.1061
FR-Bil	44.4937, -0.9561	FR-EM2	49.8721, 3.0206
FR-Fon	48.4764, 2.7801	FR-Gri	48.8442, 1.9519
FR-Hes	48.6741, 7.0646	FR-LBr	44.7171, -0.7693
FR-Lam	43.4965, 1.2378	FR-Lq1	45.6431, 2.7358
FR-Lq2	45.6392, 2.737	FR-Pue	43.7413, 3.5957
GF-Guy	5.2788, -52.9249	GH-Ank	5.2685, -2.6942
GL-NuF*	64.1308, -51.3861	GL-ZaF	74.4814, -20.5545
GL-ZaH	74.4733, -20.5503	HU-Bug	46.6911, 19.6013
HU-Mat	47.8469, 19.726	ID-Pag	2.345, 114.036
IE-Ca1	52.8588, -6.9181	IE-Dri	51.9867, -8.7518
IL-Yat*	31.345, 35.052	IS-Gun	63.8333, -20.2167
IT-Amp	41.9041, 13.6052	IT-BCi	40.5238, 14.9574
IT-Bon	39.4778, 16.5347	IT-CA1	42.3804, 12.0266
IT-CA2	42.3772, 12.026	IT-CA3	42.38, 12.0222
IT-Col	41.8494, 13.5881	IT-Cp2	41.7043, 12.3573
IT-Cpz	41.7052, 12.3761	IT-Isp	45.8126, 8.6336
IT-LMa	45.1526, 7.5826	IT-La2	45.9542, 11.2853
IT-Lav	45.9562, 11.2813	IT-Lec	43.3036, 11.2698
IT-Lsn	45.7405, 12.7503	IT-MBo	46.0147, 11.0458
IT-Mal	46.114, 11.7033	IT-Noe	40.6062, 8.1512
IT-Non	44.6902, 11.0911	IT-PT1	45.2009, 9.061
IT-Pia	42.5839, 10.0784	IT-Ren	46.5869, 11.4337
IT-Ro1	42.4081, 11.93	IT-Ro2	42.3903, 11.9209
IT-SR2	43.732, 10.291	IT-SRo	43.7279, 10.2844
IT-Tor	45.8444, 7.5781	JP-MBF	44.3842, 142.3186
JP-Mas	36.054, 140.0269	JP-SMF	35.2617, 137.0786
JP-Tak	36.1462, 137.423	JP-Tom	42.7395, 141.5149
MY-PSO	2.973, 102.3062	NL-Ca1	51.971, 4.927
NL-Haa	52.0036, 4.8056	NL-Hor	52.2404, 5.0713
NL-Lan	51.9536, 4.9029	NL-Loo	52.1666, 5.7436
NL-Lut	53.3989, 6.356	PA-SPn	9.3181, -79.6346
PA-SPs	9.3138, -79.6314	PL-Wet	52.7622, 16.3094
PT-Esp	38.6394, -8.6018	PT-Mi1	38.5406, -8.0001
PT-Mi2	38.4765, -8.0246	RU-Che	68.613, 161.3414
RU-Cok	70.8291, 147.4943	RU-Fy2	56.4476, 32.9019



