A Coupled Ground Heat Flux-Surface Energy Balance Model of Evaporation Using Thermal Remote Sensing Observations

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29 Abstract

30 One of the major undetermined problems in evaporation (ET) retrieval using thermal infrared 31 remote sensing is the lack of a physically based ground heat flux (G) model and its integration 32 within the surface energy balance (SEB) equation. Here, we present a novel approach based on 33 coupling a thermal inertia (TI)-based mechanistic G model with an analytical surface energy 34 balance model, Surface Temperature Initiated Closure (STIC, version STIC1.2). The coupled 35 model is named STIC-TI. The model is driven by noon-night (1:30 pm and am) land surface 36 temperature, surface albedo, and vegetation index from MODIS Aqua in conjunction with a clear-37 sky net radiation sub-model and ancillary meteorological information. SEB flux estimates from 38 STIC-TI were evaluated with respect to the *in-situ* fluxes from eddy covariance measurements in 39 diverse ecosystems of contrasting aridity in both northern and southern hemispheres. Sensitivity 40 analysis revealed substantial sensitivity of STIC-TI-derived fluxes due to the land surface 41 temperature uncertainty. An evaluation of noontime G (G_i) estimates showed 12-21% error across 42 six flux tower sites and a comparison between STIC-TI versus empirical G models also revealed 43 the substantially better performance of the former. While the instantaneous noontime net radiation 44 (R_{Ni}) and latent heat flux (LE_i) were overestimated (15% and 25%), sensible heat flux (H_i) was 45 underestimated (22%). Overestimation (underestimation) of LE_i (H_i) was associated with the 46 overestimation of net available energy $(R_{Ni} - G_i)$ and use of unclosed SEB flux measurements in 47 LE_i (H_i) validation. The mean percent deviations in G_i and H_i estimates were found to be strongly 48 correlated with satellite day-night view angle difference in parabolic and linear pattern, and a 49 relatively weak correlation was found between day-night view angle difference versus LE_i 50 deviation. Findings from this parameter-sparse coupled G-ET model can make a valuable 51 contribution to mapping and monitoring the spatiotemporal variability of ecosystem water stress 52 and evaporation using noon-night thermal infrared observations from future Earth Observation 53 satellite missions such as TRISHNA, LSTM, and SBG.

54 <u>Keywords</u>: Thermal remote sensing, water stress, evaporation, ground heat flux, thermal inertia,
 55 surface energy balance, STIC, terrestrial ecosystem

56 **1 Introduction**

One of the outstanding challenges in evaporation (ET) estimation through surface energy balance 57 58 (SEB) model concerns an accurate characterization of ground heat flux in the open canopy 59 architecture with mixed vegetation such as savanna or in ecosystems with low mean fractional 60 vegetation cover, prevailing water stress, and strong seasonality in soil moisture. Ground heat flux 61 (G) is an intrinsic component of SEB (Sauer and Horton, 2005), affecting the net available energy 62 for ET (the equivalent water depth of latent heat flux, LE) and sensible heat flux (H). It represents 63 an energy flow path that couples surface with the atmosphere and has important implications for 64 the underlying thermal regime (Sauer and Horton, 2005). Depending on the vegetation fraction 65 and water stress, the magnitude of instantaneous G varies greatly across different ecosystems. In 66 the humid ecosystems with predominantly dense canopies and high mean fractional vegetation 67 cover, G contributes to a small proportion in the surface energy balance. Dense canopy cover leads to less transmission of radiative fluxes through multiple layers of canopies, which results in low 68 69 warming of the soil floor. Due to persistently high soil water content, humid ecosystems generally 70 show low diurnal and seasonal variability in G. In contrast, the magnitude of G is substantially 71 large in arid and semi-arid ecosystems with sparse and open canopy and soil moisture deficits. Due 72 to prevailing feedback between the physics of ground heat flux, land-atmosphere interactions and 73 vegetation ecophysiology, evaporation modelling in the complex ecosystems remains a 74 challenging task (Wang et al., 2013; Kiptala et al., 2013). This paper addresses the challenge of 75 simultaneous estimation of G and ET by combining thermal remote sensing observations with a 76 mechanistic G model and a SEB model.

SEB models mainly emphasize on estimating sensible heat flux (H) by resolving the aerodynamic
 conductance (g_A) and computes LE as a residual SEB component as follows:

$$LE = R_N - G - H \tag{1}$$

 R_N is the net radiation. The proportion of R_N that is partitioned into G depends upon soil properties like its albedo, soil moisture, soil thermal properties such as thermal conductivity and heat capacity, which vary with mineral, organic and soil water fractions. SEB models use land surface temperature (LST or T_S) as an important lower boundary condition for estimating H and LE. Due to the extraordinarily high sensitivity of T_S to evaporative cooling and soil water content variations, thermal infrared (TIR) remote sensing is extensively used in large scale evaporation diagnostics (Kustas and Anderson, 2009; Mallick et al., 2014, 2015a, 2018a; Cammalleri and Vogt, 2015; Anderson et al., 2012). Evaporation estimation through SEB models commonly employ empirical sub-models of G in a stand-alone mode. Despite the utility of mechanistic G models is demonstrated in different studies (Verhoef, 2004; Murray and Verhoef, 2007; Verhoef et al., 2012), no TIR-based evaporation study attempted to couple a mechanistic G model with a SEB model.

91 The SEB models for ET estimation driven by remote sensing observations generally use linear and 92 non-linear relationships for estimating G and such methods generally employ R_N , T_S , albedo (α_R), 93 and NDVI (e.g., Bastiaanssen et al., 1998; Friedl, 2002; Santanello and Friedl, 2003). While the 94 inclusion of T_s and albedo serves as a proxy for soil moisture and surface characteristics effects in 95 G, inclusion of NDVI provides a scaling of G - R_N ratio for different fractional vegetation cover. 96 Unfortunately, the empirical approaches do not include any information of soil temperature or 97 daily temperature amplitude. These empirical models also lack the universal consensus. Setting G 98 as a fraction of R_N does not solve the energy balance equation and disregards the role of thermal 99 inertia of the land surface (Mallick et al., 2015b). This could introduce substantial uncertainty in 100 LE estimation because G effectively couples the surface energy balance with energy transfer 101 processes in the soil thermal regime. It provides physical feedback to LE through the effects of 102 soil moisture, temperature, and conductivity (thermal and hydraulic) (Sauer and Horton, 2005). 103 Such feedbacks are most critical in the arid and semi-arid ecosystems where LE is significantly 104 constrained by the soil moisture dry-down. The limits imposed on LE by the water stress 105 consequently result in greater partitioning of the net available energy (i.e., $R_N - G$) into H and G 106 (Castelli et al., 1999).

When LE is reduced due to soil moisture dry-down, both G and T_s tend to show rapid intraseasonal rise. Therefore, the surface energy balance equation could be linked with mechanistic G model, T_s harmonics (Verhoef, 2004), and soil moisture availability. Realizing the importance of direct estimates of G in LE and invigorated by the advent of TIR remote sensing, Verhoef et al. (2012) demonstrated the potential of a TI-based mechanistic model (Murray and Verhoef, 2007) (MV2007 hereafter) for spatiotemporal G estimates in semi-arid ecosystems of Africa. Some studies also emphasized the importance of using noontime and nighttime Ts and R_N for estimating G (Mallick et al., 2015b; Bennet et al., 2008; Tsuang, 2005). The method of MV2007 has so far been tested in a stand-alone mode, and no remote sensing method has so far been attempted to combine such a mechanistic G model (e.g., MV2007-TI model) with a SEB model for coupled energy-water flux estimation and validation.

118 By integrating T_s into a combined structure of the Penman-Monteith (PM) and Shuttleworth-119 Wallace (SW) model, an analytical SEB model was proposed by Mallick et al. (2014, 2015a, 120 2016). The model, Surface Temperature Initiated Closure (STIC), is based on finding analytical 121 solution for aerodynamic and canopy-surface conductance $(g_A \text{ and } g_S)$ where the expressions of 122 the conductances were constrained by an aggregated water stress factor. Through physically 123 linking water stress (T_s derived) with g_A and g_s, STIC established direct feedback between T_s, H 124 and LE, and simultaneously overcame the need of empirical parameterization for estimating the 125 conductances (Mallick et al., 2016, 2018a). Different versions of STIC have been extensively 126 validated in different ecological transects (Tropical rainforest to woody savanna) and aridity 127 gradients (humid to arid) (Trebs et al., 2021; Bai et al., 2021; Mallick et al., 2015a; 2016; 2018a, 128 b; Bhattarai et al., 2018, 2019). Based on the conclusions of Verhoef et al. (2012), Mallick et al. 129 (2014; 2015a,b; 2016; 2018a,b, 2022), Bhattarai et al. (2018, 2019), and Bai et al. (2021), there is 130 a need to address some of the challenges in SEB modeling, which are, (i) accurate estimation of G 131 and ET in sparse vegetation, (ii) testing the utility of coupling a TI-based G model with an 132 analytical SEB model for accurately estimating G and ET, and (iii) detailed evaluation of a coupled 133 G-SEB model at the ecosystem scale. Realizing the significance of mechanistic G model 134 (MV2007), the advantage of the analytical STIC model, and to mitigate some of the overarching 135 gaps in SEB modeling in sparsely vegetated open canopy systems, this study presents the first-136 ever coupled implementation of MV2007 G with the most recent version of STIC (STIC1.2). We 137 name this new coupled model as STIC-TI and it requires noon-night Ts and associated remotely 138 sensed land surface variables as inputs. We performed subsequent evaluation of STIC-TI in nine 139 terrestrial ecosystems in arid, semi-arid and sub-humid climate in India, the United States of 140 America (USA) (northern hemisphere) and Australia (southern hemisphere) at the eddy covariance 141 flux tower sites. The current study addresses the following research questions and objectives:

- (i) What is the performance of STIC-TI G estimates when compared with conventionally used empirical G models in ecosystems having low mean fractional vegetation cover (f_c) (≤ 0.5) and having larger soil exposure to radiation for example in Savanna?
- 145 (ii) How do the estimates from STIC-TI LE and H fluxes compare with LE and H observations in
- 146 diverse terrestrial ecosystems that represent a varied range of f_c (0.25 0.5) covering cropland,
- savanna, mulga vegetation (woodlands and open-forests dominated by the mulga tree [Acacia]
- 148 *aneura*]) spread across arid, semi-arid, sub-humid, humid climates over a vast range of rainfall
- 149 (250 to 1730 mm), temperature (-4 to 46°C) and soil regimes?
- (iii) What is the seasonal variability of G and evaporative fraction from STIC-TI model in a widerange of ecosystems having contrasting aridity and vegetation cover?

152 It is important to mention that assessing the performance of STIC-TI LE and H with respect to 153 other SEB models is not within the scope of the present study. The prime focus of the current study 154 is to assess the sensitivity of STIC-TI, temporal variability of the retrieved SEB fluxes, and cross-155 site validation of the individual SEB components.

156 A list of variables, their symbols and corresponding units are given in Table A1 in Appendix A.

157 2 Study area and datasets

158 **2.1Study site characteristics**

159 The present study was conducted using data from nine flux tower sites (four sites in India; three 160 sites in Australia; two sites in USA) equipped with Eddy Covariance (EC) measurement systems. 161 The distribution of the flux tower sites considered for the present study are shown in Fig.1 below. 162 The sites cover a wide range of climate, vegetation types, low fractional vegetation cover (f_c) of 163 around 0.5 and have contrasting aridity (Table 1). In India, a network of EC towers was set up 164 under Indo-UK INCOMPASS (INteraction of Convective Organization and Monsoon 165 Precipitation, Atmosphere, Surface and Sea) Program (Turner et al., 2019) at Jaisalmer (IND-Jai) in Rajasthan state, Nawagam (IND-Naw) in Gujarat state, Samastipur (IND-Sam) in Bihar state 166 167 and under Newton-Bhaba programme (Morisson et al., 2019 a,b) at Dharwad (IND-Dha) in 168 Karnataka state. The flux footprint for EC towers in India varied from 500 m - 1 km (Bhat et al., 169 2019). In the present study, about 90% of the fluxes came from an area within 500 m to 1 km from

170 the EC tower. Therefore, the relative contribution of vegetated land surface area to the fluxes is 171 close to 90% (Schmid, 2002; Vesala et al., 2008). The remaining percentage of fluxes originated 172 from an area beyond the flux footprint. The mean annual f_c was found to vary from 0.25 to 0.52 173 with standard deviation (SD) ranging from 0.1 to 0.16.

174 The IND-Jai site represents arid western zone over desert plains of natural grassland ecosystem. 175 The region receives very low rainfall (100 - 300 mm) during monsoon and experiences a wide 176 range in air temperature, high solar radiation, wind speed and high evaporative demand (Raja et 177 al., 2015). The IND-Naw site represents semi-arid agroecosystem in the middle Gujarat 178 agroclimatic zone of north-west India and has a pre-dominant rice-wheat cropping system. The 179 IND-Sam site has sub-humid climate of north-west alluvial plain zone in the Indo-Gangetic Plain 180 (IGP) situated in the eastern India and this site also follows rice-wheat crop rotation. IND-Dha 181 represents humid sub-tropical climate of transition zone in the southern India and this site 182 comprises of crops.

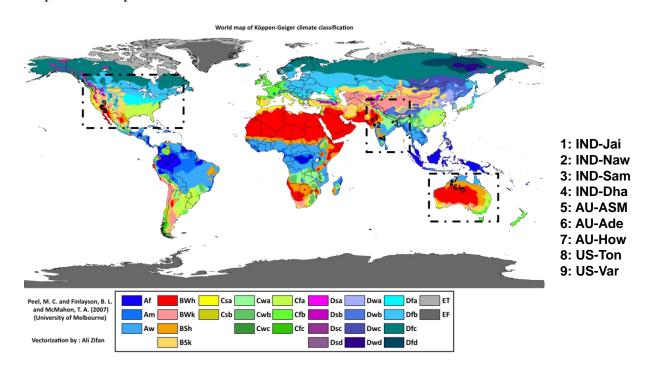


Figure 1: Locations of the flux tower sites in India, Australia and USA overlaid on climate type map. (Image Source: By Peel, M. C., Finlayson, B. L., and McMahon, T. A. (University of Melbourne) enhanced, modified, and vectorized by Ali Zifan; Hydrology and Earth System Sciences: "Updated world map of the Köppen-Geiger climate classification". Legend explanation, CC BY-SA 4.0, https://commons.wikimedia.org/w/index.php?curid=47086879)

183 In USA, two EC tower sites were located at Tonzi Ranch (US-Ton) and Vaira Ranch (US-Var), in 184 the lower foothills of the Sierra Nevada Mountains. Both the EC stations are part of the 185 AMERIFLUX Management Project (https://ameriflux.lbl.gov/). US-Ton is classified as an oak 186 savanna woodland. While the overstorey is dominated by blue oak trees (40% of total vegetation) 187 with intermittent grey pine trees (3 trees per hectare), the understory species include a variety of 188 grasses and herbs. The mean annual rainfall at this site is 559 mm. US-Var is a grassland dominated 189 site and the growing season is confined to the wet season only, typically from October to early 190 May. The mean annual rainfall at this site is 559 mm. The mean annual f_c was found to vary from 191 0.18 to 0.26 and SD of the order of 0.06 to 0.07.

192 In Australia, three EC tower sites were located at Howard Springs (AU-How), Alice Springs Mulga 193 (AU-ASM), Adelaide river (AU-Ade) in the Northern Territory as part of the OzFlux network 194 (Beringer et al., 2016) and the Terrestrial Ecosystem Research Network (TERN), which is 195 supported by the National Collaborative Infrastructure Strategy (NCRIS) 196 (http://www.ozflux.org.au/monitoringsites/index.html). The AU-How is situated in the Black 197 Jungle Conservation Reserve representing an open woodland savanna and the mean annual rainfall 198 is 1750 mm. The AU-ASM is located on Pine Hill cattle station near Alice Springs. The woodland 199 is characterized by mulga canopy and mean annual rainfall is 306 mm. AU-Ade represents savanna 200 with a mean annual rainfall of 1730 mm. The mean annual f_c varied from 0.21 to 0.48 having SD 201 range of 0.08 - 0.17. A description of Australian flux sites is given in Beringer et al. (2016). 202 Average heights of vegetation are 1.15 m at IND-Naw, 1 m at IND-Jai, 1.23 m at IND-Sam, 1.5 203 m at IND-Dha, 6.5 m at AU-ASM, 15m at AU-How, 7 m at AU-Ade, 10 m at US-Ton, and ≤ 0.5 204 m at US-Var.

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Hemisphere	Sites	Latitude (°N), Longitude (°E)	Climate & Vegetation	Mean f _c (SD)	Soil texture	T _A range (°C)	Mean Annual P (mm)	Observation period	
	Jaisalmer (IND-Jai)	26.99, 71.34	Arid grassland	0.25(±0.1)	Loamy fine sand to coarse sand	8-40	250	2017 - 2018	
	Nawagam (IND- Naw)	22.80, 72.57	Semi-arid cropland	0.41(±0.13)	Sandy loam	9 – 39	700	2017 - 2018	
Northern	Samastipur (IND- Sam)	26.00, 85.67	Humid subtropical cropland	0.52(±0.16)	Sandy loam to 10 – 39 loam		1000	2017 – 2018	
	Dharwad (IND-Dha)	15.50, 74.99	Tropical Savanna	0.36(±0.11)	Shallow to medium black clay and red sandy loam soils	12 - 40	650	2016 – 2018	
	Tonzi ranch (US-Ton)	38.43, -120.96	Woody Savanna	0.18(±0.06)	Red sandy clay loam	0-40	559	2011 - 2019	
	Vaira ranch (US-Var)	38.41, -120.95	Arid grassland	0.26(±0.07)	Rocky silt loam	0-40	559	2011 - 2019	
Southern	Alice Springs Mulga (AU- ASM)	-22.28, 133.24	Semi-arid mulga	0.21(±0.09)	Loamy sand	(-4) – 40	305	2011 - 2014	
	Howard Springs (AU-How)	-12.49, 131.15	Tropical savanna	0.48(±0.17)	Red kandasol	19 – 34	1700	2011 - 2014	
	Adelaide River (AU-Ade)	-13.07, 131.11	Savanna	0.42(±0.08)	Yellow hydrosol, shallow, loamy sand with coarse gravel	16 - 37	1730	2007 – 2009	

214 **<u>Table 1</u>**: An overview of the EC flux tower site characteristics in the present study

215 T_A: Air temperature during the observation period; P: rainfall (mm) measured using rain gauge at flux tower site during the study

216 period. IND is for India, AU is for Australia, and US is for the United States; SD is standard deviation of annual mean fc which is 217 computed from NDVI as mentioned in section 3.1.

218 2.2 Datasets

219 2.2.1 Micrometeorological data at flux tower sites

220 Standardized, controlled, and harmonized surface energy balance (SEB) flux and meteorological 221 data from nine EC towers were used in the present analysis. In Australia, H and LE were measured 222 through the EC systems and R_N were measured through net radiometers at varying heights of 15 223 m (AU-Ade), 23 m (AU-How), and 11.6 m (AU-ASM), respectively. In India, the EC 224 measurement height was maintained approximately at 8 m above the surface, except at IND-Dha 225 where it was installed at a height of 4.2 m. In USA, the SEB measurements were carried out at 226 tower heights of 23 m at US-Ton and 2 m US-Var. A summary of the instrumentation is given in 227 Table A2 of Appendix A. All the flux tower sites were equipped with a range of meteorological 228 instrumentation which measured diurnal air temperature (T_A) and relative humidity (R_H), four 229 components of the net radiation (R_N, consisting of down- and up-welling shortwave and long-wave 230 radiation (SW \downarrow , SW \uparrow , LW \uparrow and LW \downarrow , respectively)) above the vegetated canopy. In addition, the 231 diurnal soil heat flux (G) and soil temperature (T_{ST}) were measured at all the three Australian sites 232 and two US sites. In India, the diurnal soil heat flux was measured only at IND-Dha.

233 For the Indian sites, the raw EC measurements of the turbulent wind vectors (u, v and w, for234 horizontal, meridional and vertical, respectively), sonic temperature (T), and CO₂ and water vapor 235 mass density were recorded at a sampling rate of 20 Hz. Raw EC data were post-processed to 236 obtain level-3 quality controlled and harmonized surface fluxes at 30-minute flux averaging 237 intervals using EddyPRO® Flux Calculation Software (LI-COR Biosciences, Lincoln, Nebraska, 238 USA) using the data handling protocol described by Bhat et al. (2019). The EC data from the 239 OzFlux sites was averaged over 30 minutes recorded by the logger and processed through levels 240 using the PyFluxPro standard software processing scripts as mentioned in Isaac et al. (2017). The 241 Level 3 (L3) used in this paper was produced using PyFluxPro (Isaac et al., 2017) employing the 242 Dynamic INtegrated Gap filling and partitioning for Ozflux (DINGO) system as described in 243 Donohue et al. (2014) and Beringer et al. (2016). The quality checked EC data at 30 minute 244 intervals for two AMERIFLUX sites US-Ton and US-Var was acquired from 245 https://doi.org/10.17190/AMF/1245971&https://doi.org/10.17190/AMF/1245984, respectively.

246 2.2.2 Remote sensing data

247 Optical and thermal remote sensing observations available from Moderate Resolution Imaging Spectroradiometer (MODIS) (Didan et al., 2015) on-board Aqua platform were used in the present 248 249 study (Table 2) for estimating G and associated SEB fluxes. These include eight-day land surface 250 temperature (LST or T_S) at 1:30 pm and 1:30 am, and surface emissivity (ε_s) (MYD11A2), daily 251 surface albedo (α_R) (MCD43A3), 16-day NDVI (MYD13A2). The overpass times of MODIS 252 Aqua are at 1:30 pm and 1:30 am. The 8-day average values of clear-sky T_s available from 253 MYD11A2 data were used (Source: https://vip.arizona.edu/documents/viplab/MYD11A2.pdf) for 254 all nine flux tower sites. Since MYD21A2 LST product is known to provide better accuracy (1 – 255 1.5 K) (Hulley et al, 2016) as compared to MYD11A2 LST over semi-arid and arid ecosystems, 256 the former was also used in STIC-TI to compare LE and H estimates (see Table 5 in section 4.4) 257 with the estimates of MYD11A2 LST over the arid and semi-arid sites (IND-Jai, IND-Naw, US-258 Ton). The noon-night pair of thermal remote sensing observations from Aqua are close to the time 259 of occurrences of maximum and minimum soil surface temperature (see Figure 2) and are therefore 260 ideal for soil heat flux modeling using thermal inertia. The MODIS Terra overpass times are at 11 261 am and 11 pm and are far from the time of occurrences of minimum-maximum soil temperatures. 262 Therefore, MODIS Aqua acquisition times were used in the present study.

263	Table 2: A summary of MODIS Aqua optical and thermal remote sensing products used in the
264	present study

Data type	Product ID (version)	Variables used	Spatial resolution (m)	Temporal resolution	Purpose	Inputs to equation numbers
LST and emissivity	MYD11A2 (V006) MYD21A2	T_{s} (1:30 pm and am) and ε_{s}	923	8-day	For estimating R _{Ni} , G _i , LE _{i,} H _i	(5), (13), (C6), (C7), (B8)
Surface albedo	MCD43A3 (V006)	α _R	462	8-day composite from daily	For estimating R _{Ni} ,G _i	(5), (B3)

Vegetation	MYD13Q1	NDVI	250	16-day	For	(4)
index	(V006)				estimating	
					G_i	

265 **3 Methodology**

266 **3.1 Coupled soil heat flux-SEB model**

267 In this paper, we modified a thermal inertia (TI) based soil heat flux (G) model using noon-night 268 thermal remote sensing observations and thereafter coupled the TI-based G with STIC1.2. A clear-269 sky net radiation (R_N) model was also introduced into this coupled model and R_N estimation 270 algorithm is described in Appendix B. The estimation of G through modifying MV2007-TI 271 approach and its coupling with STIC1.2 is the most novel component of the modeling scheme, and 272 it is therefore described in the main body of the paper (section 3.1.1). Such a coupling enabled the 273 implementation of a mechanistic G model along with an analytical SEB model using optical-274 thermal remote sensing data. The coupled model is hereafter referred as STIC-TI.

275 **3.1.1 MV2007** soil heat flux model based on Thermal Inertia (TI)

The functional form for estimating instantaneous G (G_i , hereafter) (eq. 2 below) is based on the harmonic analysis of soil surface temperature and is described in detail by Murray and Verhoef (2007) and Maltese et al. (2013).

$$G_{i} = \Gamma \left[(1 - 0.5f_{C}) \left(\sum_{n=1}^{k} A \sqrt{n\omega} \sin \left(n\omega t + \phi_{n}' + \frac{\pi}{4} - \frac{\pi\Delta t}{12} \right) \right) \right] = \Gamma J_{S}$$
(2)

 G_i is the soil heat flux at the surface at a particular instance (Wm⁻²), Γ is the soil thermal inertia (J 279 m⁻² K⁻¹ s^{-0.5}), k is the total number of harmonics used, A is the amplitude (°C) of the nth soil 280 surface temperature (T_{ST}) harmonic, ω is the angular frequency (rads⁻¹), t is the time (s), ϕ'_n is the 281 phase shift of the nth soil surface temperature harmonic (rad), Is is the summation of harmonic 282 283 terms of soil surface temperature (K), and $\Delta t(s)$ is time offset between the canopy composite 284 temperature and the below-canopy soil surface temperature. Here, we represent G_i and A as 285 ecosystem-scale (\leq 1km) soil heat flux and surface soil temperature amplitude (averaged from soil 286 surface to 10 cm depth), respectively and assume it to be valid for different vegetated landscape.

Since we have considered a single pair (noon-night corresponding to 1:30 pm and 1:30 am) of MODIS aqua LST data in the present study, the phase shift (ϕ'_n) is taken as zero and number of harmonics is taken as one (k=1) for estimating G_i. Thus equation (2) is modified as follows:

$$G_{i} = \Gamma \left[(1 - 0.5f_{C}) \left(A \sqrt{\omega} \sin \left(\omega t + \frac{\pi}{4} - \frac{\pi \Delta t}{12} \right) \right) \right] = \Gamma J_{S}$$
(3)

With the two boundary values (i.e., $\Delta t = 1.5$ h for $f_c = 1$ and $\Delta t = 0$ for $f_c = 0$, f_c being the vegetation fraction), a linear approach is proposed here to describe the time offset Δt as a function of f_c (Maltese et al., 2013). For a given day, f_c was derived by normalizing NDVI with the upper-lower limits of annual NDVI cycle.

$$\Delta t = 1.5 f_c \tag{4}$$

294 3.1.1.1 Scaling function for estimating ecosystem-scale surface soil temperature amplitude (A)

295 Estimating ecosystem-scale A involves two steps, (a) computing point-scale soil surface 296 temperature amplitude (from surface to 0.1m depth) (T_{STA}, hereafter) from the available measurements of soil surface temperature, and (b) linking T_{STA} with remote sensing variables to 297 298 develop scaling functions for A. Point-scale soil temperature measured at different depths within 299 top 10 cm soil layer at AU-ASM, US-Ton, US-Var was averaged and considered as representative 300 surface soil temperature (0 - 10 cm). For Ind-Dha and AU-Ade, single-depth (10 cm) soil 301 temperature measurement was used. Studies also showed that LST carries some signal beneath the 302 skin of the surface (Johnston et al., 2022).

303 Several studies suggested theoretical sinusoidal trajectory of soil surface and sub-surface 304 temperatures (Gao et al., 2010), where the amplitude is maximum at the surface, and it gradually 305 decreases with depth to become 37% of surface amplitude until the damping depth (Hillel, 1982). 306 However, at deeper depths, soil temperatures remain constant with time and do not show many 307 fluctuations as compared to surface or near-surface soil temperatures. This invariant soil 308 temperature is called deep soil temperature (Mihailovic et al., 1999). However, the diurnal surface 309 soil temperature measurements (within top 0.1 m depth) across different flux tower sites showed 310 a sinusoidal-exponential behavior, i.e., sinusoidal pattern from sunrise until the afternoon and 311 exponential pattern from afternoon through sunset to the next sunrise. An illustrative example of the theoretical and observed trajectories of surface soil temperature is shown in Fig. 2. This diurnal surface soil temperature variation has a single harmonic component (Gao et al., 2010). For computing T_{STA} , theoretical half-curve of sinusoidal pattern is assumed and was derived from measurements as exemplified in Fig 2.

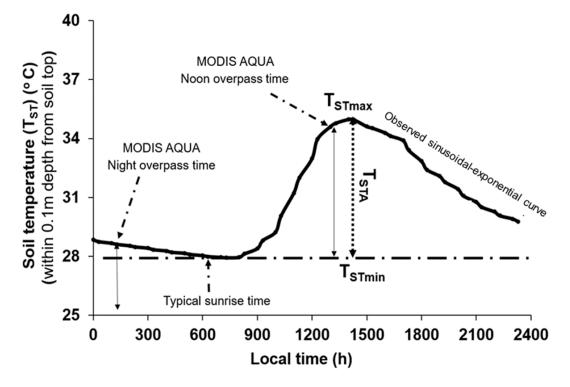


Figure 2. An illustrative example of typical diurnal variation of observed soil temperature (T_{ST}) (from surface to 0.1m depth) at OzFlux sites and timings of MODIS AQUA observations. Here, T_{STmax} and T_{STmin} are maximum and minimum point-scale observed soil surface temperatures

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317 It is evident from Fig. 2 that T_{STmin} represents minimum surface soil temperature occurring 1-1.5h 318 after sunrise and T_{STmax} occurs during 12.30 – 15.00h local time. The *in-situ* measured T_{ST} on 319 completely clear-sky days at OzFlux sites were used to extract T_{STmax} and T_{STmin} and T_{STA} was

derived as $(T_{STmax}-T_{STmin})$ from the theoretical half-curve of sinusoidal pattern.