site code	latitude, longitude	site code	latitude, longitude
RU-Fyo	56.4615, 32.9221	RU-Ha1	54.7252, 90.0022
RU-Ha3	54.7046, 89.0778	RU-Sam	72.3738, 126.4958
RU-SkP	62.255, 129.168	RU-Tks	71.5943, 128.8878
RU-Vrk	67.0547, 62.9405	RU-Zot	60.8008, 89.3508
SD-Dem	13.2829, 30.4783	SE-Abi	68.3624, 18.7948
SE-Deg	64.182, 19.5565	SE-Htm	56.0976, 13.419
SE-Lnn*	58.3406, 13.1018	SE-Nor	60.0865, 17.4795
SE-Ros*	64.1725, 19.738	SE-Sk2	60.1297, 17.8401
SE-St1	68.3541, 19.0503	SE-Svb*	64.2561, 19.7745
SJ-Adv	78.186, 15.923	SJ-Blv	78.9216, 11.8311
SK-Tat	49.1208, 20.1635	SN-Dhr	15.4028, -15.4322
UK-ESa	55.9069, -2.8586	UK-Gri	56.6072, -3.7981
UK-Ham	51.1535, -0.8583	UK-PL3	51.45, -1.2667
UK-Tad	51.2071, -2.8286	US-AR1	36.4267, -99.42
US-AR2	36.6358, -99.5975	US-ARM	36.6058, -97.4888
US-ARb	35.5497, -98.0402	US-ARc	35.5465, -98.04
US-Atq	70.4696, -157.4089	US-Aud	31.5907, -110.5104
US-Bar	44.0646, -71.2881	US-Bkg	44.3453, -96.8362
US-Blo	38.8953, -120.6328	US-Bn2	63.9198, -145.3782
US-Bn3	63.9227, -145.7442	US-Bo1	40.0062, -88.2904
US-Bo2	40.009, -88.29	US-Brw	71.3225, -156.6092
US-CRT	41.6285, -83.3471	US-CaV	39.0633, -79.4208
US-Cop	38.09, -109.39	US-Dk3	35.9782, -79.0942
US-FPe	48.3077, -105.1019	US-FR2	29.9495, -97.9962
US-Fmf	35.1426, -111.7273	US-Fuf	35.089, -111.762
US-Fwf	35.4454, -111.7718	US-GBT	41.3658, -106.2397
US-GLE	41.3665, -106.2399	US-Goo	34.2547, -89.8735
US-Ha1	42.5378, -72.1715	US-Ho1	45.2041, -68.7402
US-Ho2	45.2091, -68.747	US-IB1	41.8593, -88.2227
US-IB2	41.8406, -88.241	US-Ivo	68.4865, -155.7503
US-KS1	28.4583, -80.6709	US-KS2	28.6086, -80.6715
US-LWW	34.9604, -97.9789	US-Lin	36.3566, -119.8423
US-Los	46.0827, -89.9792	US-MMS	39.3232, -86.4131
US-MOz	38.7441, -92.2	US-Me1	44.5794, -121.5
US-Me2	44.4523, -121.5574	US-Me3	44.3154, -121.6078
US-Me4	44.4992, -121.6224	US-Me5	44.4372, -121.5668



site code	latitude, longitude	site code	latitude, longitude
US-Me6	44.3233, -121.6078	US-Myb	38.0498, -121.7651
US-NC1	35.8118, -76.7119	US-NR1	40.0329, -105.5464
US-Ne1	41.1651, -96.4766	US-Ne2	41.1649, -96.4701
US-Ne3	41.1797, -96.4397	US-ORv	40.0201, -83.0183
US-Oho	41.5545, -83.8438	US-PFa	45.9459, -90.2723
US-Prr	65.1237, -147.4876	US-SO2	33.3738, -116.6228
US-SO3	33.3771, -116.6226	US-SO4	33.3845, -116.6406
US-SP1	29.7381, -82.2188	US-SP2	29.7648, -82.2448
US-SP3	29.7548, -82.1633	US-SRC	31.9083, -110.8395
US-SRG	31.7894, -110.8277	US-SRM	31.8214, -110.8661
US-Sta	41.3966, -106.8024	US-Syv	46.242, -89.3477
US-Ton	38.4316, -120.966	US-Tw1	38.1074, -121.6469
US-Tw2	38.1047, -121.6433	US-Tw3	38.1159, -121.6467
US-Tw4	38.103, -121.6414	US-Twt	38.1087, -121.653
US-UMB	45.5598, -84.7138	US-UMd	45.5625, -84.6975
US-Var	38.4133, -120.9507	US-WBW	35.9588, -84.2874
US-WCr	45.8059, -90.0799	US-WPT	41.4646, -82.9962
US-Whs	31.7438, -110.0522	US-Wi0	46.6188, -91.0814
US-Wi1	46.7305, -91.2329	US-Wi2	46.6869, -91.1528
US-Wi3	46.6347, -91.0987	US-Wi4	46.7393, -91.1663
US-Wi5	46.6531, -91.0858	US-Wi6	46.6249, -91.2982
US-Wi7	46.6491, -91.0693	US-Wi8	46.7223, -91.2524
US-Wi9	46.6188, -91.0814	US-Wkg	31.7365, -109.9419
US-Wrc	45.8205, -121.9519	VU-Coc	-15.4427, 167.192
ZA-Kru	-25.0197, 31.4969	ZM-Mon	-15.4378, 23.2528



## 380 Appendix B: Technical details about the processing of surface reflectance

In this section we provide all specific technical details necessary to reproduce our processing steps for the surface reflectance of MODIS and Landsat.

The quality control of the MODIS reflectance-based land surface indicators included the following steps:

- 385 – Omission of the MCD43A2 BRDF\_Albedo\_Band\_Quality\_BandX flags  $\geq 3$  for each band to remove bad inversion quality from the surface reflectances.
- The flag Snow\_BRDF\_Albedo eliminated pixels that contain snow. As the gap-filling procedure used the snow information, a spatially aggregated snow flag was needed for the processing version that averages valid data within 1 km of the tower. For this, we defined the aggregated snow flag as the fraction of subpixels in the cutout that are snow covered. If  
390 more than 50% of subpixels have missing snow information for a certain day, the aggregated snow flag is set to missing as well.
- The presence of water in a scene seen by an optical sensor can strongly affect the observation. The BRDF\_Albedo\_LandWaterType flag allowed to filter for pixels exclusively on land (flag=1). This eliminated all data for many Swiss, Dutch, Italian and Finnish sites which are situated close to water bodies. Inclusion of ocean coastlines and lake shorelines (flag=2) and  
395 shallow inland water (flag=3) resulted in reasonable time series at most sites. This came at the cost of having few other sites that were affected by the presence of water. As a trade-off between data availability and quality, we decided to include land-water flags 1-3.
- After the computation of the vegetation indices from the individual spectral bands, an additional check removed possible values of the spectral vegetation indices outside their defined ranges. An outlier filter compared each value to the median  
400 of all valid values in temporal windows of 30 days (Papale et al., 2006). A large difference of a given value to the median of its surrounding values indicates a potential outlier. The threshold  $z$  as in Papale et al. (2006) was set to 2, and only a less conservative threshold of  $z=3$  acted when more than 20 valid values were available in a given window.

The empirical outlier filter for Landsat slightly differed from the one for MODIS and removed observations in the five highest and lowest percentiles of the mean seasonal cycle of an index if they differed more than 75% from their surrounding  
405 3-months moving window median. The second criterion was critical in order to preserve observations of disturbance events or recovery dynamics.

Technical details for the gap-filling:

1. The first step is a moving window median to fill short non-snow related gaps. If the entire time series has less than 40%  
410 valid data, a given moving window contains both the actual values and the median seasonal cycle for the given time of the year. The median for the moving window refers then to the distribution of both.