321 T_{STA} is generally influenced by several land surface characteristics such as surface temperature 322 and surface albedo of soil-canopy complex, surface heat capacities, fractional canopy cover and 323 thermal conductivity (White, 2013). T_S and α_R are the major thermal and reflective land surface 324 properties that have strong synergy with surface soil temperature dynamics. Hence, we have used 325 bivariate regression analysis to develop a scaling function for estimating ecosystem-scale T_{STA} (top 326 to 0.1m depth). The bivariate regression is based on the difference of noon (d) and night (n) T_S 327 data and α_R (Duan et al., 2013, Li Tian et al., 2014) from MODIS Aqua. The scaling function given 328 in eq. (5) estimates ecosystem-scale T_{STA} (symbolized as 'A' in equation 5) from surface to 0.1 m 329 soil depth:

$$A = B_1(T_{Sd} - T_{Sn}) + B_2(\alpha_R) + B_3$$
(5)

Here, B1, B2, B3 are coefficients of regression model; T_{Sd} and T_{Sn} are noon and nighttime LST, respectively. The results of this regression analysis are elaborated in section 4.1.

332 3.1.1.2 Estimating Γ

333 Γ is the key variable for estimating G_i using eq. (2). MV2007 adopted the concept of normalized

thermal conductivity (Johansen, 1975) and developed a physical method to estimate Γ as follows:

$$\Gamma = e^{\left[\Upsilon' \left(1 - S_r^{(\Upsilon' - \delta)}\right)\right]} (\tau_* - \tau_0) + \tau_0$$
(6)

where τ_* and τ_0 are the thermal inertia for saturated and air-dry soil (J m⁻²K⁻¹s^{-0.5}); $\tau_0 = D_1\theta_* + D_2$; 335 $\tau = D_3 (\theta^{-1.29}); \Upsilon'(-)$ is a parameter depending on the soil texture (Murray and Verhoef, 2007; 336 Minasny, 2007; Anderson et al., 2007); $S_r(m^3 m^{-3})$ is relative saturation and is equal to (θ/θ_*) ; δ 337 (unitless) is the shape parameter which is dependent on the soil texture. θ * (m³ m⁻³) is the soil 338 porosity (equal to the saturated soil moisture content when soil moisture suction is zero), θ (cm³ 339 cm^{-3}) is the volumetric soil moisture and D₁, D₂, D₃ are coefficients which were derived from a 340 large number of experimental data. The reported global values of D₁, D₂, and D₃ were taken as -341 1062.4, 1010.8, 788.2, respectively (Maltese et al., 2013). The value for θ * and shape parameter 342 343 for soil textures across study sites were specified according to Van Genuchten et al. (1980). The 344 details are mentioned in Table E1 of Appendix E.

In the present study, the relative soil moisture saturation, $S_r(\theta/\theta^*)$ is represented in terms of an aggregated moisture availability (M) of canopy-soil complex through a linear function (eq. 12). In case of zero canopy cover, M represents the soil moisture availability from surface to 0.1 m depth. In sparse and open canopy, rates of moisture availability from soil to root and root to canopy were assumed same. Theoretically, M is expressed as available soil moisture fraction between field capacity (θ_{fc}) and permanent wilting (θ_{wp}) point as given in eq. (7) below.

$$M = \frac{\theta - \theta_{wp}}{\theta_{fc} - \theta_{wp}}$$
(7)

Where, θ_{fc} (m³ m⁻³) is the volumetric soil moisture at the field capacity (at a suction of 330 hPa) and θ_{wp} (m³ m⁻³) is the volumetric soil moisture at the permanent wilting point (at suction of 15000 hPa) (Singh, 2007). Since θ_{fc} , θ_* , θ_{wp} are characteristic volumetric soil moisture contents corresponding to specific suctions and depends on the soil texture, dividing the numerator and denominator in eq. (7) by θ_* gives the following expression:

$$M = \frac{\frac{\theta}{\theta_*} - \frac{\theta_{wp}}{\theta_*}}{\frac{\theta_{fc}}{\theta_*} - \frac{\theta_{wp}}{\theta_*}}$$
(8)

357 Due to their dependence on soil texture, the ratios (θ_{fc}/θ_*) and (θ_{wp}/θ_*) are treated as constants. 358 These are represented as C and C' in the later equations (eq. 9, 10, and 11). The constants, C and 359 C' vary from 0.3 to 0.8 and from 0.1 to 0.4 (Murray and Verhoef, 2007; Minasny et al., 2011; 360 Anderson et al., 2007), respectively over different soil textures.

$$M = \frac{\frac{\theta}{\theta_*} - C'}{C - C'}$$
(9)

$$M(C - C') = \left(\frac{\theta}{\theta_*}\right) - C'$$
(10)

By replacing S_r in eq. (6) as θ/θ_* and by rearranging eq. (10), the following linear function is obtained.

$$S_{r} = \frac{\theta}{\theta_{*}} = M (C - C') + C' = M'$$
(11)

363 Thus, the modified equation to calculate Γ is given by eq. (12) as follows:

$$\Gamma = e^{\left[Y'\left(1 - M'^{(Y'-\delta)}\right)\right]}(\tau_* - \tau_0) + \tau_0$$
(12)

By substituting the values obtained from eq. (4), (5) and (12) into eq. (3), we obtained the instantaneous ecosystem-scale G_i corresponding to MODIS Aqua noontime overpass. The intrinsic link between G_i estimates through MV2007-TI and SEB scheme in STIC1.2 is made through M, where the computation of M follows the procedure as described in Mallick et al. (2016, 2018a, b) and Bhattarai et al. (2018). (description in Appendix C).

369 3.1.1.3 Estimating M

In STIC1.2, an aggregated moisture availability (M) of canopy-soil complex is expressed as the ratio of the 'vapor pressure difference' between the aerodynamic roughness height of the canopy (i.e., source/sink height) and air to the 'vapor pressure deficit' between aerodynamic roughness height to the atmosphere:

$$M = \frac{(e_0 - e_A)}{(e_0^* - e_A)} = \frac{(e_0 - e_A)}{\kappa(e_S^* - e_A)} = \frac{s_1(T_{0D} - T_D)}{\kappa s_2(T_S - T_D)}$$
(13)

Where e_0 and e_0^* are the actual and saturation vapor pressure at the source/sink height; e_A is the 374 atmospheric vapor pressure; e_s^* is the saturation vapor pressure at the surface; T_{0D} is dew point 375 376 temperature at the source/sink height; T_S is the LST; T_D is the air dew point temperature; s₁ and s₂ 377 are the psychrometric slopes of the saturation vapor pressure and temperature between $(T_{0D} - T_D)$ versus $(e_0 - e_A)$ and $(T_S - T_D)$ versus $(e_S^* - e_A)$ relationship; and κ is the ratio between $(e_0^* - e_A)$ 378 379 and $(e_s^* - e_A)$. To solve the eq. (13), estimation of T_{0D} is necessary. An initial estimate of T_{0D} [T_{0D} = $[(e_S^* - e_A) - s_3T_S + s_1T_D]/(s_1 - s_3)]$ and M were obtained following Venturini et al. (2008) where 380 381 s₁ and s₃ were approximated in T_D and T_S, respectively. However, eq. (13) cannot be directly 382 solved because there are two unknowns in one equation. However, since T_{0D} also depends on LE 383 (Mallick et al., 2016, 2018a), an iterative updation of T_{0D} (and M) was carried out by expressing 384 T_{0D} as a function of LE $[T_{0D} = T_D + (\gamma LE/\rho c_p g_A s_1)]$ which is described in detail by Mallick et al. 385 (2016, 2018a) and Bhattarai et al. (2018). In the numerical iteration, s_1 was not updated to avoid 386 numerical instability and it was expressed as a function of T_D.

387 3.1.2 STIC-TI: Coupling modified MV2007-TI and STIC 1.2

388 The initiation of the coupling between MV2007-TI and STIC1.2 was executed through linking G_i 389 estimates from the modified MV2007-TI with M estimates from STIC1.2. Having the initial 390 estimates of M (through eq. 13), an initial estimation of G_i was made from eq. (2) where S_r in eq. 391 11 was replaced with the initial estimates of M'. From the initial estimates of G_i (eq. 2) and R_{Ni} 392 (equations in Appendix B), initial estimates of LE_i and H_i were obtained through the Penman-393 Monteith Energy Balance (PMEB) equation. Analytical expressions of the conductances for 394 estimating H and LE through the PMEB equation were obtained by solving the state equations as 395 described in the Appendix. The process was then iterated by updating T_{0D} [$T_{0D} = T_D$ + 396 (yLE /pcpgAS1)] and M in every time step (as mentioned in Mallick et al., 2016, 2018a), and re-397 estimating G_i (using eq. 3), net available energy ($R_{Ni}-G_i$), conductances, LE_i and H_i, until stable 398 estimates of LE_i were obtained. The conceptual block diagram and algorithm flow of STIC-TI is 399 shown in Fig. 3a and Fig 3b, respectively.

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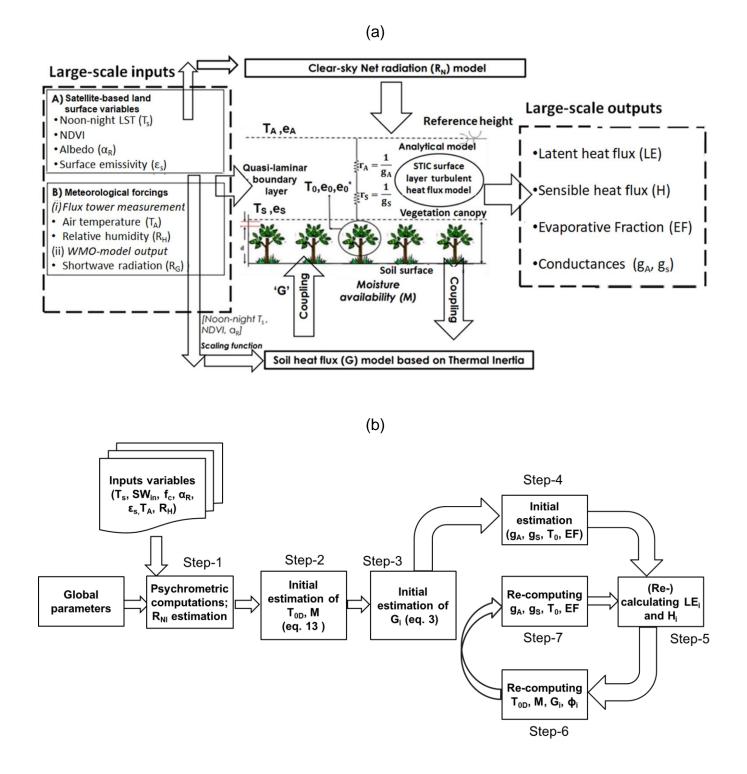


Figure 3: (a) Conceptual diagram of STIC-TI model showing different input variables and model outputs, (b) Algorithmic flow for estimating G and associated SEB fluxes through STIC-TI.

Examples of iterative stabilization of G_i and LE_i for Indian, Australian and US ecosystems of India
are shown in Fig. 4. The iterative stabilization of G_i and LE_i was obtained between 8-25 iterations
for all sites.

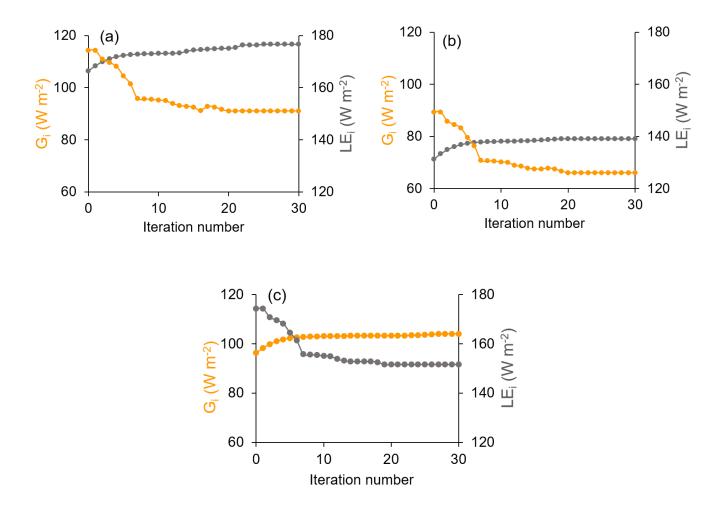


Figure 4: Illustrative examples of iterative stabilization of STIC-TI G_i (yellow marker line) and LE_i (grey marker line) in (a) IND-Jai, (b) AU-ASM, (c) US-Ton

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410 The noteworthy features of STIC-TI are: (1) estimating G by modifying the mechanistic MV2007-

411 TI model using noon and midnight T_s information from thermal remote sensing observations

412 available through polar orbiting satellite platform (e.g. MODIS Aqua), (2) coupling the

413 mechanistic MV2007-TI G model with STIC1.2 to simultaneously estimate surface moisture

414 availability (M), G, and SEB fluxes, (3) introducing water stress information in G (through M) to

415 better constrain the aerodynamic and canopy-surface conductances as well as the SEB fluxes, and

416 (4) derivation of amplitude of ecosystem-scale surface soil temperature (from top soil to 0.1 m soil417 depth).

418 **3.1.3 Generation of remote sensing inputs**

419 Two of the key variables in SEB modeling are Ts and ε_s . These two variables were retrieved at 420 923m spatial resolution from MODIS Aqua noon-night TIR observations (MYD11A2) in bands 421 11.03 µm and 12.02 µm using a generalized split-window algorithm (Wan et al., 2015). For 422 optimal retrieval, tractable sub-ranges of atmospheric column water vapor and lower boundary air 423 surface temperature were used. Land surface emissivity was estimated from land cover types and 424 anisotropy factors. The MYD21A2 LST product was generated using Temperature-Emissivity 425 Separation (TES) algorithm (Hulley et al, 2016) and improved water vapor scaling method to 426 remove the atmospheric effects. Albedo was estimated from MODIS (MCD43A2 Version 6.0) 427 Bidirectional Reflectance Distribution Function and Albedo (BRDF/Albedo) daily dataset (Schaaf 428 et al., 2002)) at 462 m spatial resolution. Actual albedo is a value which is interpolated between 429 white-sky and black-sky albedo as a function of fractional diffuse skylight (which is a function of 430 aerosol optical depth). NDVI was obtained from level 3 MODIS vegetation indices product 431 (MYD13Q1, version 6.1), which are generated every 16-day at 250 meter (m) spatial resolution. 432 All the input remote sensing variables mentioned in table 2 were resampled to spatial resolution 433 of MYD11A2 product (923 m).