2. The second step fills reflectance values with a constant value in the presence of snow (snow flag  $\geq 0.1$ ). Partly long periods with missing snow information in the Snow\_BRDF\_Albedo flag needed special treatment. Some of these gaps appeared systematically in early winter in higher latitudes, so also times of missing snow information are considered as snow covered. However, also during the growing season long periods of missing snow information occur in several sites globally. The following criteria check whether a period that is considered snow covered by high values or missing snow flags is filled with a constant baseline value or not:

- If a given site has less than 60 days with valid snow coverage (i.e. Snow\_BRDF\_Albedo=1) in the total record, snow typically does not occur at the site. In this case the gap-filling procedure does not apply this gap-filling step at all for this site.
- The gap-filling with a constant value only addresses gaps with a minimum length of 20 consecutive days with snow flag missing or 1. This avoids filling very short intermittent snow periods or short gaps in snow information during the growing season.
- This gap-filling step does not consider gaps due to missing snow information if the median seasonal cycle of snow coverage indicates  $\leq 5\%$  of snow cover at the given time of the year and the difference between the fill value and the median seasonal cycle is large (i.e. exceeds the 85<sup>th</sup> percentile of the differences in times of missing snow information).

The constant baseline value that is used to fill snow periods in the time series for a site represents the 3<sup>rd</sup> percentile of the median seasonal cycle of the spectral vegetation indices. If a given index typically has high values outside the growing season, the baseline value represents the 97<sup>th</sup> percentile instead. However, if for a given winter the average over the last 5 valid data points at the end of the growing season or over the first 5 valid data points at the beginning of the next growing season is lower than the baseline value (higher than the baseline for indices which are typically high outside the growing season), the baseline takes the value of this average for the given winter (similar to Beck et al., 2007).

3. Linearly scale the median seasonal cycle to the time series to fill longer gaps (Verger et al., 2013). Calibration happens in moving temporal windows of 80 days, and application of the scaling in steps of 20 days.

### Appendix C: Technical details about the processing of MODIS LST

In this section we provide all specific technical details necessary to reproduce the processing steps for the MODIS LST. The empirical filter to remove potential outlier values (Papale et al., 2006) followed the same procedure like for the vegetation indices, but used a constant z-value of 1.5 as it provided the best trade-off between filter success, wrong positives and wrong negatives.

Estimates of LST in data gaps originate from the following steps:



– In contrast to the procedure for the reflectance-based vegetation indices, the distribution of values in the temporal windows of 8 days is not supplied by the median seasonal cycle in case of low data availability. The moving window median was not applied for windows with less than three valid values.

445 – Filling by linearly scaling the median seasonal shift between any two of the four MODIS LST time series to each other (Crosson et al., 2012; Li et al., 2018). The following explains this gap-filling step for  $TERRA_{day}$  as the 'imputed' time series:

1. Obtain the median seasonal cycle (MSC) of the shift between  $TERRA_{day}$  and  $AQUA_{day}$ :

$$MSC( \Delta(TERRA_{day}, AQUA_{day}) ).$$

450 2. Linearly scale  $MSC( \Delta(TERRA_{day}, AQUA_{day}) )$  to  $\Delta(TERRA_{day}, AQUA_{day})$  in temporal windows of 80 days (provided a minimum of 10 valid values in a given window). Apply the scaling in windows and steps of 20 days.

$$\Delta(TERRA_{day}, AQUA_{day})_{t=k:k+80} = f ( MSC( \Delta(TERRA_{day}, AQUA_{day}) )_{t=k:k+80} )$$

$$\Delta(TERRA_{day}, AQUA_{day})_{t=k:k+20}^* = m \cdot MSC( \Delta(TERRA_{day}, AQUA_{day}) )_{t=k:k+20} + n.$$

3. Add the scaled average shift to the  $AQUA_{day}$  to obtain an estimate of  $TERRA_{day}^*[AQUA_{day}]$ .

455 
$$TERRA_{day}^*_{t=k:k+20}[AQUA_{day}] = AQUA_{day=t=k:k+20} + \Delta(TERRA_{day}, AQUA_{day})_{t=k:k+20}^*$$

Analogously to  $TERRA_{day}^*[AQUA_{day}]$ , also the night-time LST observations contributed to estimate  $TERRA_{day}^*[TERRA_{night}]$  and  $TERRA_{day}^*[AQUA_{night}]$ . All three estimates  $TERRA_{day}^*[AQUA_{day}]$ ,  $TERRA_{day}^*[TERRA_{night}]$  and  $TERRA_{day}^*[AQUA_{night}]$ , served to fill gaps in  $TERRA_{day}$ , namely in the order of increasing standard deviation of the differences between valid  $TERRA_{day}$  and each of the three estimated  $TERRA_{day}^*$ .