434 **3.2 Sensitivity and statistical analysis**

435 The accuracy of STIC-TI heavily depends on the accuracy of T_S , NDVI, and α_R due to the dual 436 role of T_S in estimating M and G_i, the role of NDVI in G_i, and the combined role of T_S and α_R in 437 estimating R_{Ni} . Therefore, one-dimensional sensitivity analysis was conducted to assess the 438 impacts of uncertainty in T_s, NDVI and α_R on G_i, H_i and LE_i. The sensitivity was assessed by 439 varying noon-time T_S by ± 0.5 K, ± 1.0 K and ± 1.5 K (keeping nighttime T_S constant so that 440 amplitude can vary automatically); varying NDVI by ± 0.05 ; ± 0.10 , ± 0.15 ; and varying albedo by 441 $\pm 0.02, \pm 0.05, \pm 0.10$, respectively. SEB fluxes were computed by using T_s, NDVI, and α_R for three 442 different periods of the year in all the eight ecosystems. Sensitivity analyses were conducted by 443 increasing and decreasing systematically T_s , NDVI, α_R from its central value while keeping the 444 other variables and parameters constant. This procedure was selected because the fluxes and

intermediate outputs of the STIC-TI model reflect an integrated effect due to uncertainty in T_S. In the first run, SEB fluxes were computed using *in-situ* T_S measurements obtained from the flux tower outgoing longwave radiation measurements. Then T_S was increased and decreased at constant interval and a new set of fluxes were estimated. In the similar way, α_R and NDVI were increased and decreased at constant intervals and new set of fluxes were computed. The sensitivity of STIC-TI was assessed by the equation 14.

Sensitivity =
$$\frac{E_{i0} - E_{iM}}{O_i} * 100$$
 (14)

451 E_{i0} is the estimated (original) model output and E_{iM} is the estimated (modified) output obtained by 452 changing the variable whose sensitivity is to be tested. O_i is actual measurements. Apart from the 453 sensitivity analysis, the following set of statistical metrics were used to assess model performances.

$$R^{2} = \left(\frac{\sum_{i=1}^{n} (E_{i} - \overline{E}) (0_{i} - \overline{0})}{\sqrt{\sum_{i=1}^{n} (E_{i} - \overline{E})^{2}} \sqrt{\sum_{i=1}^{n} (0_{i} - \overline{0})^{2}}}\right)^{2}$$
(15)

RMSE =
$$\sqrt{\sum_{i=1}^{n} \frac{(E_i - O_i)^2}{n}}$$
 (16)

$$BIAS = \frac{\sum_{i=1}^{n} (E_i - O_i)}{n}$$
(17)

$$MAPD = \frac{100}{n} \sum_{i=1}^{n} \left| \frac{E_i - O_i}{O_i} \right|$$
(18)

$$KGE = 1 - \sqrt{(r-1)^2 + \left(\frac{\sigma_E}{\sigma_o} - 1\right)^2 + \left(\frac{\overline{E}}{\overline{0}} - 1\right)^2}$$
(19)

454 Where R^2 is the coefficient of determination, RMSE is root-mean-square error, BIAS is the mean 455 bias, MAPD is the mean absolute percent deviation, KGE is Kling-Gupta efficiency, n is the total 456 number of data pairs, the bar indicates mean value of the measured variable and model estimates 457 of the same variable. E_i and O_i are the model estimated and measured SEB fluxes, r is the Pearson's 458 correlation coefficient and \overline{O} is the average of measured values and \overline{E} is the average of estimated

459 values and σ_o is standard deviation of observation values and σ_E is the standard deviation of 460 estimated values. The KGE has been widely used for calibration and evaluation hydrological 461 models in recent years and it combines the three components of Nash-Sutcliffe efficiency (NSE) 462 of model errors (i.e., correlation, bias, ratio of variances or coefficients of variation) in a more 463 balanced way. But it has not been widely used for analyzing the ET model performances. KGE = 1464 indicates perfect agreement between modeled estimates and observations. The performance of a 465 model is considered 'poor' for KGE between 0 and 0.5 and models with negative KGE values is 466 considered 'not satisfactory'.

467 **4 Results**

468 **4.1 Ecosystem- scale surface soil temperature amplitude (A)**

469 The scaling functions developed to estimate ecosystem-scale (1km) surface soil temperature 470 amplitude (A) from point-scale T_{STA} were used to estimate G_i. However, before the development 471 of the scaling functions, analysis was carried out to investigate the relationship of soil temperature 472 amplitude between the two different spatial scales. The scatterplot (Fig. 5a) of noon-night LST 473 difference (ΔTs) versus T_{STA} for different albedo classes showed a linear increase in ΔTs with 474 increasing T_{STA}. However, some divergence of data points within the cluster were also noticed 475 which could be associated with different albedo (α_R) levels. Bivariate linear function was fitted 476 between T_{STA} as predictand (Y) versus ΔTs (T_{sd} - T_{sn}) and α_R as predictors (X1 and X2, 477 respectively). The function was found to be Y = 0.59X1 - 51.3X2 + 8.66 by combining the data 478 of nine ecosystems (r = 0.86). The coefficients in the above expressions correspond to B1 (0.59), 479 B2 (51.3), B3 (8.66) of eq. 5 in section 3.1.1.1. The estimated amplitude from this ecosystem-scale 480 predictors and scaling functions was treated as ecosystem-scale surface soil temperature amplitude 481 (A).

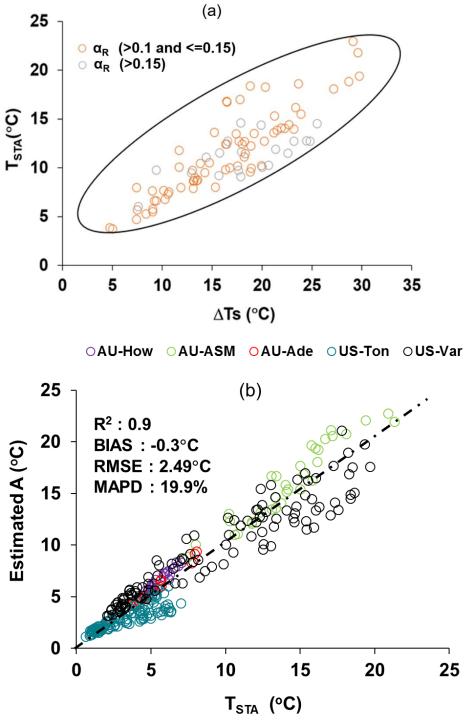


Figure 5. (a) Two-dimensional scatterplots between (ΔTs) versus T_{STA} at different α_R levels over different ecosystems. Here T_{STA} in y-axis is the observed soil temperature amplitude that is used to develop the scaling function and delta ΔTs is noon-night LST difference of MODIS AQUA; (b) Validation of ecosystem-scale estimates of A from the above functions over different sites.

482 Validation of ecosystem-scale estimates of A from the above functions over different sites is 483 shown in Fig. 5b with respect to T_{STA} for the independent datasets. The estimated A was found to 484 have MAPD of 19.9%, negative bias, and $R^2 = 0.90$ over different ecosystems. The temporal 485 variation of estimated A and T_{STA} is shown in Fig D1 in Appendix D. Further analysis was carried out to investigate the bias in A at three fractional vegetation cover (f_c) classes ($f_c < 0.3$; $0.3 \le f_c$ 486 487 ≤ 0.5 ; f_c >0.5) representing bare soil (class 1), 30 – 50% canopy cover (class 2) and more than 50% 488 canopy cover (class 3), respectively. While negative bias was noted for class 1 and class 3 (-0.54°C 489 and -0.83° C), the bias was positive (0.49° C) in the intermediate f_c which represents sparse and 490 patchy canopy cover. The signals of surface albedo, emissivity and temperatures of soil surface 491 and canopy are relatively pure in class 1 and class 3 as compared to class 2, where the surface 492 signal carries more heterogeneity. Given T_{STA} is computed from the in-situ measurements, it is 493 likely to carry more heterogeneity in class 2 as compared to the other two classes. The land surface 494 emissivity in MYD11A2 was estimated from land cover types and anisotropy factor, which have 495 differential impacts on T_{ST} and T_S leading to such opposite bias in class 2 as compared to class 1 496 and class 3.

497 4.2 Sensitivity analysis of STIC-TI Gi, LEi and Hi to land surface variables

498 **4.2.1 Sensitivity of Gi to land surface variables**

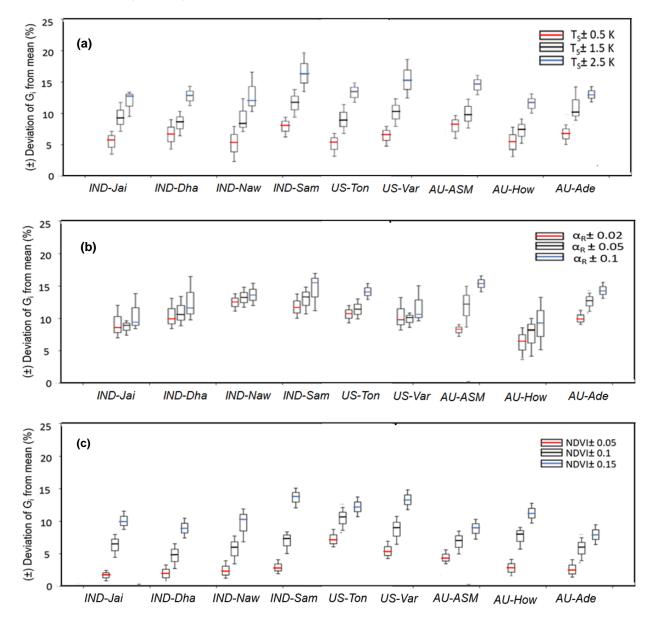
499 The average sensitivity of G_i to three land surface variables (T_s, NDVI, α_R) by combining the 500 estimates of wet and dry periods is shown in Fig. 6. G_i was found to be substantially sensitive to 501 T_s with error magnitude ranging from 2 – 18% due to T_s uncertainties of $\pm 0.5 - 2.5$ K (Fig. 6a), 502 with greater sensitivity to T_S during the summer season. The median sensitivity of G_i due to $\pm 5 -$ 503 10% uncertainty in α_R varied from 5 to 12% in all the ecosystems (Fig. 6b). The uncertainties in 504 NDVI revealed 2 to 15% error in G_i estimates (Fig. 6c), and no significant difference in the mean 505 sensitivity due to NDVI uncertainties was noted between the ecosystems. The sensitivity of G_i 506 decreased with increasing values of NDVI.

507 4.2.2 Sensitivity of LE_i and H_i to land surface variables

508 Both LE_i and H_i were sensitive to T_s to the order of 2 - 29% (LE_i) and 5 - 35% (H_i) for T_s 509 uncertainty of $\pm 0.5 - 2.5$ K from its mean values (Table 3). Interestingly, LE_i was more sensitive 510 to T_s uncertainties as compared to H_i in the rainfed ecosystems. The highest mean sensitivity of

25

511 LE_i to T_s was found in arid (IND-Jai: 2 - 28%), semi-arid (AU-ASM: 5 - 21%), tropical savanna 512 (IND-Dha: 3 – 26%), savanna (US-Ton: 4-29%) and arid (US-Var: 3-26%) ecosystems. The mean 513 sensitivity of H_i to T_S was maximum in sub-humid (IND-Sam: 2 – 32%), semi-arid (IND-Naw: 2 514 -28%), savanna (AU-Ade: 8 – 17%) (Table 3). A greater sensitivity of the SEB fluxes due to α_R 515 uncertainties was found than due to NDVI. The median sensitivity of LE_i and H_i due to 10% 516 uncertainty from mean α_R varied within 2 – 16% in all the ecosystems (Table 3). By contrast, 517 errors in the two SEB fluxes were substantially low (2 - 13%) due to $\pm 0.05 - 0.15$ uncertainty 518 from mean NDVI (Table 3).



26

Figure 6: Sensitivity of STIC-TI G_i due to uncertainties in T_S (a), α_R (b), and NDVI (c) for eight flux tower sites in India and Australia. The uncertainties were introduced by taking the mean values of these variables during three different periods (summer, rainy and winter) of a year. Mean uncertainties of the three periods are presented in the figure.

	Sensitivity of LE _i and H _i to T _S , NDVI and α_R (percent change)								
	T _s unc	ertainty	$\alpha_{\rm R}$ unce	ertainty	NDVI uncertainty $(\pm 0.05 - 0.15)$				
Study sites	(±0.5 -	- 2.5 K)	(±5–	10%)					
-	LEi	H _i	LEi	H _i	LEi	H _i			
IND-Jai	2-28	1-6	3-14	2-13	2-8	2-6			
IND-Dha	3-26	2-8	2-12	3-12	3-10	3-9			
IND-Naw	1-20	2-28	2-10	3-10	2-7	2-6			
IND-Sam	1-16	5-32	4-13	6-11	2-5	2-7			
US-Ton	4-29	4-12	3-12	4-12	3-8	5-7			
US-Var	3-26	6-14	4-11	2-10	4-10	2-8			
AU-ASM	5-21	2-10	3-12	2-13	2-10	2-11			
AU-How	8-13	2-15	2-11	4-16	3-12	3-13			
AU-Ade	2-17	8-17	3-12	2-10	3-10	3-9			

519 **<u>Table 3</u>**: Sensitivity (in percent) of LE_i and H_i due to T_S, NDVI, and α_R uncertainties

520 **4.3 Comparative evaluation of STIC-TI and contemporary Gi models**

521 The performances of STIC-TI and existing G_i models were evaluated and compared with respect 522 to *in-situ* G_i measurements. The existing models reported by Moran et al. (1989), Bastiaanssen et 523 al. (1998), Su (2002), and Boegh et al. (2004) have been considered for comparing with TI-based 524 model. These four existing models are referred here as MOR89, BAS98, SU02 and BO04, 525 respectively. While the models MOR89, SU02 and BO04 are based on linear regression between 526 G versus NDVI, BAS98 is based on multivariate regression of G with NDVI, LST and α_R . The 527 performance of the STIC-TI was substantially better as compared to MOR89, SU02 and BO04 528 with respect to MAPD (19%), RMSE (22 Wm⁻²) and coefficient of determination ($R^2 = 0.8$) when 529 compared with in-situ measurements over one Indian, three Australian and two US flux tower sites

530 (Table 4) and comparable with BAS98 G_i model. The validation plot of retrieved noontime Gi

from STIC-TI is shown in Fig. 7.

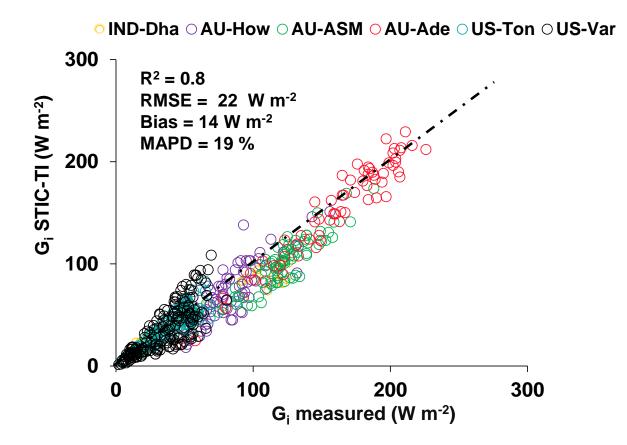


Figure 7: Validation of noontime (1:30 pm) G_i estimates with respect to in-situ measurements in different ecosystems. The regression between the two sources of G_i is G_i (STIC-TI) = 0.90 G_i (tower) -0.10.

532 <u>**Table 4**</u>: A comparison of error statistics of G_i estimates from STIC-TI and existing G_i models 533 over different ecosystems

G models	R ²	RMSE (W m ⁻²)	MAPD (%)	KGE
STIC-TI	0.80	22	19	0.74
MOR89	0.70	31	29	0.46
BAS98	0.80	20	18	0.61
SU02	0.80	30	26	0.54
BO04	0.70	35	29	0.48

534 The RMSE varied from 9 to 20 W m⁻² with MAPD ranging from 12 to 21% across individual flux

535 tower sites. High magnitude of G_i was predicted in the arid and semi-arid systems (120 – 240 W

536 m⁻²) as compared to the humid systems ($20 - 90 \text{ W m}^{-2}$), which was in close correspondence with 537 the observations. The model also captured the range of G_i that are generally found in different 538 biomes ($20 - 140 \text{ W m}^{-2}$ for grasslands, $20 - 90 \text{ W m}^{-2}$ for cropland) (Purdy et al., 2016). Due to 539 the paucity of G_i measurements, direct validation of G_i was only possible for 32 days (concurrent 540 to MODIS overpass) at the IND-Dha site. Overall, STIC-TI tends to provide reasonable G 541 estimates for the terrestrial ecosystems having soil temperature amplitude above 5°C.

542 4.4 Evaluation of STIC-TI LE_i, H_i, and EF

- 543 The modeled versus measured LE_i and H_i showed good agreement in all the nine ecosystems with
- 544 RMSE in LE_i and H_i estimates using MYD11 LST product to the order of 29 62 W m⁻² and 26 62
- 545 61 W m⁻², MAPD of 9 31% and 20 36%, BIAS of -29 to 38 W m⁻² and -44 to 32 W m⁻² (Fig.
- 546 8a, b; Table 5) and high R^2 of 0.8.

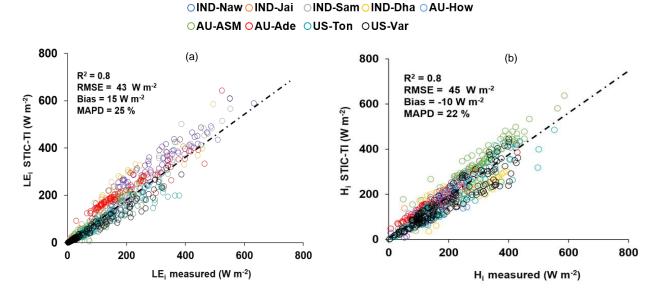


Figure 8: (a) Validation of STIC-TI LE_i estimates with respect to *in-situ* measurements in different ecosystems; (b) Validation of STIC-TI H_i estimates with respect to *in-situ* measurements in different ecosystems.

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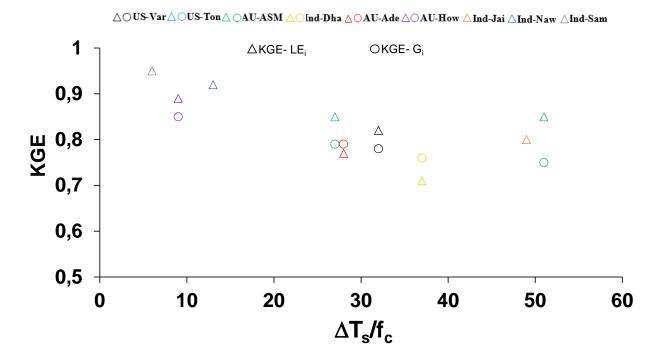
- 548 <u>**Table 5**</u>: Error statistics of STIC-TI LE_i and H_i estimates with respect to EC measurements in
- 549 different ecosystems of India, US, and Australia using MYD11A2 LST product for all nine sites
- and using MYD21A2 LST product for three semi-arid and arid sites. The statistics obtained by
- using MYD21A2LST are shown in the parentheses.
- 552

Sites	STIC-TI (LE _i and H _i)									
	\mathbb{R}^2		BIAS		RMSE		MAPD		KGE	
			(W m ⁻²)		(W m ⁻²)		(%)			
	LEi	Hi	LEi	Hi	LEi	Hi	LEi	Hi	LEi	Hi
IND-Jai	0.90	0.90	-21	12	57	27	31	22	0.80	0.76
IND-Jai	(0.91)	(0.92)	(-16)	(9)	(45)	(21)	(24)	(19)	(0.82)	(0.79)
IND-Naw	0.90	0.80	19	-26	44	51	17	28	0.92	0.71
IIND-INAW	(0.92)	(0.85)	(12)	(-16)	(37)	(46)	(16)	(25)	(0.92)	(0.73)
IND-Dha	0.90	0.90	38	-44	43	35	27	25	0.71	0.64
IND-Sam	0.90	0.80	12	-10	32	61	9	27	0.95	0.70
US-Ton	0.90	0.90	-29	-32	53	34	25	17	0.85	0.91
03-101	(0.91)	(0.92)	(-18)	(-21)	(45)	(27)	(22)	(15)	(0.87)	(0.93)
US-Var	0.90	0.80	-19	-28	49	39	27	20	0.82	0.89
AU-ASM	0.90	0.90	-3	22	46	26	29	20	0.94	0.83
AU-ASM	(0.93)	(0.91)	(6)	(16)	(37)	(18)	(24)	(17)	(0.95)	(0.85)
AU-How	0.90	0.90	16	-25	42	27	17	21	0.89	0.85
AU-Ade	0.90	0.90	21	15	29	53	28	36	0.77	0.80

⁵⁵³

554 Arid ecosystems in India (IND-Jai), US (Ton and Var) and semi-arid ecosystem in Australia (AU-555 ASM) revealed relatively high MAPD (31%, 25%, 27%, and 28%) (Table 5). In general, STIC-TI 556 was able to produce the dominant convective heat fluxes with respect to the EC measurements as 557 evident through low RMSE for H_i and high RMSE for LE_i in the IND-Jai, US-Ton, US-Var, and 558 AU-Ade where LE_i is inherently low except few rainy days. A uniform distribution of data points 559 around 1:1 validation line (Fig. 8a) indicated overall low BIAS in LE_i estimates. However, 560 modeled H_i was consistently lower than the observations (negative BIAS) in the tropical savanna 561 (IND-Dha and AU-How) and semi-arid (IND-Naw) ecosystems [(-44) – (-25) W m⁻² and -26 W m⁻²) while a consistent positive BIAS was observed in the AU-ASM (semi-arid) and AU-Ade 562 (savanna), US-Var (arid) (Fig. 8b; Table 5). This consequently led to overall low negative BIAS 563 (-10 W m⁻²), relatively low R² in H_i (R² = 0.8) as compared to the errors in LE_i (BIAS = 15 W m⁻²) 564 565 ², $R^2 = 0.9$). The regression between the modeled and tower measurements of LE_i is LE_i(STIC-TI) 566 $= 0.98 LE_i(tower) - 0.266$. The regression between the modeled and tower measurements of H_i is 567 H_i (STIC-TI) = 0.93 H_i (tower) + 4.90. The KGE statistics varied in the range of 0.71 – 0.95 for LE_i 568 and in the range of 0.64 - 0.91 for H_i, respectively across all nine flux tower sites, thus revealed 569 reasonably high efficiency of the model to capture the magnitude and variability of SEB fluxes.

570 The impact of MODIS Aqua day-night view angle difference (δ VZA) on STIC-TI fluxes was 571 further investigated. Estimated errors in terms of mean percent deviation in LE_i, H_i and G_i with respect to measurements for each 10° bin over 16 angular bins within ±80° were analysed in 572 573 response to mean δVZA of each angular bin. G_i errors (X) were found to be significantly correlated with δVZA (Y) in a parabolic (Y = $0.0027X^2 - 0.0025X + 1.4919$; r = 0.73) pattern (Refer 574 575 Appendix F, Figure F(a)). Errors in G_i to the order of 5 to 10%, 10-15% and >15% were largely 576 found to be within $\pm 30^{\circ}$, $\pm 45^{\circ}$, and $>45^{\circ}$ to -80° δ VZA, respectively. The errors in H_i were found to have strong linear (Y = -0.1452X + 1.1146, r = 0.77) dependence on δ VZA (Refer Appendix F, 577 578 Figure F(b)). However, a weak dependence of LE_i errors (Y = -0.0878X + 2.0314, r = 0.5) on 579 δ VZA (Refer Appendix F, Figure F(c)) was found, as majority of the errors were within ±10% that 580 corresponded to $\pm 60^{\circ}$ δ VZA. The nature of relations and degree of dependency of model flux 581 errors on δVZA in this study would be helpful to minimize the error budget in surface energy 582 balance fluxes from future thermal infrared missions having day-night observations.



<u>Figure 9</u>: Relationship between KGE of STIC-TI (G_i and LE_i) with $\Delta T_s/f_c$ in different terrestrial ecosystems.

583 Further investigation was made on whether KGE for STIC-TI G_i and LE_i follow any systematic

pattern and the ratio ΔT_s and f_c was used as proxy for surface heterogeneity and dryness. The plot

585 of KGE of G_i and LE_i with this ratio is shown in Fig. 9. KGE-G_i was found to show a systematic 586 decrease with increase in Δ Ts-fc ratio up to 40, after which it remained unchanged with increase 587 in the ratio. Although KGE of LE_i also decreased (20% reduction) with increase in Δ Ts-f_c ratio, 588 KGE-LE_i was found to increase beyond $\Delta Ts-f_c$ 40. This revealed that the model efficiency 589 remained high (>0.8) within certain dryness limits (Δ Ts-f_c ratio <20 and >50) and the efficiency 590 reduced moderately (within 0.7 - 0.8) for intermediate dryness. Interestingly, the use of 591 MYD21A2 LST in STIC-TI showed improvements (see the parentheses in different columns in 592 Table 5) in LE_i and H_i error statistics as compared to using MYD11A2 LST in terms of higher R^2 593 and KGE, and lower RMSE in LE_i (3-8% less) and H_i (2-3% less) for semi-arid and arid sites such 594 as IND-Jai, IND-Naw and US-Ton.

An independent evaluation of multi-temporal heat fluxes over two US flux sites for the years 2016-2018 is shown in Fig. 10 and Fig 11. STIC-TI G_i estimates with MYD11A2 LST product showed close match with *in-situ* measurements with respect to intra and inter-annual variability in G_i followed by LE_i and H_i. This further demonstrates the merit of the coupled model for reproducing ecosystem-scale G_i estimates especially for shorter and open canopies.

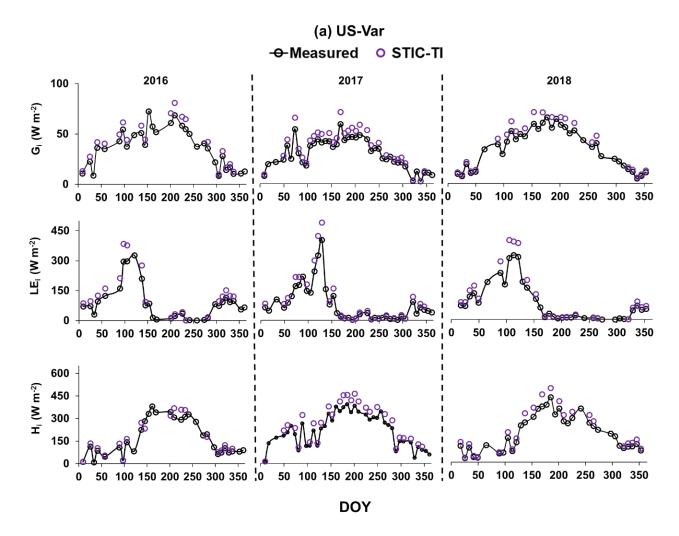


Figure 10: Illustrative examples of temporal evolution of STIC-TI derived fluxes using MYD11A2 LST product versus observed SEB fluxes for three consecutive years from 2016 to 2018 in a grassland ecosystem in United States (e.g., US-Var).

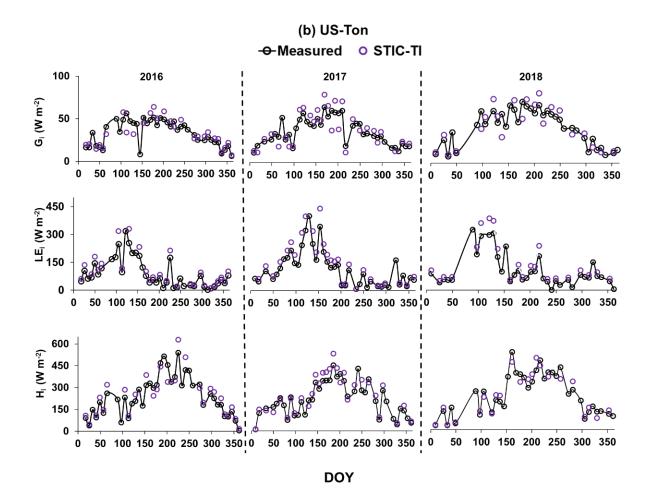


Figure 11: Illustrative examples of temporal evolution of STIC-TI derived fluxes using MYD11A2 LST product versus observed SEB fluxes for three consecutive years from 2016 to 2018 in a woody savanna ecosystem in the United States (e.g., US-Ton).