460 The procedure analogously filled  $AQUA_{day}$ ,  $TERRA_{night}$  and  $AQUA_{night}$  accordingly using valid observations of the remaining three, respectively.

– Linearly scale the MSC of one LST time series to the actual time series in temporal windows. As in step 2, the calibration happened in temporal windows of 80 days, while the scaling was applied in windows of 20 days. Exemplarily for

$$TERRA_{day}: TERRA_{day=t=k:k+80} = f ( MSC( TERRA_{day} )_{t=k:k+80} )$$

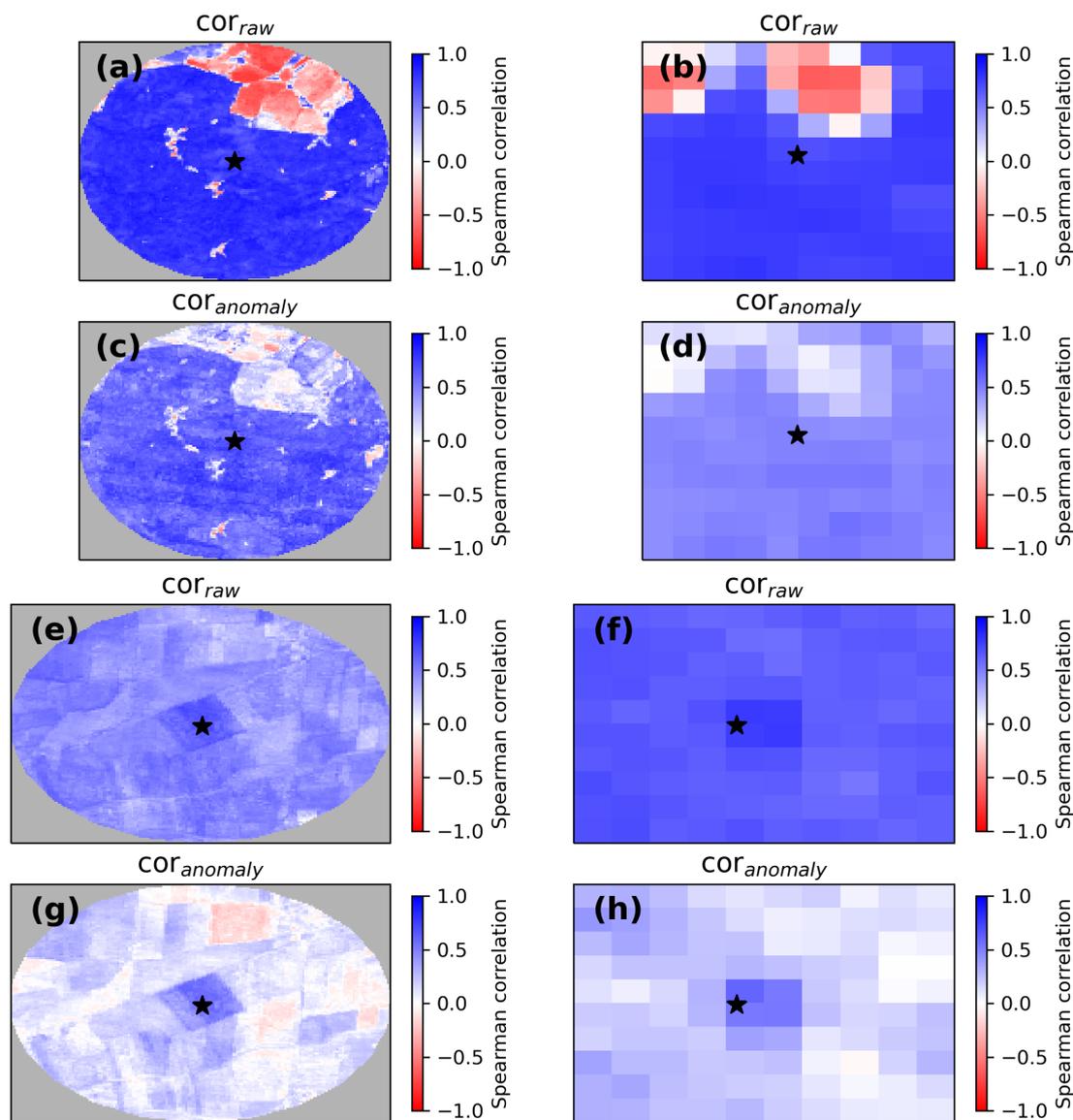
465 
$$TERRA_{day}^*_{t=k:k+20} = m \cdot MSC( TERRA_{day} )_{t=k:k+20} + n$$