600 The temporal behavior of STIC-TI and observed evaporative fraction (EF) (ratio of LE and R_N – 601 G) (Fig. 12) along with observed monthly rainfall (P) distinctly captured the substantial temporal 602 variability in EF during the dry-to-wet transition in the Indian study sites, which also corresponded 603 to low (high) θ and P. In IND-Naw and IND-Sam, a marked rise (>0.4) in STIC-TI EF was noted 604 during day-of-the-year (DOY) 25 to 75 where wheat is grown under assured irrigation. The impact 605 of irrigation is thus captured by the substantial increase in EF in the absence of P. In contrast, the 606 rainfed grassland system (IND-Jai) showed peak EF (~0.8), which corresponded to south-west 607 monsoon rainfall during June to September and a progressive decline in EF during the dry down 608 period in October to April corresponding to post south-west monsoon phase. Some intermittent 609 spikes in EF were also noted during dry-down phase in both STIC-TI and observations. The 610 intermittent EF spikes during the soil moisture dry down phase could be due to enhanced LE 611 through moisture advection from the surrounding vegetation causing an enhancement of 612 evaporation than expected. This is known as the 'clothesline effect' which frequently occurs in 613 semi-arid and arid ecosystems. In addition to IND-Jai, the response of both modeled and measured 614 EF to wet and dry spells was also noted during south-west monsoon period at all other flux tower 615 sites of India.

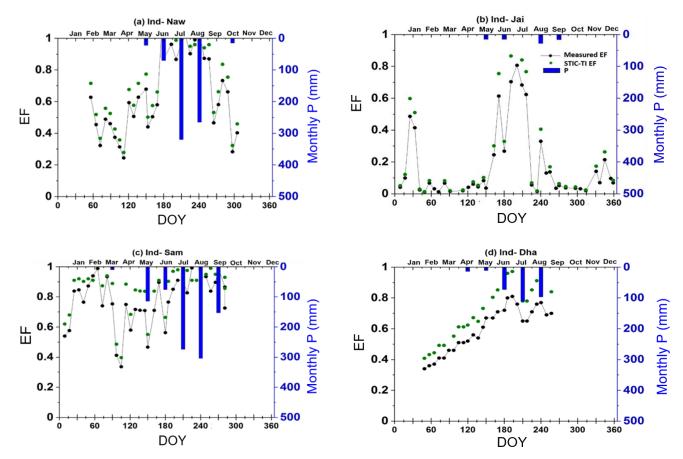


Figure 12: Illustrative examples of temporal variation of STIC-TI derived EF using MYD11A2 LST product with respect to measured EF and P in (a) IND-Naw, (b) IND-Jai, (c) IND-Sam, and (d) IND-Dha

The temporal behavior of EF from STIC-TI using MYD11A2 LST product and EC measurements along with measured θ and P at the OzFlux and AmeriFlux sites also revealed (Fig. 13) close correspondence of STIC-TI with EC observations. Low EF (0.05 – 0.40) during the dry season around DOY 100 – 250 and high EF (>0.4) during the wet season (DOY 1 – 120 and 300 to 360) in AU-ASM, US-Ton and US-Var was observed. The analysis showed that STIC-TI EF can capture the annual variability of observed EF and its responses across different ecosystems during wet and dry seasons. The plots of STIC-TI EF versus measured θ (in the inset of Fig. 13) revealed triangular scatter close to right-angled triangle with positive slope of hypotenuse in three ecosystems AU-ASM, US-Var and US-Ton. This showed that in the water-controlled ecosystems, where distinct wet-dry seasons exist, the positive EF- θ relationship is an outcome of the soil moisture controls on transpiration during the dry season.

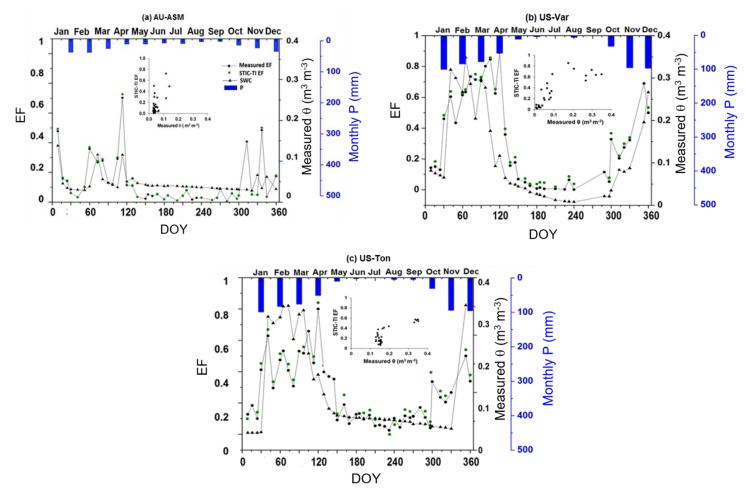


Figure 13: Comparison of temporal variation of STIC-TI derived EF using MYD11A2 LST with respect to measured EF, θ and Pin (a) AU-ASM, (b) US-Var, (c) US-Ton. The scatterplots in the inset shows the relationship between STIC-TI EF with respect to measured θ .

627 **5 Discussion**

628 5.1 Interaction of flux and internal SEB metrices

629 From section 4.1 we found relatively reduced sensitivity of G_i to Ts uncertainties. In any given

630 condition, if an over(under) estimation of M due to noontime T_s uncertainties (through eq. 13)

631 leads to an over(under) estimation of Γ , the effects of such over(under) estimation of Γ (due to

632 noontime T_s uncertainties) tend to be compensated by under(over) estimation of amplitude A (in 633 eq. 5), ultimately leading to a reduction of the sensitivity of G_i to T_s . While the scatter between G 634 versus A for a wide range of Γ (Fig. 14a) revealed large scatter with increasing amplitude under 635 the dry conditions (low Γ), the scatter between Γ versus T_s for different M (Fig. 14b) revealed 636 exponential reduction of Γ with increasing Ts and dryness, and almost no significant change in Γ 637 with increasing T_S at a constantly high dryness (M<0.25). Thus, the confounding effects of Γ , A, 638 and M through eq. 3, 5, 12 and 13 led to a reduction of sensitivity of G to T_S, as exemplified in 639 Fig. 14.

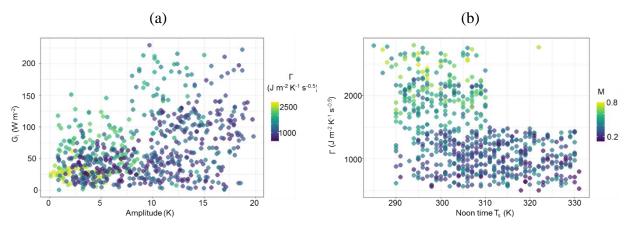


Figure 14: Response plots among parameters of TI-based G_i model, such as (a) G_i versus Amplitude (A) for varying Γ , and (b) Noon-time T_S versus Γ with varying M.

640

641 Concerning LE_i and H_i, dual uncertainties could be propagated in both the fluxes through 642 daytime T_s (through M and G_i), leading to high sensitivity of these two SEB fluxes due to T_s perturbations. The relatively high sensitivity of LE_i to T_S (as compared to H_i) in the non-643 644 irrigated ecosystems could be due to partial compensation of g_A/g_S in both numerator and 645 denominator of the PMEB equation for H (eq. C7 of Appendix C). A recent study (Fig.10 in 646 Mallick et al., 2018a) showed high sensitivity of g_s due to T_s (1% change in T_s led to 5.2–7.5% 647 change in g_S) as compared to g_A sensitivity to T_S (1% change in T_S led to 1.6–2% change in g_A), 648 suggesting that errors in g_S due to T_S uncertainty tend to be larger than errors in g_A . Partial 649 cancellation of the conductance errors in the numerator of eq. (C7 of Appendix C) might have 650 resulted in compensation of H_i errors in the water-limited ecosystems. In this environment, the 651 variability of LE_i is mainly dominated by g_A/g_S , which makes LE_i highly sensitive due to T_S 652 uncertainties. Combined uncertainty due to g_A/g_S in the denominator and g_A in the numerator of eq. (C6 of Appendix C) resulted into greater sensitivity in LE_i to T_S in the arid and tropical savannah ecosystems (Mallick et al., 2015, 2018a; Winter & Eltahir, 2010). The very low sensitivity of LE_i and Hi due to uncertainties in NDVI is because NDVI was not used in the conductance parameterizations and effects due to NDVI in STIC-TI was only propagated through G_i. The sensitivity of LE_i and H_i to albedo was mainly due to the dependence of net radiation (R_{Ni}) on albedo, and any resultant uncertainty in R_{Ni} (due to albedo) tends to be reflected in the sensitivity of LE_i and H_i to albedo.

660 **5.2 Possible sources of errors in SEB flux evaluation**

661 In STIC-TI, underestimation and overestimation errors in G_i in different ecosystems (Fig. 7) could 662 originate due to the errors in MYD11A2 LST product. A host of studies previously reported that 663 the standard deviations of errors in retrieved emissivity in bands 31 and 32 are 0.009, and the 664 maximum error in retrieved Ts of MOD11A1 LST falls within 2-3 K, which is mainly due to the 665 errors in surface emissivity correction (Duan et al., 2017; Wan, 2014; Lei et al., 2018). In the present analysis, we found an overestimation error of MODIS T_s in the range of 0.5 - 1.5 K when 666 667 compared with *in-situ* infrared temperature measurements at the tropical savanna site. As 668 mentioned in section 3.1, a positive (negative) bias in T_S would tend to an overestimation 669 (underestimation) of amplitude (A) in eq. (5); underestimation (overestimation) of M in eq. (13), 670 and consequent underestimation (overestimation) of Γ (eq. 12) and G_i, respectively. Furthermore, 671 the standard deviation of NDVI surrounding the tower sites varied from 0.01 - 0.05 when 672 compared to the ground measurements, which could be another source of error in the STIC-TI 673 model. In addition, NDVI saturates at LAI > 3. However, STIC-TI provides direct estimates of 674 ecosystem G and is independent of R_N.

Despite the comparable accuracy of current G estimates with the G model of Bastiaanssen et al (1998), the foundation of STIC-TI lies in the use of soil moisture characteristics with varying soil textural types which are known to influence the soil heat conductance and thereby G. Thus, the control of soil moisture on evaporation is explicitly included in STIC-TI as opposed to the semiempirical G function of Bastiaanssen et al (1998). The higher accuracies of TI-based thermal diffusion model as compared to R_N -based empirical G models were also reported by Purdy et al. (2016) at daily or longer time scales in cropland, grassland. All these G model estimates many a times differ from in situ measurements because of the no accounting of leaf litter presence or layeron soil floor in the remote sensing-based G-model.

684 The overestimation (underestimation) of LE_i(H_i) is also due to the effects of spatial resolution of 685 different input variables on these two SEB fluxes and conducted statistical evaluation with respect 686 to the measured SEB fluxes. Eswar et al. (2017) demonstrated the need for spatial disaggregation 687 models for monitoring LE_i at field scale using contextual models by disaggregation of evaporative 688 fraction (Λ) and downwelling shortwave radiation ratio (R_G). Using different disaggregation 689 models, they estimated LE_i at 250m spatial resolution and reported RMSE of 30 - 32 W m⁻² as compared to LE_i obtained at 1000m spatial resolution with RMSE of 40 - 70 Wm⁻² over different 690 691 sites in India. Anderson et al. (2007) reviewed different validation experiments conducted in 692 diverse agricultural landscapes (Anderson et al., 2004, 2005; Norman et al., 2003) and reported 693 RMSE in LE_i in the range of 35 - 40 W m⁻² (15%) at 30 - 120 m disaggregated spatial resolution. 694 Current analysis also brought out the need for noon-night thermal imaging with spatial resolution 695 finer than 1000m to adequately capture the magnitude and variability of LE_i in the terrestrial 696 ecosystems especially agroecosystems where average field sizes are less (< 0.5 ha) and fragmented 697 such as in India and other sub-continents.

698 As seen in Fig. 8a and Table 5, there is a gross overestimation of LE_i with respect to the tower 699 observations when MYD11A2 LST was used. The consistent positive BIAS in STIC-TI LE_i in 700 five out of nine sites is presumably due to the overestimation of R_{Ni} (Figure B1 of Appendix B) 701 and underestimation of G_i. Figure 7 shows overestimation of G_i for three OzFlux sites and US sites 702 and underestimation of G_i for Indian site with G_i (STIC-TI) = 0.90 G_i (tower) - 0.10 and 703 overestimation of R_{Ni} at the ecosystem-scale, with R_{Ni} (STIC-TI) = 0.78 R_{Ni} (tower) +58.92 704 (Appendix-B2). This means a systematic overestimation of net available energy $(R_{Ni} - G_i)$ will be 705 obvious in cases where STIC-TI shows underestimation of G_i. Since available energy is an 706 important component for estimating LE through the PMEB equation, an overestimation of net 707 available energy leads to an overestimation of LE by STIC-TI. Sensible heat flux will be 708 consequently underestimated due to the complementary nature of the PMEB equation. It may be 709 also noted that the use of MYD21A2 LST led to relatively better accuracy in LE_i (3-8%) and H_i 710 (2-3%) as compared to using MYD11A2 LST in semi-arid and arid ecosystems. The higher 711 retrieval accuracy of MYD21A2 LST using TES (Temperature-Emissivity Separation) algorithm

- over MYD11A2 LST that uses split-window algorithm (Wan et al, 2015) is the main reason for
 obtaining higher accuracy in LE_i and H_i estimates.
- The standard deviations of MODIS Aqua day-night overpass time over study sites were found to
- be within 30-45 minutes (Sharifnezhadazizi et al, 2019) and the expected deviation in LST from
- the mean local time would be around ± 0.75 K (Sharifnezhadazizi et al, 2019). Sensitivity analysis
- showed that resultant uncertainties in STIC-TI heat flux estimates would be of the order of $\pm 5-7\%$
- 718 due to this LST uncertainty.

719 **5.3 Effects of SEB closure**

Given there is a widespread lack of SEB closure $(H + LE \neq R_N - G)$ or residual energy balance, 720 721 knowledge of the impact of different vegetation types and climatic variables on SEB 'non-closure' 722 is essential. A recent study by Dare-Idowu et al. (2021) covering 8 growing seasons and 3 crops 723 (maize, wheat, and rapeseed) in two sites of south-western France showed that the systematic effect 724 of each site on SEB closure was stronger than the influence of crop type and stage. Same study 725 revealed a greater percentage of SEB closure under unstable atmospheric conditions and in the 726 prevailing wind directions, and sensible heat advection accounted for more than half of the 727 imbalance at both the sites.

728 In our study, using unclosed SEB observations for Indian sites in the absence of in-situ Gi 729 observations also added to the consistent positive BIAS in the statistical evaluation of LE_i. A 730 ubiquitous lack of energy balance closure to the order of 10 - 20% worldwide at most of the EC 731 sites is reported in different literatures (Stoy et al., 2013; Wilson et al., 2002), which implies a 732 systematic underestimation (overestimation) of LE_i (EC tower) (and/or H_i(EC tower)). 733 Accommodating an average 15% imbalance in LE_i (EC tower) would tend to diminish the positive 734 BIAS in STIC-TI. Therefore, the pooled gain (0.98) and positive BIAS between the STIC-TI and 735 tower LE_i is determined by the overestimation of $(R_{Ni} - G_i)$, combined with the underestimation 736 of measured LE_i from the EC towers. An underestimation of H_i (negative BIAS) is associated with 737 two reasons; (a) ignoring the two-sided aerodynamic conductance of the leaves (Jarvis and 738 McNaughton, 1986; Monteith and Unsworth, 2013; Schymanski et al., 2017), which could lead to 739 substantial underestimation of H_i, and (b) due to the complementary nature of the PMEB equation, 740 if LE_i is overestimated, H_i will be underestimated. In addition, frequent micro-advection fluxes 741 alter measured in situ H and LE fluxes. But these advection conditions are not explicitly accounted in the current STIC-TI model. At the EC tower sites, the fraction of residual energy balance to R_N
can be quantified with respect to vegetation/crop growth characteristics or biophysical properties.
However, where G observations are lacking such as in many Indian EC tower sites, the TI-based
G model can be used to fill up the missing G observations to quantify residual energy balance and
to correct the SEB non-closure.

747 **6** Summary and conclusions

748 This study addressed one of the long-term uncertainties in thermal remote sensing of evaporation 749 modeling in open canopy heterogeneous ecosystems, which is associated with empirical methods 750 of estimating ground heat flux. We demonstrated for the first-time physical integration and 751 coupling of a mechanistic ground heat flux model with an analytical evaporation model (Surface 752 Temperature Initiated Closure, STIC) within the surface energy balance equation. The model is 753 called STIC-TI, which uses satellite-based land surface temperature from MODIS Aqua and 754 associated biophysical variables, and it has minimal independence on *in-situ* measurements. The 755 estimation of evaporation through STIC-TI does not require any empirical function for inferring 756 the biophysical conductances. STIC-TI is independent of uncertain parameterizations of surface 757 roughness and atmospheric stability and does not also involve any look-up table for biome or plant 758 functional attributes. By linking noon-night land surface temperature with harmonics equation of 759 thermal inertia and soil moisture availability, STIC-TI derived the ground heat flux, and 760 subsequently coupled it with evaporation. Independent validation of STIC-TI with respect to eddy 761 covariance flux measurements from nine terrestrial ecosystems in arid, semi-arid and sub-humid 762 climate in India, USA and Australia led us to the following conclusions:

(i) While the MODIS Aqua day-night view angle difference showed strong impact on ground
 heat flux and sensible heat flux estimate deviations of STIC-TI (with respect to
 measurements), relatively weak dependence of latent heat flux errors on the day-night view
 angle difference was noted.

(ii) The most notable advantages of STIC-TI are firstly, it provides direct estimates of ground heat
 flux while simultaneously integrates the effects of soil water stress on ground heat flux and
 evaporation through the inclusion noon-night land surface temperature information. Secondly,

the ecosystem-scale surface soil temperature amplitude used in the ground heat flux modelcan advance our understanding on associated terrestrial ecosystem processes.

772 The requirement of few input variables in STIC-TI generates promise for surface-atmosphere 773 exchange studies using readily available data from the current generation remote sensing satellites 774 (e.g., MODIS, VIIRS) that have noon-night land surface temperature observations. STIC-TI can 775 also be potentially used for distributed ET mapping from future high spatial resolution (~ 50-60776 m) TIR missions having noon-night observations with high revisit such as the Indo-French 777 mission, TRISHNA (Thermal infrared Imaging Satellite for High-resolution Natural Resource 778 Assessment) (Lagouarde et al., 2018, 2019), ESA's LSTM (Land Surface Temperature 779 Monitoring), and NASA SBG (Surface Biology and Geology), respectively.

780 Author contributions

KM and BKB conceptualized the idea; DD conducted STIC-TI model coding, simulations; BKB and DD conducted the data analysis in consultation with KM; DD and BKB wrote the first version of the manuscript with KM writing the introduction, discussions, and conclusions; BKB and KM conducted all the analysis and writing during revision; all authors contributed to discussions, editing and corrections; BKB and KM jointly finalized the manuscript.

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800 Data and code availability

801 Harmonized time series datasets the study grids available in over are 802 https://doi.org/10.5281/zenodo.5806501. The model code is available to the first author upon 803 reasonable request.

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1102 Appendix A

Attributes	Symbol	Description		
Temperature	T _A	Air temperature (° C)		
	T _{Max}	Maximum air temperature (° C)		
	T _{Min}	Minimum air temperature (° C)		
	TD	Air dew-point temperature (° C)		
	T _{STA}	point-scale soil temperature amplitude		
	ΔTs	noon-night LST difference (° C)		
-	T _{ST}	Soil temperature (° C)		
-	Ts	Land surface temperature (LST) (°C)		
Humidity,	R _H	Relative humidity (%)		
vapor	eA	Atmospheric vapor pressure at the level of T _A measurement (hPa)		
pressures	e_A^*	Saturation vapor pressure at the level of T _A measurement (hPa)		
	e_s^*	Saturation vapor pressure at surface (hPa)		
	DA	Atmospheric vapor pressure deficit at the level of T_A measurement		
		(hPa)		
Radiation	R _G	Downwelling shortwave radiation (or global radiation) (W m ⁻²)		
	R _R	Upwelling or reflected shortwave radiation (W m ⁻²)		
	$R_L\!\!\downarrow$	Downwelling longwave radiation (W m ⁻²)		
	$R_{L}\uparrow$	Upwelling longwave radiation (W m ⁻²)		
-	τ_{sw}	Atmospheric transmissivity for shortwave radiation (unitless)		
-	α_{R}	Broadband shortwave surface albedo (unitless)		
SEB	LEi	Latent heat flux (W m ⁻²); subscript 'i' signifies 'instantaneous'		
components	Hi	Sensible heat flux (W m ⁻²); subscript 'i' signifies 'instantaneous'		
	Gi	Ground heat flux (W m ⁻²); subscript 'i' signifies 'instantaneous'		
	R _{Ni}	Net radiation (W m ⁻²); subscript 'i' signifies 'instantaneous'		
	ф	Net available energy (W m ⁻²); i.e., $R_N - G$		
	А	Ecosystem-scale surface soil temperature amplitude (°C)		

Table A1: A list of symbols, their descriptions and units used in the present study

MV2007	T _{Sd}	Daytime T _S (° C)		
model	T _{Sn}	Nighttime T _S (° C)		
	ω	Angular frequency (rad s ⁻¹)		
	φ _n '	Phase shift of the n th soil surface temperature harmonic (rad)		
	Δ	Shape parameter (unitless)		
	Sr	Relative soil moisture saturation (m ³ m ⁻³)		
	f_s	Sand fraction (unitless)		
	θ_{fc}	Soil water content at field capacity (m ³ m ⁻³)		
	θ_{wp}	Soil water content at permanent wilting point (m ³ m ⁻³)		
	θ*	Soil porosity (cm ³ cm ⁻³)		
	Js	Summation of harmonic terms of soil surface temperature (K)		
	Ϋ́	Soil textural parameter (unitless)		
	Γ	Soil thermal inertia (J K ⁻¹ m ⁻² s ^{-0.5})		
	τ0	Thermal inertia of air-dry soil (J K ⁻¹ m ⁻² s ^{-0.5})		
	τ*	Thermal inertia of saturated soil (J K ⁻¹ m ⁻² s ^{-0.5})		
	ť'	Time of satellite overpass (seconds)		
	Δt	Time offset between the canopy composite temperature and the		
		below-canopy soil surface temperature (seconds)		
	κ	Total number of harmonics used (unitless)		
	f_c	Vegetation fraction (unitless)		
	θ	Volumetric soil moisture (cm cm ⁻³)		
Clear-sky R _{Ni}	R _{ns}	Net shortwave radiation (W m ⁻²)		
model	R_{nl}	Net long wave radiation (W m ⁻²)		
	G _{sc}	Solar constant (1367 W m ⁻²)		
	β_e	Sun elevation angle (⁰).		
	ε _s	Infrared surface emissivity (unitless)		
	ε _a	Atmospheric emissivity (unitless)		
	Е	Eccentricity correction factor due to variation in Sun-Earth distance		
		(unitless)		
	М	Aggregated moisture availability (0-1)		

STIC-TI	g A	Aerodynamic conductance (m s ⁻¹)		
model	gs	Canopy-surface conductance (m s ⁻¹)		
-	T_0	Aerodynamic temperature (or source/sink height temperature) (°C)		
-	T _{0D}	Dewpoint temperature at the source/sink height (°C)		
-	Λ	Evaporative fraction (unit less)		
-	e ₀	Vapor pressure at the source/sink height (hPa)		
-	e_0^*	Saturation vapor pressure at the source/sink height (hPa)		
-	D ₀	Vapor pressure deficit at source/sink height (hPa)		
-	S ₁	Psychrometric slope of vapor pressure and temperature between (T _{0D}		
		$-T_D$) versus (e ₀ -e _A) (h Pa K ⁻¹)		
-	S 2	Psychrometric slope of vapor pressure and temperature between (Ts-		
		T _D) versus (e_s^* - e_A) (h Pa K ⁻¹)		
	S 3	Psychrometric slope of vapor pressure and temperature between (T_{0D})		
		-T _D) versus (e_s^* - e_A).		
-	κ	Ratio between $(e_0^* - e_A)$ and $(e_s^* - e_A)$ (unitless)		
	S	Slope of saturation vapor pressure vs. temperature curve (h Pa K ⁻¹)		
-	α	Priestley-Taylor coefficient (unitless)		
Ancillary	U	Wind speed at 8 m height (m s ⁻¹)		
meteorological	u*	Friction velocity (m s ⁻¹)		
variables				
	Р	Precipitation (mm d ⁻¹)		
	γ	Psychrometric constant (h Pa k ⁻¹)		
Constants	c _p	Specific heat capacity of air at constant pressure (MJ kg ⁻¹ K ⁻¹)		
	ρ	Density of air (Kg m ⁻³)		
	σ	Stefan–Boltzmann constant (5.67 x 10 ⁻⁸ Wm ⁻² K ⁻⁴)		
Sensor view	VZA	MODIS Aqua sensor view angle (°)		
geometry	δVZA	Difference in MODIS Aqua day-night sensor view angle (°)		

- 1106 **Table A2:** Summary of instruments used, height or depth and period of measurements, measured
- 1107 variables at nine EC flux tower sites

Type of primary instruments	Measurement height/ depth (m)	Measured variables		
used for in situ data recording				
at flux tower sites				
Net radiometer	 3m (IND-Naw, IND-Jai, IND-Sam) 15m (AU-Ade) 12.2m (AU-ASM) 23m (AU-How)2m (US-Ton, US-Var) 	Four radiation flux components: shortwave incoming (R_G) and outgoing (R_R); longwave incoming ($R_L\downarrow$) and outgoing ($R_L\uparrow$)		
EC assembly with IRGA (Infrared Gas Analyzer), three- dimensional sonic anemometer, TC probe	 8m (IND-Naw; IND-Jai; IND-Sam) 4.5m (IND-Dha) 15m (AU-Ade) 11.6m(AU-ASM) 23m (AU-How) 2m (US-Ton, US-Var) 	High response wind vectors (u , v and w), sonic temperature, and CO ₂ - water vapor mass at 10/20 Hz frequency		
Humidity and temperature probe	 8m (IND-Naw, IND-Jai, IND-Sam) 4.5m (IND-Dha) 15m (AU-Ade), 11.6m (AU-ASM) 23m (AU-How), 70m (AU-How) 2m (US-Ton, US-Var) 	$T_{\rm A} and R_{\rm H}$		
Soil temperature probe	 -0.1m (IND-Dha) -0.15m (AU-Ade) (-0.02, -0.06m) (AU- ASM) -0.08m (AU- How) -0.02m, -0.04m, -0.08m, and - 0.16m (US-Ton, US-Var) 	T _{ST}		
Soil heat flux plates	 Ground, 0.1 m (IND-Dha) Ground, -0.15 m (AU-Ade) Ground, -0.08 m (AU-ASM) Ground, -0.15 m (AU-How) -0.01m (US-Ton, US-Var) 	Soil heat flux (G)		

1108 Appendix B

1109 **B1: Clear-sky instantaneous net radiation** (**R**_{Ni}) model

- 1110 Net radiation (R_N) is defined as the difference between the incoming and outgoing radiation, which
- 1111 includes both longwave and shortwave radiation at the Earth's surface.

1112 Terrestrial R_N has four components: downwelling and upwelling shortwave radiation (R_G and R_R), 1113 downwelling and upwelling longwave radiation ($R_L\downarrow$ and $R_L\uparrow$), respectively.

$$\mathbf{R}_{\mathbf{N}} = (\mathbf{R}_{\mathbf{G}} - \mathbf{R}_{\mathbf{R}}) + (\mathbf{R}_{\mathbf{L}\downarrow} - \mathbf{R}_{\mathbf{L}\uparrow}) \tag{B1}$$

1114 Out of these four terms mentioned in eq.(B1), R_G and $R_L\downarrow$ are dependent on various factors such 1115 as geographic location, season, cloudiness, aerosol loading, atmospheric water vapor content and 1116 less on surface properties. On the other hand, the upwelling radiations in eq. (B1) strongly depends 1117 on the surface properties such as surface reflectance and emittance, land surface temperature, and 1118 soil water content (Zerefos and Bais, 2013).

1119 Instantaneous net radiation (R_{Ni}) can be derived using eq. B2 as follows (Mallick et al., 2007):

$$R_{\rm Ni} = R_{\rm ns} - R_{\rm nl} \tag{B2}$$

$$R_{ns} = (1 - \alpha_R) R_G \tag{B3}$$

$$R_{nl} = R_{L\downarrow} - R_{L\uparrow} \tag{B4}$$

1120 Where, R_{ns} is net shortwave radiation (W m⁻²), R_{nl} is net longwave radiation (W m⁻²).and α_{R} is the 1121 broadband surface albedo shortwave spectrum.

A WMO (World Meteorological Organization) shortwave radiation model (Cano et al.,1986)
calibrated over Indian conditions (Mallick et al., 2007, 2009) was used to compute R_G using the
following equation:

$$R_{G} = \tau_{sw}G_{sc}E (sin\beta_{e})^{1.15}$$
(B5)

1125 Where, τ_{sw} is the is the global clear sky transmissivity for the shortwave radiation (0.7), G_{sc} is the 1126 solar constant (1367 Wm⁻²), ϵ is the eccentricity correction factor due to variation in Sun-Earth 1127 distance and β_e is the sun elevation in degrees.