#### Appendix D: Details about the analysis of spatial context

For the analysis at DE-Geb and ES-LM1 we used night-time partitioned GPP (Reichstein et al., 2005) with the mean of the variable  $u^*$ -threshold (GPP\_NT\_VUT\_MEAN) from the Drought 2018 Team and ICOS Ecosystem Thematic Centre (2020) data release (Migliavacca et al., 2020; ICOS Ecosystem Thematic Centre and Gebese, 2019). We computed the actual flux  
470 footprints after Kljun et al. (2015) from ICOS drought 2018 data (Drought 2018 Team and ICOS Ecosystem Thematic Centre, 2020) using the R-code version (V1.41) of the FFP-tool. As a flux footprint for the intersection with EVI we define the area that contributes 80% to the flux footprint probability density function (80% isoline of the monthly/daily cumulative flux footprint for Landsat and MODIS, respectively).

Flux footprint calculation followed the same procedure for the three measurement heights at RU-Zo2. Surface temperature was  
475 inverted from long-wave outgoing radiation measured at a fixed height of 302 m using Stefan-Boltzmann law. As the inverted surface temperature was compared to LST AQUA<sub>day</sub>, the average of half-hourly outgoing long-wave radiation for the nominal overpass time at 1.30pm  $\pm$  1.5 hours was taken. Surface emissivity is unknown and we assumed emissivity=1 throughout the year. Only days with good quality in both the LST and the long-wave outgoing radiation are used according to the following  
480 criteria: i) more than 90% of the EO cutout have valid (i.e. non-gapfilled) values which restricts the comparison to clear-sky conditions, and ii) at least 50% of the half-hourly long-wave fluxes in a given day are of good quality. A larger cutout of 5x5 km<sup>2</sup> was extracted for MODIS LST to fully cover also the extent of the flux footprint of the highest measurement level, but is used only for illustrative purposes and not in the data provided in the FluxnetEO collections.



**Figure D1.** Spearman correlation between EVI and GPP using monthly Landsat (a, c, e, g) and daily MODIS (b, d, f, h) data for ES-LM1 (a-d) and DE-Geb (e-h) Fluxnet sites. The correlation estimates were computed on the raw time series (a, b, e, f) and on the anomalies (c, d, g, h).



*Author contributions.* JN and UW compiled the site coordinates and established the pipeline to obtain EO data from Google Earth Engine, and unified formats. SW developed the processing steps for MODIS data with the input from MJ, MM and JN. SB adapted the processing to  
485 Landsat data. SE provided model coefficients, code and guidance on its usage for the LST geometrical correction. SW and UW created the files that are offered to the community. TE computed flux footprints for the example sites and analysed them with respect to the satellite data together with SW and SB. SW wrote the manuscript with contributions from all authors, especially MM, SB, TE.

*Competing interests.* The authors declare no competing interests.

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