1128 $R_{L}\downarrow$ at any instance was calculated as follows:

$$R_{L\downarrow} = \varepsilon_a \sigma \left(273.14 + T_A\right)^4 \tag{B6}$$

- 1129 Where, σ is the Stefan–Boltzmann constant (5.67 x10⁻⁸ Wm⁻²K⁻⁴); T_A is the air temperature (⁰C);
- 1130 ε_a is the atmospheric emissivity.

1131 Atmospheric emissivity (ε_a) was computed using the following equation (Bastiaanssen et al.,1998):

$$\varepsilon_a = 0.85 - \ln \tau_{sw}^{0.09}$$
 (B7)

1133 $R_{L\uparrow}$ at any particular instance was calculated as follows:

$$R_{L\uparrow} = \varepsilon_s \sigma (273.14 + T_s)^4 \tag{B8}$$

1134 Where, ε_s is the surface emissivity in thermal infrared (8 – 14 µm) spectrum and T_s is the land 1135 surface temperature (⁰C).

1136 **B2: Evaluation of STIC-TI R**_{Ni}

1137 Comparison of the clear-sky R_{Ni} estimates with respect to *in situ* measurements revealed RMSE in 1138 R_{Ni} to the order of 27 – 72 W m⁻², MAPD 8 –24%, BIAS (-67) – 50 W m⁻², and R² varying from 1139 0.62– 0.90 across all the sites (Fig. B2, Table B2). Among the nine sites, a consistent 1140 underestimation of R_{Ni} was noted in IND-Dha, US-Ton, and US-Var (with BIAS of -23 W m⁻², -1141 61 W m⁻² and -67 W m⁻²), whereas substantial overestimation of R_{Ni} was found in IND-Sam, IND-1142 Naw, and AU-ASM with a BIAS of 50 W m⁻², 37 W m⁻² and 43 W m⁻², respectively (Table B2).

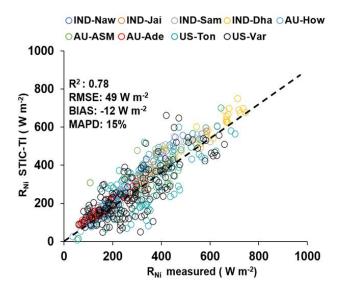


Figure B2: Validation of STIC-TI derived R_{Ni} estimates with respect to *in situ* measurements in different ecosystems. The regression equation between modeled versus in-situ R_{Ni} is, R_{Ni} (STIC-TI) = 0.78 R_{Ni} (tower) +58.92.

1143 **Table B2:** Performance evaluation statistics of clear-sky R_{Ni} estimates in nine different 1144 agroecosystems

Sites	Error statistics of clear-sky R _{Ni} model				
	estimates				
	R ²	BIAS RMSE		MAPD	
		(W m ⁻²)	(W m ⁻²)	(%)	
IND-Jai	0.81	-9	32	8	
IND-Naw	0.81	37	56	12	
IND-Dha	0.81	-23	42	9	
IND-Sam	0.64	50	67	15	
US-Ton	0.68	-61	69	21	
US-Var	0.62	-67	72	24	
Au-How	0.87	7	27	15	
AU-ASM	0.88	43	50	14	
AU-Ade	0.90	11	27 16		

1145 Appendix C

1146 C1: Estimating SEB fluxes using STIC1.2 analytical model and thermal remote sensing data

1147 STIC1.2 (Mallick et al., 2014, 2015a,b, 2016, 2018a) is a one-dimensional physically based SEB 1148 model and is based on the integration of satellite LST observations into the Penman-Monteith 1149 Energy Balance (PMEB) equation (Monteith, 1965). In STIC1.2, the vegetation-substrate 1150 complex is considered as a single unit. Therefore, the aerodynamic conductances from individual 1151 air-canopy and canopy-substrate components is regarded as an 'effective' aerodynamic 1152 conductance (g_A) , and surface conductances from individual canopy (stomatal) and substrate 1153 complexes is regarded as an 'effective' canopy-surface conductance (gs) which simultaneously 1154 regulate the exchanges of sensible and latent heat fluxes (H and LE) between surface and 1155 atmosphere. One of the fundamental assumptions in STIC1.2 is the first order dependence of these two critical conductances on M through T_S. Such an assumption enabled an integration of satellite 1156 1157 LST in the PMEB model (Mallick et al., 2016). The common expression for LE and H according 1158 to the PMEB equation is as follows:

$$LE = \frac{s\phi + \rho c_P g_A D_A}{s + \gamma \left(1 + \frac{g_A}{g_S}\right)}$$
(C6)

$$H = \frac{\gamma \varphi \left(1 + \frac{g_A}{g_S}\right) - \rho c_P g_A D_A}{s + \gamma \left(1 + \frac{g_A}{g_S}\right)}$$
(C7)

In the above equations, the two biophysical conductances (g_A and g_S) are unknown and the STIC1.2 methodology is based on finding analytical solutions for the two unknown conductances to directly estimate LE (Mallick et al., 2016, 2018a). The need for such analytical estimation of these conductances is motivated by the fact that g_A and g_S can neither be measured at the canopy nor at larger spatial scales, and there is no universally agreed appropriate model of g_A and g_S that currently exists (Matheny et al., 2014; van Dijk et al., 2015). By integrating T_S with standard SEB theory and vegetation biophysical principles, STIC1.2 formulates multiple state equations in order 1166 to eliminate the need to use the empirical parameterizations of the g_A and g_S and also to bypass the 1167 scaling uncertainties of the leaf-scale conductance functions to represent the canopy-scale 1168 attributes. The state equations for the conductances are expressed as a function of those variables 1169 that are mostly available as remote sensing observations and weather forecasting models. In the 1170 state equations, a direct connection to T_s is established by estimating M as a function of T_s. The 1171 information of M is subsequently used in the state equations of conductances, aerodynamic 1172 variables (aerodynamic temperature, aerodynamic vapor pressure), and evaporative fraction, 1173 which is eventually propagated into their analytical solutions. M is a unitless quantity, which 1174 describes the relative wetness (or dryness) of a surface and also controls the transition from potential to actual evaporation; which implies $M \rightarrow 1$ under saturated surface conditions and $M \rightarrow 0$ 1175 1176 under extremely dry conditions. Therefore, M is critical for providing a constraint against which 1177 the conductances are estimated. Since T_s is extremely sensitive to the surface moisture variations, 1178 it is extensively used for estimating M in a physical retrieval scheme (detail in Appendix A3 of 1179 Bhattarai et al., 2018; Mallick et al., 2016, 2018a). It is hypothesized that linking M with the 1180 conductances will simultaneously integrate the information of T_S into the PMEB model. To 1181 illustrate, we express the state equations by symbols, $sv_1 = f \{c_1, c_2, c_3, v_1, v_2, v_3, v_4, sv_3, sv_5\}$; sv_2 1182 $= f \{v_4, sv_1, sv_5, sv_6\}; sv_3 = f \{c_3, v_3, v_4, sv_4, sv_5\}; sv_4 = f \{c_3, v_3, sv_1, sv_2, sv_7, sv_8\}.$ Here, f, sv, v, 1183 and c denote the function, state variables, input variables (5 input variables; radiative and 1184 meteorological), and constants (3 constants), respectively. Here sv_1 to sv_4 are g_A , g_S , aerodynamic 1185 temperature (T₀), evaporative fraction (Λ), and sv₈ is M. Given the estimates of M, net radiative 1186 energy ($R_{\rm Ni}$ – $G_{\rm i}$), $T_{\rm A}$, $R_{\rm H}$, the four state equations are solved simultaneously to derive analytical 1187 solutions for the four state variables and to produce a surface energy balance "closure" that is 1188 independent of empirical parameterizations for g_A , g_S , T_0 , and Λ . However, the analytical solutions 1189 to the four state equations contain three accompanying unknown state variables (effective vapor 1190 pressures at source/sink height, and Priestley-Taylor variable), and as a result there are four 1191 equations with seven unknowns. Consequently, an iterative solution was found to determine the 1192 three additional unknown variables as detailed in this section above and also described in Mallick 1193 et al. (2016, 2018a) and Bhattarai et al. (2018). The state equations of STIC are given below.

$$g_{A} = \frac{\Phi}{\rho c_{P} \left[(T_{0} - T_{A}) + \left(\frac{e_{0} - e_{A}}{\gamma} \right) \right]}$$
(C1)

$$g_{S} = g_{A} \frac{(e_{0} - e_{A})}{(e_{0}^{*} - e_{0})}$$
 (C2)

$$T_0 = T_A + \left(\frac{e_0 - e_A}{\gamma}\right) \left(\frac{1 - \Lambda}{\Lambda}\right)$$
(C3)

$$\Lambda = \frac{2\alpha s}{2s + 2\gamma + \gamma \frac{g_A}{g_S}(1+M)}$$
(C4)

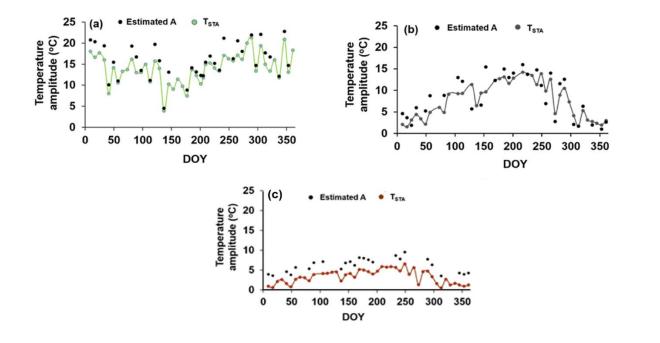
1194 Detailed derivations of these four state equations are given in Mallick et al. (2016). Given the values of M, R_N, G, T_A, and R_H or e_A, the four state equations can be solved simultaneously to 1195 1196 derive analytical solutions for the four unobserved variables and to simultaneously produce a 1197 'closure' of the PMEB model that is independent of empirical parameterizations for both g_A and gs. However, the analytical solutions to the four state equations contain three accompanying 1198 unknowns; e0 (vapor pressure at the source/sink height), e0* (saturation vapor pressure at the 1199 1200 source/sink height), and Priestley-Taylor coefficient (α), and as a result there are four equations 1201 with seven unknowns. Consequently, an iterative solution was needed to determine the three 1202 unknown variables (as described in Appendix A2 in Mallick et al. 2016). Once the analytical solutions of gA and gS are obtained, both variables are returned into eq. (13) to directly estimate 1203 1204 LE.

In STIC-TI, an initial value of α was assigned as 1.26; initial estimates of e_0^* were obtained from 1205 T_{s} through temperature-saturation vapour pressure relationship, and initial estimates of e_{0} were 1206 obtained from M as, $e_0 = e_A + M(e_0^* - e_A)$. Initial T_{0D} and M were estimated according to 1207 1208 Venturini et al. (2008) as described in section 3.2, and initial estimation of G was performed from 1209 initial M using the equation sets eq. (2) - eq. (11). With the initial estimates of these variables; first estimate of the conductances, T_0 , Λ , H, and LE were obtained. The process was then iterated 1210 by updating e_0^* , D_0 , e_0 , T_{0D} , M, and α (using eq. A9, A10, A11, A17, A16 and A15 in Mallick et 1211 al., 2016), with the first estimates of g_S, g_A, T₀, and LE, and re-computing G, ϕ , g_S, g_A, T₀, A, H, 1212 and LE in the subsequent iterations with the previous estimates of e_0^* , e_0 , T_{0D} , M, and α until the 1213

- 1214 convergence of LE was achieved. Stable values of G, conductances, LE, H, T₀, e₀^{*}, e₀, T_{0D}, M, and
- 1215 α were obtained within ~25 iterations. The inputs needed for computation of LE_i (eq.C6) are air
- 1216 temperature (T_A) , land surface temperature (T_S) , relative humidity (R_H) , net radiation (R_{Ni}) and
- 1217 soil heat flux (G_i).

1218 Appendix D

- 1219 The temporal variation of estimated A and T_{STA} is shown in Fig. D1. The annual variations of T_{STA}
- 1220 in different ecosystem was found to be within the ranges of 1 4° C.



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1222 Figure D1: Temporal variation of A and T_{STA} in (a) AU-ASM (2013), (b) US-Ton (2014), (c) US-

- 1223 Var (2014).
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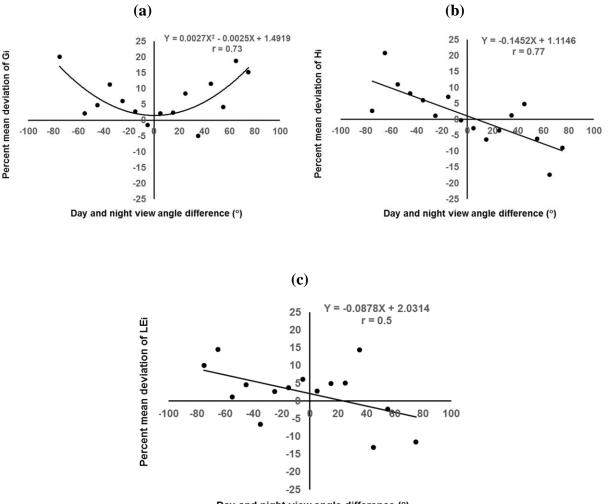
1230 Appendix E

Table E1: Soil textural properties and their values used in the present study (Murray and Verhoef,
2007; Minasny et al., 2011; Anderson et al., 2007).

Soil texture	Water retention Shape parameter (δ)	Field capacity (vol/vol) (%) θ_{fc}	Wilting point (vol/vol) (%) θ _{wp}	Sand fraction (f _s)	Saturated soil moisture (vol/vol) (%) 0*
Sand	2.77	10	5	0.92	43
Loamy Sand	2.39	12	5	0.82	41
Sandy loam	2.27	18	8	0.58	41
Loam	2.20	28	14	0.43	43
Silty loam	2.22	31	11	0.17	45
Sandy clay loam	2.17	27	17	0.58	39
Clay loam	2.14	36	22	0.40	41
Silty clay loam	2.14	38	22	0.10	43
Sandy clay	2.11	36	25	0.52	38
Silty clay	2.12	41	27	0.06	46
Clay	2.10	42	30	0.22	38

1233 Appendix F

1234 Day-night view angle effect on errors of STIC-TI heat flux estimates from measurements is shown1235 in Figure F.



Day and night view angle difference (°)

Figure F: Dependence of STIC-TI model flux error in terms of mean percent deviation from measurements on day-night view angle difference of MODIS Aqua expressed as mean of 10° bin for (a) G_i, (b) H_i, and (c) LE_i.

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